Name:

Enrolment No:



UPES

End Semester Examination, December 2023

Course: STELLAR EVOLUTION

Semester: V

Program: B.Sc (H) Physics Time : 03 hrs. Course Code: PHYS 3024 Max. Marks: 100

Instructions: 1. All questions in Section A are mandatory, while question # 9 and 11 of Sections B and C, respectively, have internal options.

2. Give deserving descriptions too to questions in Section A

SECTION A	
(5Qx4M=20Marks)	

S. No.		Marks	CO
Q 1	What are the main 2 forces acting on a star on main sequence?	4	CO1
Q 2	What are the preceding stages of its development before a star is born?	4	CO1
Q 3	Why, do you think, is convection important in a star?	4	CO1
Q 4	What are AGNs? Give the full form of the acronym and describe them.	4	CO2
Q 5	Supernova is well heard of! What is Nova?	4	CO1
	SECTION B		
	(4Qx10M= 40 Marks)		
Q 6	One of the ends of a star is the white dwarf stage. And even amongst the white dwarfs there are differences. Explain what and how.	10	CO1
Q 7	Explain what solar wind is. Evaluate the effects it has had and still makes on the solar system since its formation.	6+4	CO2
Q 8	There is interstellar dust and then there is interstellar gas. Are they different? Describe both and their relation, if any.	2+8	CO3
Q 9	Describe the nuclear reactions that take place in the core of a Sun like stars.	10	
	OR		CO2
	The energy production in the Sun happens only at its core. Hence, the temperature falls off radially outwards. Till this region is arrived at! Which one? And why this rise?	3+7	

SECTION-C (2Qx20M=40 Marks)				
Q 10	Stars are formed from giant clouds in space. Appraise how such clouds are maintained before the collapse sets in. What is the single most property which determines the evolution of a star? In to how many classes are the stars binned according to this property?	16+2+2	CO3	
Q 11	Analyze the evolution of a high mass star. Divide this evolution into discrete phases and make this analysis. Do highlight the underlying Physics in the different phases.	20	CO1	
	 a. Apprise what degeneracy means. Discuss the phenomenon of electron degeneracy in stellar evolution. b. Appraise the state of a neutron star with the rich Physics involved. 	10 10		