## DESIGN \& COMPUTATIONAL ANALYSIS FOR SATELLITE ATTITUDE CONTROL SYSTEMS AND PERTURBATION CORRECTION ANALYSIS FOR NANOSATELLITES IN LOW EARTH ORBIT WITH UTILIZATION OF KALMAN FILTERS

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## DECLARATION BY SCHOLAR

I hereby declare that this submission is my own work and that, to the best of my knowledge and belief, it contains no material previously published or written by another person nor material which has been accepted for the award of any other degree or diploma of the university or other institute of higher learning, except where due acknowledgment has been made in the text.
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## THESIS COMPLETION CERTIFICATE

This is to certify that the thesis on "DESIGN \& COMPUTATIONAL ANALYSIS FOR SATELLITE ATTITUDE CONTROL SYSTEMS AND PERTURBATION CORRECTION ANALYSIS FOR NANOSATELLITES IN LOW EARTH ORBIT WITH UTILIZATION OF KALMAN FILTERS" by
RAJA M in Partial completion of the requirements for the award of the Degree of Doctor of Philosophy Engineering is an original work carried out by him under our joint supervision and guidance. It is certified that the work has not been submitted anywhere else for the award of any other diploma or degree of this or any other University.


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#### Abstract

An Attitude control system plays the important role to maintain the satellite to desired attitude orientations. The intended application of NANO satellite in low earth orbits (LEO) helps in determining its per-determined orbits from disturbed or perturbed orbit. The LEO orbits typically at an altitude in the range of 160-2000 km. The LEO satellites are widely used for remote sensing, navigation, and military surveillance applications. There is more development in the field of Nano satellite system design and control of a variety of applications. The Numerical ODE integration process helps to find initial position and velocity vectors of the Nano satellite by Runge - Kutta using Cowell's Method. The perturbation simulation is analyzed with MATLAB, General Mission Analysis Tool (GMAT) open source software developed by the National Aeronautics and Space Administration (NASA). The Keplerian results are validated with General mission analysis tool. The design parameters of Nano satellites such as Moment of inertia, North American Aerospace Defense Command (NORAD) two-line element commands, and Geometry parameters are considered. The Nano satellite Attitude control systems (ACS) for SRM Satellite, Pratham (IIT Bombay), NPSAT-1 described in the research work. The high pointing accuracy attitude estimation and feedback control systems are presented. The dynamics attitude controller with feedback is implemented by MATLAB/SIMULINK package. The small satellite is very much essential to collect the information in the space environments. The most of the Keplerian or orbital elements are considered as ideal condition. A satellite is expected to move in the orbit until its life is over. This would have been true if the earth was a true sphere and gravity was the only force acting on the satellite. Under the initial condition the motion of two bodies like Earth -Satellites are considered. Practically, this is not possible. The motion of the body includes the disturbing forces in the orbit. However, a satellite has deviated from its normal path due to several forces. This deviation is termed as orbital perturbation. The changes in the orbital element with respect to secular variations are considered. This work describes the conservative forces like aerodynamic drag and solar pressure. In this thesis, an


overview of orbital perturbation of the six Keplerian elements like semi-major axis, True anomaly, Longitude of ascending node, eccentricity, inclination, Argument of perigee are presented. The numerical simulation to demonstrate the performance of SRM Satellite, Pratham, and International Space Station (ISS) is Performed. The perturbation algorithm is implemented in MATLAB Environment. To maintain the orientation of the satellite it is necessary to design the attitude estimation technique. The Satellite is considered as rigid body representing the attitude parameter. An inertia matrix describes the rigid body dynamics. The attitude orientation of the satellite using Quaternion and Euler angle is derived. Calculation of the Euler angles (Roll angle, Pitch angle, Yaw angle) with Direction cosine matrixes (DCM) have singularity and computational problems compared to the Quaternion method are discussed. In Low earth orbits, satellite will have an enormous amount of aerodynamic drag acting on the satellite body rapidly in low earth orbit due to centripetal force and gravitational attraction because of that satellite dwell time is reduced. An attitude estimation is measured by the orientation of the vectors. An Attitude sensor is used to measure the satellite orientation in the reference frame which will help in accurately predicting the orbit deviation.

In the presents work, LEO orbit satellite attitude control is implemented by Armature control DC Motor acting an actuator. The magnetic torque is having a solenoid coil generating the magnetic flux which interacts with GEO magnetic fields with the help of magnetometers. The GEO magnetic fields modeling is considered in the International Geomagnetic Reference Field IGRF (IGRF-12). The magnetic moments or control torques are generated by the magnetic coil in satellite body or torqrods to determine the attitude angles and angular rates of the body. The attitude (Roll, Pitch, and Yaw) estimations of Nano satellite NPSAT-1 using Kalman filter and fuze the data to on-board attitude sensors like INS/GPS, Magnetometer reference with low earth orbit satellite. A low-cost sensor is used for simulation of Nano satellite with a Kalman filter. This filter predicts the future estimates state from the magnetometer and attitude quaternions. The design specifications are taken to meet the accuracy requirements (desired value $\leq 0.2$ seconds) of Nano satellite
attitude control. This research work presents the Kalman algorithm with magnetometer and Inertial sensor information. The stabilization of a Nano satellite using magnetic torquer concepts are considered with principle moment of inertia of the model. The feedback signal from on-board sensors compare with reference orbit trajectory and implementation of the Proportional Derivative (PD) controller is constructed. The spacecraft control system used to improve the transient response like overshoot and settling time of the system. Thus, in the design of attitude control rise time, setting time, (desired value $\leq 0.2$ seconds), minimum overshoot, and no steady state error were achieved

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## LIST OF ABBREVIATIONS AND ACRONYMS

$K E=$ Keplerian elements

EOM: Equation of motion
COE: Classical orbital elements
NORAD: North American Aerospace Defense
HEO: High earth orbit
GEO: Geo-synchronous Earth orbit
CMG: Control momentum gyroscope
SMA: Semi-Major Axis
INS: Inertial Navigation System
GPS: Global Position System
IMU: Inertial Measurement Unit
ISRO: Indian space research organization
NASA: National Aeronautics and Space Administration
CF: Coordinate frame
IF: Inertial frame
ECEF: Earth centered Earth Fixed frame
DCM: Direction cosine matrix
EP: Ecliptic Pole
AN: Ascending node
DN: Descending node
GP: Gyroscopic precision
MT: Magnetic torque
FW: Flywheels
MW: Momentum wheels
AD: Atmospheric drag
as: arcsecond
ISS: International Space Station
MDM: Magnetic dipole moment
VCMGs $=$ Variable-speed CMGs

SGVSCMG: Single gimbal Variable speed control momentum gyroscope
DGV: Double-gimbal VSCMG
LOS: Line of sight
OP: Orbital plane
ADCS: Attitude Determination and Control System
RAAN: Right assertion to ascending node
ODE: Ordinary Differential Equation
PID: Proportional-Integral-Derivative controller
RG's: Rate Gyro's
IGRF: International Geomagnetic Reference Field
IVAB: Inner Van Allen Belt
OVAB: Outer Van Allen Belt
PMS: Permanent Magnet Stabilization
GGS: Gravity Gradient Attitude Stabilization
UT: Unscented transformation
KF: Kalman filter
GNSS: Global Navigation Satellite system

CM: Covariance matrix
UKF: Unscented Kalman Filter
GMAT: General Mission Analysis Tool
TA: True anomaly

SF: Solar Flux (1353 W/m²)
EO: Earth's Oblateness
EP: Equatorial plane
KF-RE: Kalman Filter Rate Estimator
AP: Argument of perigee
SSE: State Space Equation
EKF: Extended Kalman Filter
GM: Greenwich meridian
VE: Vernal Equinox

## LIST OF SYMBOLS

G: Gravitational constant, $\left(G=6.674 * 10^{-11} \mathrm{Nm}^{2} \mathrm{~kg}^{-2}\right.$ )
$r_{\text {earth }}:$ Radius of the earth, $\left(6.37 * 10^{6} \mathrm{~m}\right)$
$m_{1:}$ Earth mass, $\mathrm{Kg}\left(5.972 * 10^{\wedge} 24 \mathrm{Kg}\right)$
$m_{2}$ : Satellite mass, Kg
$F_{i}$ : Forces acting on the $i^{\text {th }}$ body
K: Gaussian parameter
$\omega$ : Angular velocity
$\tau$ : Torque
$\mu$ : Gravitational constant Let, $\mu=G * m_{l}$
$a_{x}, a_{y}, a_{z}:$ Acceleration (x, y, z-directions)
( $e_{r}$ and $e_{\theta}$ ): Unit vectors
$h$ : Angular momentum per unit mass
a: Semi major axis
b: Semi minor axis
e: Eccentricity, ecc
$\Omega_{\text {RAAN: }}$ Longitude of ascending node
$\omega$ : Argument of perigee
i: Orbital inclination
$\xi:$ Real Value
$v$ : True anomaly, $T_{\text {anomaly }}$
$v_{c}$ : Circular velocity
T: Orbital period
( $\theta, \psi, \varphi)$ : Satellite Attitude Rates
q: Quaternion
$R_{\text {max }}$ : Rotation matrix, $R_{B}^{o}$ (Satellite body to Orbit)
(I, J, K): Inertial reference frame
( $x_{o}, y_{o}, z_{o}$ ): Orbital frame
$T_{\text {Satellite: }}$ Satellite torque
$R_{\text {apogee }}$ : Distance between satellite and Earth at apogee point
$R_{\text {perigee: }}$ Distance between satellite and Earth at perigee point
$r$ : position vector
$v$ : Velocity vector
c: Velocity of light $\left(3 \times 10^{8} \mathrm{~m} / \mathrm{s}\right)$
ת: Orbital angular velocity
$J_{2}$ : Zonal perturbations
$\frac{m_{2}}{C_{d} A}:$ Ballistic constant
$C_{d}$ : Coefficient of drag
A: Area of the satellite
$A_{\Theta}$ : Area of the SUN
O: orbit
$F_{\text {drag }}:$ Aerodynamic drag force, $F_{A D}$
$\rho$ : Air density
H: Atmospheric scale height of density
$R=x_{i}+y_{j}+z_{k}:$ Position vectors
$V=u i+v j+z k:$ Velocity vectors
$\Omega_{\text {SUN }} \& \Omega_{\text {Moon: }}$ Longitude of ascending nodes SUN \& LUNAR (degrees/day)
$\omega_{\text {SUN }} \& \omega_{\text {moon }}$ : Argument of perigee SUN \& LUNAR (degrees/day)
$a_{r}$ : Solar radiation force
$a_{p}$ : Perturbing acceleration
N: North pole

B: GEO Magnetic field (Earth's Magnets)
$\phi=$ Flight path angle
M: Dipole moments
I: Current in the coil
N: No of turns in the coil
$T_{\text {Required: }}$ Required torque
$N$ : Nadir
Z: Zenith
$\delta$ : Damping ratio
$\omega_{n}$ : Undamped natural frequency
$H_{s}$ : Satellite angular momentum
$H_{w}$ : Wheel angular momentum
$l, m, n$ : Satellite with the principle axis of body
( $I_{x x}, I_{y y}, I_{z z}$ ): Satellite principle Moments of inertia
$J_{W}:$ Moment of inertia of Wheel
$J_{G}:$ Moment of inertia of Gimbal
$c_{r}:$ Reflectivity of SUN
$T_{\text {gravity: }}$ Gravity gradient torque
$K_{\text {Gain: }}$ Kalman Gain
$x$ : system state variable
$P_{C U}$ : Covariance updates
$w_{k}$ : Process Noise
$v_{k}$ : Measurements Noise

A(nxn): Plant (or) Process Matrix

B: Control matrix

F: Controller Gain
$u_{d}:$ Perturbation forces

H: Measurement matrix
$Z_{k}$ : Actual Measurements
$Q_{k}$ : Process covariance
$R_{k}$ : Measurements covariance

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## CHAPTER 1

## INTRODUCTION

### 1.1 Overview

The attitude control System (ACS) of satellite is important to maintain the position and orientation of the vehicle into the original orbital plane. There are two methods used for attitude control system, one is active control system and another passive control system. The active control technique requires a signal from feedback sensor signal and applies the signal to vehicle actuator to maintain the orientation of the satellite body. The passive attitude control method uses the gravity gradient technique to stabilize the satellite with respect to its orientation [1]. These methods utilize numerical computation techniques to apply the feedback signal to actuation systems, magnetic torquers, fly wheel, momentum wheels. This work analyzes the various perturbations or disturbances in low earth orbits (LEO) and attitude control of Nano-Satellite with error estimation using Kalman Filters [2]. An especially in a LEO, due to the gravitational attraction, satellites have an enormous amount of aerodynamic drag compared to the high earth orbits (HEO) with little or no aerodynamic drag. The attitude determination ( AD ) is the integral part of the satellite used to calculate the attitude information for various times with the help of attitude sensors like INS/GPS (Inertial Navigation System/Global Position System), IMU (Inertial Measurement Unit), and Magnetometer, Sun Sensor, Star Sensor, and the Earth sensor [2]. The Satellite attitude governor/control is very important to stabilize the satellite along with its predetermined orientation. Due to the perturbation forces (or) environmental disturbances the satellite may get affected and the original orbit may be changed due to these perturbations. Thus, it is very important to reduce these perturbation forces, as they can cause life term of the satellite to be significantly reduced [3]. It is essential to maintain the satellite pointing accuracy for the proper orientation of the communication, such as GPS antenna pointing towards the ground station on Earth's surface to transmit and receive the data for telemetry, or the solar panel pointing towards the SUN. This has great importance, since thrusters require energy from sun
to re-orient the satellite to the desired path [3]. Mostly, communication system using directional antennas sends/receives signal transmitter to receiver by directional transmission methods or line of sight (LOS) communications. The signals from the satellites are directed towards to the ground station on Earth or inter satellite communications. The solar panels point towards the SUN at an orientation because it consumes more optimal electricity or power. This is especially important, since thrusters require energy from the sun to re-orient the satellite to the desired path [4]. The Remote sensing satellite to capture data from spectrometers/HD cameras focuses terrain surface all the time, and it requires that the camera should be pointing towards Earth's surface for recording accurate data collections. The Telescopic satellite to focus on one location requires the proper orientation schemes. It is also required to use the attitude stabilizing techniques to maintain the required path [5]. The environmental forces also differ at various altitudes. For example, LEO satellites will be affected more by aerodynamic drag and gravitational attraction due to the proximity of earth as compared to the other perturbations. Three body problems can occur at HEO, with Earth-Satellite-Moon attraction problem. The Spacecraft stabilization can be classified into two methods, one as the spin stabilization method and second is gravity gradient method. The spin stabilization methods require energy to control the satellite. The actuators are used to reorient the satellite in the desired path from perturbed path and they act as magnetic torquers at LEO Nanosatellites. An attitude sensor is used to measure the attitude (pitch, roll, and yaw) of satellite and these feedback signals are compared with pre-determined orientations [6]. These signals trigger the dynamics of the vehicles to move the control surface according to the reference trajectory. The magnetic thruster is basically a dipole solenoid coil which interacts with the earth magnetic flux and creates the control torque, which requires maintaining its specific path for low earth orbiting satellites. At the High earth orbit satellite such as GEO Synchronous orbit satellite control momentum gyroscope (CMG) acts as an actuator to generate the necessary control torques [6]. There are various factors affect the lifetime of satellites, such as types of the orbits, perturbation forces, type of satellite parameters, as well as the exposures to the different types of environment. The accurate attitude determination techniques using with attitude sensors (Magnetometer, INS/GPS, and IMU) and control techniques
using actuator (Magnetic torquer, Thruster, CMG) are used to increase the life time/mission life and performances of the satellite [7], [8]. The Attitude determination and control system (ADCS) is very important to keep the satellite from disturbed orbit to original orbit. It requires signal from various on-board attitude sensors to compare this signal with the reference trajectory based on how it controls the actuator with suitable (or) required orientation. India has launched 120 satellites at single launcher by "Indian space research organization" (ISRO) which includes the weather monitoring, remote sensing, disaster management, communication \& navigation, military operation, and earth observation satellites. The Chandrayan-1 satellite launched into lunar orbit-lunar observation, Mangalyaan-Mars orbiter mission, Chandrayan-II tentatively scheduled launch in 2019 between the months of January to March. The ISS is a major satellite space station that orbits in LEO. However, due to the perturbations, the satellite altitude varies with time. To maintain the accurate planned orbit, it's highly recommended that the attitude determination and control algorithm are used with frequent thruster to re-orient the satellite into mission orbit [9].

### 1.2. Research Motivation

The improvement of aerospace engineering and space research is becoming highly integrated, complex, knowledge intensive and globally distributed. The aerospace industry needs more accurate results and successful mission to meet the competitive global markets. Therefore, it becomes essential to study, evaluate and analyze the existing systems and identify areas for the improvement of aerospace/space systems. The mathematical modeling and numerical methods are easy to be modelling systems dynamics and behaviors [10]. The simulation tools are widely used to analyze various models of transient responses of actuators and plant dynamics. Its highly recommend, the more pointing accuracy of attitude controls of satellite for getting the optimum results/Mission performances. To design the low cast attitude error estimation using Kalman filters (KF) [11]. The space industry is spending lots of money on the design and development of space mission operations such as launching the satellite, trajectory transfer, formation flying, planetary observations and earth observations to achieve accurate performances and mission effectiveness with qualitative/optimum
results, it is requiring satellite high pointing accuracy attitude determination and control systems. The proper ADCS used to increase the lifetime of the vehicle and mission performances [12]. The research focus on to design and analysis of the satellite attitude control with orbital perturbation at low earth orbiting satellite and error estimation using Gaussian filter like Kalman Filters used for updates and measurements of present states and future estimates from satellite attitude sensors. The author Giebelmann J worked on "Development of an active magnetic attitude determination and control system for picosatellites on highly inclined circular low earth orbits" [13]. This work helped to recognize the basic ideas about dynamics of Small/Nano satellite in low earth orbit and attitude control techniques. The research discussed about the magnetic stabilization methods for attitude correction. This impressed me to start the magnetic torquers concepts, how the GEO magnetic field varies with atmosphere inner van Allen radiation belts and outer Van Allen belts used to generate the required torque to adjust the satellite form perturbed orbit to original satellite orbit. Another research work helped me to understand the Kalman Filter, Murty S. Challa has presented [14] "A Simple Attitude Unscented Kalman Filter: Theory and Evaluation in a Magnetometer-Only Spacecraft Scenario" gives me the strong understand about Kalman filter and its operation. The attitude sensor at low earth orbit Nano satellite more rely on Magnetometer and the model is implemented based upon Euler angles. This work mainly focuses on the on-board sensor used to estimate the attitude errors with different condition. The major part of work explains the earth-pointing spacecraft undergoing only small rotation angles. In this section, understand the Kalman algorithm and developed and implemented in MATLAB program. It's very important to specify the work presented by Karatas S "LEO Satellites: Dynamic Modelling, Simulations and Some Nonlinear Attitude Control Techniques" [15]. The research contributed the various dynamics and simulation of LEO satellite attitude control methods widely used for small satellites. The research work discussed about the various perturbations in LEO and these affect the attitude and altitude of the satellite and its orientation. The concepts of various actuation systems at like magnetic torquers, Thruster, CMG, Momentum wheels. The actuator requires the input from the controllers, like P-Proportional, I-Integral, D-Derivatives, PI, PD, PID and make the required control torque to the satellite dynamics [16], [17].

This research work encouraged to focus on research topics to design \& analyze of various attitude controls of pitch dynamics, roll dynamics, and yaw dynamics, with a suitable controller incorporated to reduce the oscillation in the system because of the orbital perturbation forces.

### 1.3. Attitude Modeling Tools

The attitude modeling and simulation for satellite systems are classified into two types; one is attitude determination (AD) and another attitude control system (ACS). To analyze the various orbital perturbation forces at Low Earth Orbiting NANO satellites and motion equations in the form of ordinary differential equation (ODE) using Runge - Kutta method was adopted [18]. The International Space Station (ISS), SRM Satellite, Pratham (IITB) Satellite Cowells perturbation algorithm implemented by using MATLAB/Python Environments. The perturbation forces change the satellite six orbital elements like Semi-major axis - distance between the perigee to apogee, Angle of inclination, Eccentricity of orbit, Argument of perigee - angle between perigee to line of nodes where satellite crossing of the equator, south to the north pole, Longitude of ascending node (reference with vernal Equinox), True anomaly/Mean anomaly [19]. The attitude determination parts of LEO system are configured by mathematical modeling and simulation using MATLAB (Version 2014) and Python (Version 3.7) open source software. The Nano satellites attitudes (Pitch, Yaw, Roll) control parts of SRM Satellite, Pratham (IIT Bombay), NPSAT-1 modeled and developed using SIMULINK (Version 2014) control system Toolbox. The analytical model is developed with the help of satellite dynamic equations (refer: https://opensource.gsfc.nasa.gov/projects/GMAT/index.php) and compares the results with General Mission Analysis Tool, GMAT (2014 Version 1.1) open source software for validation. The GMAT is developed by the National Aeronautics and Space Administration (NASA) used for Orbit determination, Visualize the various trajectory optimizations, Mission analysis, attitude maneuvers [20], [21]. The Attitude determination and control system (ADCS) of low earth orbiting NANO satellites are analyzed \& evaluated with MATLAB/SIMULINK and GMAT and Python (32-bit).

### 1.4.Organization of Thesis

The construction of the thesis as tracks
Chapter 1 The first chapter contains the overview of attitude control in satellite systems, research motivation and research questions, modeling tools, and Research summary.

Chapter 2 Covers the literature review of LEO Satellites: Dynamic Modelling, Simulations, Attitude control, Error estimation, Analysis perturbation. Discuss the problem statements about research work, how to design the high pointing accuracy attitude determination and control system (ADCS), Research objectives, Scope of the work.

Chapter 3 Briefly describe the satellite attitude reference frames (Inertial frame, Earth frame, satellite body frame, ECEF Frame, Orbit frame), Geometry of orbits, Keplerian elements, Rotation matrix, DCM, Quaternion, Orbit frame to an inertial frame, Earth frame to an inertial frame, Orbit frame to Satellite body frame.

Chapter 4 Presents the mathematical modeling of satellite dynamics equations and linearization of the equation of motion (EOM), Estimation of the disturbance torques subjected into the satellite, various perturbations (Aerodynamic drag and Solar drag) in low earth orbits. Perturbation models by Runge-Kutta using Cowells methods, analysis of the perturbation of Nano-Satellites (International Space Station, Pratham (IIT) Bombay Satellite, and SRM Satellite in LEO.

Chapter 5 Discuss the attitude estimation using Kalman Filter, Priori state estimation, Posteriori state estimation, Estimation of the satellite (NPSAT-1) attitude errors (Pitch, Roll, Yaw) using a Kalman algorithm with the help of INS/GPS and Magnetometers.

Chapter 6 Present the design of an attitude controller of Nano Satellites (Pratham, SRM satellite, ISS, NPSAT-1) Simulink model of rate controllers with perturbation, Comparison of attitude responses/output from Proportional-Derivatives (PD) controller.

Chapter 7 (Conclusion and recommendation of future work) presents conclusions derived from the work and significant contributions. A brief scope for further research has been identified to provide direction and possible extensions to the work. The reference presents the details of the technical papers referred in this thesis work.

### 1.5.SUMMARY

Introductory chapter presents the overview of the dissertation in terms of its research area, motivation and preliminary research questions. The summary of the research work contains introduction, analysis and characteristics of perturbations and attitude control systems in LOW EARTH ORBIT, Aerodynamic drag, solar drag, Kalman Filter Rate Estimator (KF-RE) and state space theory of attitude sensor [23]. This projects briefly discusses perturbation forces, such as aerodynamic drag and solar disturbances and how it affects six orbital elements or classical orbital elements (COE) angle between the orbital plane and equatorial plane, longitude of ascending node, satellite position with respect to perigee, size of the orbit, perigee to apogee distance, argument of perigee, Mean or Eccentric anomaly for circular orbit [24], [25]. The modeling and simulation tools for the analysis of complex system have been introduced. Attitude sensors such as INS/GPS, IMU, Magnetometer, Rate Gyros, SUN sensors, Earth Sensors, Horizon sensors are used for satellite attitude determination [26]. In this thesis low earth orbit attitude sensors like INS/GPS, IMU and magnetometers have been used for NANO satellites attitude determination. The high Earth orbits (HEO) like Geo synchronous earth orbit (GEO) satellites are heavy satellite. The magnetic torquers are used to control the attitude of low earth orbiting NANO satellites. This is also referred as magnetic actuators. The control momentum gyroscope (CMG) is widely used for satellite attitude control at high earth orbits satellite [27]. The magnetometers widely used for attitude control. It consists of solenoid coil or bar magnets, such as permanent magnet or an electromagnet [28]. The coil or conductor wounded in a core material is mounted into the satellite body frame. The current carrying the conductor produces the magnetic flux which can interact with Earth magnets and generate the controlled forces $(\tau=\mathrm{mB}$, where $\tau$ is torque, $\mathrm{m}=$ NIA is dipole and B is GEO-magnetic field) [29]. The torque generated in the device is based upon the number of turns $(\mathrm{N})$ of the conductors and the amount of current (I) passing through the conductor, cross section area (A) of the conductor. The reaction forces generated in the coil creates the magnetic flux which interacts with the earth magnetic fields used to maintain the satellite into the desired orbits [29]. For heavy or large satellites Thruster, Momentum wheels, fly wheels are used to control the actuators of satellite dynamics. The algorithm of LEO attitude determination and
control mainly based upon satellite attitude sensors INS/GPS, Magnetometer, Inertial Units. The Attitude control algorithms mainly using magnetometer and Rate Gyro's at low earth orbiting small/Nano satellite. The accuracy and precision of Rate Gyro's are highly reliable at low altitude with time [30]. The magnetometer gives proper attitude data when satellite crosses to near earth's atmosphere and the equator. The earth's atmosphere consists of various magnetic belts like Inner Van Allen Belts (IVAB) and Outer Van Allen Belts (OVAB). The High earth orbit such as GEO synchronous orbit attitude control using control momentum gyroscope consist of small motor attached to the satellite body [30]. The inertia of the satellite $\left(\mathrm{I}_{\mathrm{xx}}, \mathrm{I}_{\mathrm{yy}}, \mathrm{I}_{z z}\right)$ measures with the help of disc attached to the motor. The satellite on-board attitude sensors measure the orientation information of the vehicle compare with pre-determine attitude based on that it produces the error signals to actuate the actuator with the help of suitable controllers. The controllers (Proportional P, Integral I, Derivative D, PI, PD, PID) create required signal to the actuators. The CMG act as an actuator which produces the control signals to satellite model. Mostly, the motor is used to deflect the control surface of the satellite model to re-orient from perturbed path to actual orbit path [30]. The proposed method is used to design, mathematical modeling, Simulation of analyzing the perturbation in LEO, Rate control, Attitude error estimator using Kalman Filter developed by MATLAB/SIMULINK and GMAT package.

The Kalman filter is widely used in Global position system/Inertial Navigation system to estimate errors in the autonomous based navigation system. The main part of this research works to estimate the attitude errors from the sensor. The Kalman filter algorithm is used to find angular rates and minimize the errors in the satellite system. This is recursive in nature. The probabilistic theory white Gaussian or Normal distribution methods are used to minimize the error covariance matrix and standard deviation. The Kalman filter interacts with the satellite control and navigation components such as Rate Gyro's and Magnetometer.

## CHAPTER 2

## LITERATURE REVIEW

### 2.1. Outline of Research Problem

A perturbation is a deviation from some typical or expected movement. These irritations, or varieties in the orbital components, can be arranged in view of how they influence the Keplerian components. In orbit, the common varieties speak to a straight variety in the component, brief period varieties are occasional in the component with a period not as much as the orbital period. The satellite orbit changes due to the perturbation forces affecting on it. It requires the highest pointing accuracy attitude control and the determination system which is used to keep the satellite into the predicted trajectory. The attitude control design needs a suitable controller to trigger the actuator for the required demands. At LEO, satellite control relies on magnetic torquers, which can interact with GEO (Earth) magnetic field as the resultant torque is used to control the Nano or Small satellites.

Fischel1, R. E, (1963) "Passive Magnetic Attitude Control for Earth Satellites", Advances in the Astronautical Sciences, Volume 11, Western Periodical Company Hollywood, Calif. The research worked on the vertical stabilization scheme which has been incorporated in the low earth orbit [31]. The research discussed about the satellite attitude control using magnetic actuator (or) magnetic torquers in low earth orbiting satellite. The magnetic torquer consists of solenoid coil, the current ' $I$ ' in the coil produce the magnetic flux which interacts with GEO magnetic field vectors. The idea was also put through by the research how this magnetic field generates the control torque to controlling the satellite with suitable algorithm/control laws [31]. The control torque from magnetic torquers perpendicular to the Earth field vectors (or) GEO magnetic fields. The attitude control system of small satellite relies on magnetometer signal.
S.K Shrivastava, (1976) "Effects of solar radiation pressure and aerodynamic forces on satellite attitude Dynamics and their utilization for control" [32], The research explains the relative magnitudes of torques due to various forces arising from gravitygradient, solar radiations, earth and earth-reflected radiations atmospheric and magnetic forces, cosmic dust, etc., depend on the orbital elements and the satellite's shape, size, surface conditions, mass distribution and orientation. The work presents the satellite dynamics and modelling based on the orbital perturbation aerodynamic drag and solar disturbances [33]. The major part of the work discussed upon aerodynamic drag and solar radiation perturbations are the major perturbations, while in low earth orbit satellites the perturbation which has the maximum magnitude is an aerodynamic drag perturbation because the satellite orbit is too close to Earth's atmosphere which is the primary source of perturbation.
M. D. Shuster and S. D. Oh, (1981) "Attitude Determination from Vector Observations", Journal of Guidance and Control, Volume 4: Page No.70-77, the research work explains the different attitude determination techniques, the differences between them and finally contributed towards the selection of most efficient method for attitude determination [34]. Also, the work discussed the mathematical models to collect the inertial frame of reference \& the vector component in the satellite body. The work based upon mathematical modeling of satellite dynamics is modelled using Euler's equations for a rigid body motion under the influence of internal and external torques. The author discussed the mathematical modelling of satellite dynamics are modelled by satellite rigid body dynamics with Euler angle equations. The angular motion of satellite changes due to internal disturbances and external disturbances. The attitude these components are used, typically in the form of a quaternion, Euler angles. It takes at least two vectors to estimate the attitude [35]. The main area of the work singular results for certain rotations and the Unit Quaternions are computationally less intense. Therefore, Unit Quaternions are more efficient method to be used for attitude determination.

Toshio Fukushima, (1996) "Generalization of Encke's Method and its Application to The Orbital and Rotational Motions of Celestial Bodies", The Astronomical Journal, National Astronomical Observatory, 2-21-1 Ohsawa Mitaka, Tokyo 181, Japan,

Volume 112, Number 3, Received 1995 November 1; revised 1996 May 28. This paper described the several differential equation formulations such as Encke's method and their comparison of a given set of perturbations for a Low Earth Orbit satellite and finally contributed towards the selection of most efficient and accurate method in determining the exact perturbations [36].

WH Steyn, (2001), "Comparison of Low-Earth Orbiting Satellite Attitude Controllers Submitted to Controllability Constraints", the author worked on the Keplerian orbit how the satellite attitudes change with perturbation forces affecting on it [37]. The works present the magnetic torque control the satellite dynamics. The attitude control mainly introduces in the satellite orbital plane only not in the equatorial plane. It was further proposed that if the magnetic field can be taken as periodic changes in the earth's atmosphere. The stabilization platform is achieved by the magnetic moments produced by the satellite body consists of magnetic torque rods. This field interacts with a GEO magnetic field and produces the control signal to the actuator. In the orbit, Dipole is considered as a non-rotating platform with the combined earth magnetic field [38]. By the application of the above-mentioned theory on the stabilization of small/Nano satellite attitude control is achieved by magnetic torques or magnetic moments.

Jonas Elfving, (2002), "Attitude and Orbit Control for Small Satellites", the author presents various Sensors and estimations are used to predict the satellites current position, velocity, attitude and angular velocity [39]. The current work discusses about the designs pointing accuracy of a satellite when using different sensors and actuators so that a craft does not get too expensive. The work based on experimentally analyzed the satellite minimum solar radiation pressure at higher altitude has more effect on satellite, but its radiation influences aerodynamic drag that we have discussed in the atmospheric drag [39]. The orbital elements in the satellite change due to the periodic variations in solar pressure. For calculating the radiation in solar pressure, it's more important the satellite must focus to the SUN. It's clearly understood the author conclude the more orbital decay because of maximum solar radiation pressure.
A.M. Mohammed, A. Boudjemai, S. Chouraqui, (2006) "Magnetorquer Control for Orbital Maneuver of Low Earth Orbit Microsatellite", This work discussed various permanent magnet stabilization (PMS) and Gravity gradient attitude stabilization (GGS). The PMS method introduced the electromagnet into the satellite body frame. The work concentrates the alignment on the magnetic field in the coil with local magnetic moments. The author presents the satellite orbit rapidly decays because of the current in the electromagnet reduces when it is switched off. The satellite alignment controlled by radio signals commend to the electromagnets [40]. In Gravity gradient attitude stabilization, the Earth's gravity gradient is used to achieve vertical stabilization of a satellite. The GGS method the satellite vertical axis is perpendicular to the local magnetic fields. The author explains the attitude dynamics of the satellite along its axis. The attitude representation using Euler methods is widely discussed. The Euler method is used to measure the axis of rotation in the two axes of the satellite body with respect to the inertial reference frame. The satellite Z axis (or) Symmetry axis along with the direction of electromagnets used to stabilize the satellite frame [40]. The differences in Earth's gravitational pull across the satellite mass due to the minor changes in the distance from earth become a significant source of torques in orbits.

Karatas S, (2006) "LEO Satellites: Dynamic Modelling, Simulations and Some Nonlinear Attitude Control Techniques", A Thesis Submitted to the Graduate School of Natural and Applied Sciences of Middle East Technical University [15]. The author worked on the various types of perturbations and the variation of their magnitude of different orbits, which finally contributed towards the selection of most dominant perturbations for Low Earth Orbit satellites. There are two types of magnitude and density variation in solar maxima time and solar minima time. The variation in the solar cycle range is 200 to 250 kilometers and temperature vary from 600 K to 1150 K . The literature only included the Aerodynamic drag perturbations in the formulation of differential equations.

Diaz, Orlando X, (2010) "Analysis and comparison of extended and unscented Kalman filtering methods for spacecraft attitude determination", the author discussed the Extended Kalman filter (EKF) is used widely for nonlinear error estimation [41].

The objectives of the work minimize probability distribution function and standard deviation. The spacecraft attitude measurements were done using different on-board attitude sensors. The linear variation of sensor measurement KF predicts the future estimates [41]. For nonlinear variations of data from attitude sensors modeled by EKF, UKF and predict the state future estimate.

Wang, P(2010) "Attitude Control of Low-orbit Micro-satellite with Active Magnetic Torque and Aerodynamic Torque", The author presents the satellites use this principle for passive attitude stabilization by deploying gravity gradient booms and somehow dumping initial post-launch angular momentum those spacecraft's can maintain Earthoriented position through their orbit [42]. There are gaseous and liquid particles present in the atmosphere of every planet and it provides resistance or a force of resistance which is termed as drag to the satellite body when passes through it. As the spacecraft comes to this type of the planet's atmosphere, it experiences the drag forces and it's greater during launch and reentry of a spacecraft into the space. The author worked on LEO satellite is in the altitude of 90 to 800 km so the effect of this drag is more, and it will bring down from its nominal orbit [43]. There is some limit satellite enter the atmosphere. It has the effect on the atmospheric density and resulting in an increase in the density increase the atmospheric drag. In altitude 90 Kilometers rapid changes in temperature affect the orientation of the satellite. The author explains the disturbance torque like aerodynamic, solar pressure, magnetic flux, homogeneous of earth [43]. Also discussed the how this disturbance affects in the low earth orbiting satellite and its effects.

Murty S. Challa, (2016) "A Simple Attitude Unscented Kalman Filter: Theory and Evaluation in a Magnetometer-Only Spacecraft Scenario" The author worked on an Unscented Kalman filter to estimate the errors in the satellite from different on-board attitude sensor measurements [14]. The research work presents the various attitude orientation using Euler angle and Quaternion methods. The state space satellite model is developed with plant matrix and measurement matrix with Unscented Kalman Filter. The magnetometer used for attitude measurement with different angles. This work mainly minimizes the error covariance matrix and standard deviation [14].

### 2.2. Problem Statement

From literature reviews, the satellite attitude control system uncertainties are associated with the performance due to dynamic variations in orbital elements, with this background, the current work planned with the objective to study the performance comparison and enhancement of satellite attitude error modeling and control system design. The attitude of the satellite changes due to several parameters such as orbit types, perturbation forces, satellite parameters, space environments, etc. Because of the attitude variation the pointing accuracy of the spacecraft decreases gradually. The lifetime of the satellite depends upon the proper attitude control mechanism drawn from literatures. The control system is not designed properly; it will affect the mission performances and qualitative results. At low altitude is will have an enormous amount of drag due to the area and mass of the satellite and drag coefficient referred as ballistic $\left(\frac{m 2}{C_{d} A}\right)$ components. There are more gravitational attractions because of the earth's surface at LEO; whereas for GEO satellites having less/or no aerodynamic drag. These perturbation forces change the attitude/orientation of the Nano satellite Semi-Major axis, Inclination of the orbit, Eccentricity, Satellite position with Perigee location, Longitude of Ascending Node, Ascending node to perigee location, and circular orbit Mean anomaly or Eccentric anomaly. The satellite needs accurate orientation while Transmit/Receive the data and telemetry. For Example, The Solar panel should be deflecting line of sight with SUN. Communication unidirectional antennas transmit the signal to ground station require an accurate attitude. Remote sensing satellite needs capturing the terrain information through HD camera need accurate pointing accuracy. It's understood to design the accurate attitude control system of satellite very much important to save the mission life and proper data collection between satellites to control station. The satellite is launched from rocket booster. After firing all the fuels, the satellite moves under the influence of the gravitational field. Because of the initial momentum de-dumpling the satellite in 3D space. The satellite separated from rockets; requires the proper attitude control system. Until that can't serve the purpose solar panel not produces the power because of improper orientation towards SUN. The life span of the satellite based upon how much power in the satellite. For design the accurate attitude control system increases,
the satellite mission life and performances. To avoid the space debris or space junks. To find the rate information star tracker fail to measure the rate information during high slew rates $0.5 \mathrm{deg} / \mathrm{Sec}$. The rate gyros measure the accurate rate information at all the times.

### 2.3. Research Objectives and Description

The following are the performance measures considered in the system
a) To analyze the perturbation in Low earth orbit satellite
b) To design and develop the attitude control system for Nano Satellite using suitable controllers.
c) To estimate the attitude errors from on-board sensors and implement the algorithm with Kalman Filter

In the real-world behavior compare with mathematical modeling and simulation are highly recommendable. Matching the results with environmental condition mostly requires in the space industry. The prime objectives for this research to design the high pointing accuracy attitude control for small satellite in low earth orbit. To minimize the errors, deviation and required control torque to command the actuator.

Orbital Perturbation: Low earth orbit perturbations like, Aerodynamic drag, solar drag, etc.

Controller: PD compensator designed for attitude control of Nano Satellite.
Kalman Filter: An attitude error modeling and predicts the position of states and angular rates of the Model.

### 2.4. Scope of the work

The Scope of the thesis is improvement of algorithm (Attitude control/ Attitude Error Estimation) in future applications and complex problems. The attitude control system (ACS) increases the lifetime of vehicle by incorporating the accurate control techniques in the satellite. The Attitude sensor like Rate Gyro's highly reliable for attitude control. The reliability of the sensor reduces with time. The satellite losses
mission effectiveness and performances because of the less pointing accuracy attitude mechanism and control methods. This problem is overcome with suitable controllers and attitude sensors added to the system. This is economically big challenge in the Aerospace Industry. It is important to maintain the accurate pointing accuracy and orientation of the satellite in the orbit. This will save the money and revenue in the space industry.

The proposed future work in the relevant area may be as follows:

1) Perturbation analysis at LEO
2) Developing the attitude control algorithm
3) Attitude error estimation from attitude sensors

The scope of this work analyzes the perturbation in the low earth orbit and how this affects the satellite parameters and orientation. To design the suitable controller for reducing the oscillation in the orbit and keep the satellite more time in the orbit. This project estimates the deviation of satellite from the nominal orbit to disturb or perturbed orbit. That variation measured by attitude sensors in the mission. To generate the counter moments in the actuators to deflect the control in the satellite like magnetic torques. The International Space Station (ISS) changes the altitude and attitude because of perturbation forces which required frequent thruster to keep the satellite into same altitude also in the same orbit. The ADCS is very much important to keep the satellite into the same orbit orientation. The accurate orientation schemes to achieve the most presided data collection of communications antennas, HD cameras, deploy the solar panels. At 1000 KM altitudes there are more changes in the atmospheric environment due to gravitational attractions. It is necessary to generate more control actuation in LEO compare with high earth orbit GEO Synchronous Orbit (above 1000 KM ). The velocity of the low earth orbit is greater than the GEO orbits. The accurate navigation and control techniques mainly increase the mission life and income in the space industry. Attitude control is used to receive the proper data collection form the sensors and defects the solar panel with the same line of sight focus to SUN especially for back of power during eclipse time.

## CHAPTER 3 <br> VEHICLE ORIENTATION IN MICROGRAVITY CONDITION AT LEO

Vehicle orientations is described by coordinate system, two basic methods of coordination system used to represent the vehicle position with respect to inertial axes and vehicle body axes such as Quaternion method and Euler angles methods [44].

### 3.1 Representing Attitude Information

To represent the attitude (Pitch, Roll, and Yaw) of satellite, Euler angles transformation is accurate enough but sometimes singularity occurs in coordinate transformation. Quaternion representations are used as to avoid singularities [45]. To find the angular components (velocity) in satellite body frame with respect to the inertial reference frame, we assume the earth is an inertial reference frame originated with the center of Earth [46].

### 3.1.1 Euler Angles Methods

Euler angle method for describing the alignment of the vehicle body axes to inertial coordinate system [47]. To converts one frame into another frame direct cosine matrix (DCM), it is mostly used. Euler angles are Roll angle $\varphi$ - Satellite Rotates about x -axis, Pitch angle $\boldsymbol{\theta}$-Satellite Rotates about y -axis, and Yaw angle $\boldsymbol{\psi}$ Satellite Rotates about z-axis. The satellite rotations of body axis of the inertial reference axes shown in Figure (3.1). Euler angle has singularity problems as it won't measure the two rotations on the same axes. Overcome this, it is replaced by quaternion orientation to avoid singularity problems [47].


Figure (3.1) Satellite body axis to the inertial axis [47]
The rotation matrix equations are used to finding the position of vehicle frame to reference frame [48]. The satellite frames are $b_{1} \& b_{2} \& b_{3}$. The most conjoint method of rotation is 313 types. First rotations about the body from place to place b3 axes, Second, rotates the body from place to place b1, third rotates the body from place to place b3. Euler angles (See Figure 3.2) requires the latest transformation about satellite body frame to inertial frame. The roll angles $\phi$ deflects satellite about the xaxis, the pitch angles $\theta$ deflects satellite about the y -axis \& the yaw angles $\psi$ deflects satellite about the z axis. Euler angles (3.1) given below [48]

$$
\Theta=\left(\begin{array}{c}
\phi  \tag{3.1}\\
\theta \\
\psi
\end{array}\right)
$$

The rotation matrices (3.2), (3.3), (3.4) are given as follows:
$R_{x, \phi}=\left(\begin{array}{ccc}1 & 0 & 0 \\ 0 & \cos \phi & -\sin \phi \\ 0 & \sin \phi & \cos \phi\end{array}\right)$
$R_{y, \theta}=\left(\begin{array}{ccc}\cos \theta & 0 & \sin \theta \\ 0 & 1 & 0 \\ -\sin \theta & 0 & \cos \theta\end{array}\right)$
$R_{z, \psi}=\left(\begin{array}{ccc}\cos \psi & -\sin \psi & 0 \\ \sin \psi & \cos \psi & 0 \\ 0 & 0 & 1\end{array}\right)$

As a result, the rotation matrix (3.5) $R_{B}^{o}$ converts the body to orbit frame

$$
R_{B}^{o}=R_{z}(\psi) R_{y}(\theta) R_{x}(\phi)\left(\begin{array}{ccc}
c \psi c \theta & -s \psi c \phi+c \psi s \theta s \phi & s \psi s \phi+c \psi c \phi s \theta  \tag{3.5}\\
s \psi c \theta & c \psi c \phi+s \psi s \theta s \phi & -c \psi s \phi+s \psi c \phi s \theta \\
-s \theta & c \theta s \phi & c \theta c \phi
\end{array}\right)
$$



Figure (3.2) Euler Angles representation [48]

$$
\begin{equation*}
\boldsymbol{R}^{(\text {inertial to Body })}=\operatorname{Rmat}_{i}^{T}(\psi) \operatorname{Rmat}_{j}^{T}(\theta) \operatorname{Rmat}_{k}^{T}(\phi) \tag{3.6}
\end{equation*}
$$

In real time applications, Euler angle attitude representation methods is well suited techniques to implement Nano satellite attitude control with on- board attitude sensors such as Inertial navigation system (INS) / Global position system (GPS) and Inertial measurement unit (IMU).

### 3.1.2. Quaternions Method



Figure (3.3) Quaternion diagram of transformation from satellite frame [49]
Quaternions method finds the orientation of the body with the help of Euler axis defined by unit vectors $e_{i}=\left(e_{1} \& e_{2} \& e_{3}\right)$ and the vehicle can be rotated by angle
$\theta$. Both parameters $\left(\mathrm{e}_{\mathbf{i}}, \theta\right)$ used to find the orientation with respect to the inertial frame in space. The dual quaternions have both magnitude and angle. Quaternion doesn't have any singularity problems [49]. It is used to measure the two rotations in same axes.

Table 3.1: Pros and cons of various orientation methods

| Attitudes <br> Transformation <br> Methods | DCM | Euler Angles | Quaternions |
| :---: | :---: | :---: | :---: |
| Advantages | Satellite orientation <br> define by direct <br> cosine matrix | Satellite attitudes <br> (Roll, Pitch, Yaw) If <br> given, the inimitable <br> orientation is <br> defined | No Singularity <br> Problems |
| Dis-Advantages | Six limitation must <br> be met, non- <br> instinctive | Singularity <br> Problems Exist | Requires Transform <br> Techniques |

The proposed research work implements the Quaternion which converts the rotation matrix and transforms it to another frame. In Table 3.1 gives the advantage and disadvantages of various orientation techniques. It has four elements compared to the Euler transformation which has nine elements. The propagation of satellite orientation is calculated by a quaternion. Quaternion method is a more useful transformation from the body frame (satellite) to the inertial frame as compared to the Euler method [50].

Quaternion ( q ) defines as $\mathrm{q}=\xi+\mathrm{i} \epsilon_{l}+\mathrm{j} \epsilon_{2}+\mathrm{k} \epsilon_{3}$
It consists of 4 elements, $\xi$ is real value, $\mathrm{i} \epsilon_{1, \mathrm{j}} \mathrm{j}_{2}$, and $\mathrm{k} \epsilon_{3}$ is imaginary values.

$$
\eta=\cos \frac{\varphi}{2}, \varepsilon=\left[\begin{array}{l}
\varepsilon_{1}  \tag{3.7}\\
\varepsilon_{2} \\
\varepsilon_{3}
\end{array}\right]=\left[\begin{array}{l}
\mathrm{k}_{\mathrm{x}} \sin (\varphi / 2) \\
\mathrm{k}_{\mathrm{y}} \sin (\varphi / 2) \\
\mathrm{k}_{\mathrm{z}} \sin (\varphi / 2)
\end{array}\right] \quad, \quad q=\left[\begin{array}{l}
\eta \\
\varepsilon_{1} \\
\varepsilon_{2} \\
\varepsilon_{3}
\end{array}\right]
$$

The quaternions expressed in the satellite body to orbit (rotation matrix) as: [50]

$$
\begin{equation*}
R_{B}^{0}(\mathrm{q})=R_{\eta, \varepsilon}=\mathrm{I}_{3 \times 3}+2 \eta \mathrm{~S}(\varepsilon)+2 \mathrm{~S}^{2}(\varepsilon) \tag{3.8}
\end{equation*}
$$

Quaternion rotation matrix is shown in equation (3.9) $R_{B}^{o}$ can be written as: [50]

$$
R_{B}^{o}=\left(\begin{array}{ccc}
1-2\left(\varepsilon_{2}^{2}+\varepsilon_{3}^{2}\right) & 2\left(\varepsilon_{1} \varepsilon_{2}-\varepsilon_{3} \eta\right) & 2\left(\varepsilon_{1} \varepsilon_{3}+\varepsilon_{2} \eta\right)  \tag{3.9}\\
2\left(\varepsilon_{1} \varepsilon_{2}+\varepsilon_{3} \eta\right) & 1-2\left(\varepsilon_{1}^{2}+\varepsilon_{3}^{2}\right) & 2\left(\varepsilon_{2} \varepsilon_{3}-\varepsilon_{1} \eta\right) \\
2\left(\varepsilon_{1} \varepsilon_{3}-\varepsilon_{2} \eta\right) & 2\left(\varepsilon_{2} \varepsilon_{3}+\varepsilon_{1} \eta\right) & 1-2\left(\varepsilon_{1}^{2}+\varepsilon_{2}^{2}\right)
\end{array}\right)
$$

When compared to both attitude determination techniques (Euler \& Quaternion), Unit Quaternion method is more commonly used, because this method is not singular for any rotations, while Euler Angles can give singular results for certain rotations. The Unit Quaternions are computationally less intense [50]. Therefore, unit quaternions are more efficient method used for attitude determination.

It has both real and complex numbers coveting the vector from quaternion operation having 30 floating value and 45 processes to transform from a quaternion to a matrix. It was first defined by William Rowan Hamilton. It avoids the singularity difficult in Euler angle transformation from one frame to another frame. The above diagram (Figure 3.3) discuss the two reference frames ( $\mathrm{x}_{\mathrm{A}}, \mathrm{y}_{\mathrm{A}}, \mathrm{z}_{\mathrm{A}}$ ) and ( $\mathrm{x}_{\mathrm{B}}, \mathrm{y}_{\mathrm{B}}, \mathrm{z}_{\mathrm{B}}$ ), satellite rotates about z -axis. In case both vectors in both the frames are same, but different value in $x$ component and the $y$ component [51]. The vector ' $u$ ' can be described in any frame. For example, vector $u$ has same length and it is significant to describe the quaternion. Quaternion frame 1 to frame 2 converts a vector. Frame 'A' vector denoted to Frame ' B ' vector. Quaternion defines the magnitude and direction.

This same quaternion might be notated instead as $q B A$, where the order of the subscripts now indicates that the quaternion converts the position and velocity in the frame ' $A$ ' vector quantity into frame ' $B$ ' vector quantity. This form can be helpful when combining quaternions in order as discussed below. Both are used in practice and it is important to always verify the rotation represented.

### 3.1.3. Geometrical Definitions

The transformation (See Figure: 3.4) from an inertial reference (I, J, K) frame to local orbital frame ( $\boldsymbol{x}_{\mathrm{o}}, \boldsymbol{y}_{\mathrm{o}}, \boldsymbol{z}_{\mathrm{o}}$ ) is performed by means of the Euler rotation sequence via the rotation angles $(\Omega, i, u)$ [53]. The latitude angles denote, $u=$ argument of perigee ( $\grave{\omega}$ )+ true anomaly (v). The angle of inclination is given by $i$ (rad/sec), and vernal equinox with respect to the perigee is given by $\Omega$ [53].


Figure (3.4) Inertial Coordinate (I, J, K) and Orbital Coordinate [53]
The local orbital coordinates (See Figure 3.5) reference ( $\boldsymbol{x}_{\mathrm{o}}, \boldsymbol{y}_{0}, \boldsymbol{z}_{\mathrm{o}}$ ) system and Earth Frame is satellite frame is attached to the center of the body called as reference Frame ( $\left.\boldsymbol{x}_{\mathrm{b}}, \boldsymbol{y}_{\mathrm{b}}, \boldsymbol{z}_{\mathrm{b}}\right)$. The nominal attitudes RF is along the orbital reference axes; $\boldsymbol{x}_{\mathrm{b}}=\boldsymbol{x}_{\mathrm{o}}$, $y_{b}=y_{o}, z_{b}=z_{0}$. In this case, the $y_{b}$ axis is nominally along the velocity direction $\boldsymbol{V}$ [54].



Figure (3.5) (a) Orbital Frame and (b) Body centered Reference frame [54]

The proposed methods study the series of transformations from the orbital coordinate system to the reference frame is 3-2-1 via the (Yaw, Pitch, Roll) angles. [54]


Figure (3.6) Satellite reference frame to orbital reference frame [55]

For, Satellite rotates about the
Satellite rotation about the $\mathrm{x}_{0}$ axis called as $\phi$ (Roll angle)
Satellite rotation about the $y_{o}$ axis called as $\theta$ (Pitch angle)
Satellite rotation about the $\mathrm{z}_{\mathrm{o}}$ axis called as $\psi$ (Yaw angle)
Consider the sketch of Figure 3.6, where $z$-axis indicates yaw rotations and the $y$ axis indicates pitch rotations, and $x$-axis indicates roll rotations [55].

### 3.1.4. Orientation of the Satellite

Coordinate frame (CF) is used to find the location of the vehicle with respect to the orientation of reference frame as shown in Figure. 3.7. Here, satellite frame is called as a rotating body frame [56]. The following assumption are considered for the simulations, Earth is an inertial reference frame which is fixed. It denotes as a nonrotating frame [56].


Figure (3.7) Different types of coordinate frame [56]
Earth-Centered Earth Fixed Frame (ECEF): The ECEF is located at the center of earth. The X -axis \& Y-axis rotates in the Earth center inertial (ECI) frame or nonrotating frame [57]. We consider earth is an inertial reference frame. The Z-axis located in North Pole. The X-axis crosses among the Greenwich Meridian (GM) and the Equator. This point both longitude and latitude are considered as $0^{\circ}$ degree. The Y -axis considers as right-hand coordinate system [58].

Satellite Body Frame: The body frame is located at the satellite body mass in the center. The body frame attached to the vehicle frame. The Sz-axis located in the Nadir
direction of the satellite [57]. The Sx-axis and Sy-axis crosses the orbit frame when the attitude of satellite referred as $0^{\circ}$ degree [57]. In this case all the attitude of satellite pitch, yaw, and roll angles become zero.

Earth center Inertial (ECI) Frame: ECI frame is fixed in space. The Ix-axis located from vernal equinox at satellite concentration on the Earth's surface. This Iz-axis represent the angular velocity direction of the orbit. The axis Iy is orthogonal to Ix and Iz [58].

Orbital Frame: The orbit frame originates with satellite (vehicle) center of the mass. The Nadir position is represented by the axis $\mathrm{O}_{\mathrm{z}}$ [57]. This point, the satellite concentration of the earth's surface. The axis Ox represents Satellite motion in the body. Also, axis ox $\perp$ oz. The axis Oy represents the complete right-hand coordinates [58].

### 3.2. The Earth-Satellite System (Satellite Equation of Motion)

We consider the mass $\mathrm{m}_{1}$ referred as the earth mass and $\mathrm{m}_{2}$ as the satellite mass; Newton law of gravitation states that the force of attraction increases when Earth mass and satellite mass increase. Forces of attraction reduce when the distance between the two masses increase [24].


Figure (3.8) Gravitational force between two masses [24]

Where: F1=F2= Attraction forces between two bodies (Large body =Earth; Small body = Satellite), N

$$
\begin{gathered}
\mathrm{G}=\mathrm{Gravitational} \text { element (constant), }\left(\mathrm{G}=6.674 * 10^{-11} \mathrm{~N} \mathrm{~m}^{2} \mathrm{~kg}^{-2}\right) \\
\mathrm{m}_{1}=\text { Consider earth mass, large mass, } \mathrm{Kg} \\
\mathrm{~m}_{2}=\text { Consider flying object or satellite, small mass, } \mathrm{Kg} \\
\mathrm{r}=\text { Line connecting from the earth center to satellite center, } \mathrm{m}
\end{gathered}
$$

Two bodies problem considers the Earth and satellite body. This is the easy methods of N-Body problems. Large mass (Earth) considers $\mathrm{m}_{1} \&$ Small mass (satellite) consider $\mathrm{m}_{2}$. The n -body equation becomes [24]

$$
\begin{equation*}
F_{1}=m_{1} r_{1}=G m_{1} m_{2} \frac{r_{2}-r_{1}}{\left|r_{2}-r_{1}\right|^{3}}(3.10) \quad F_{2}=m_{2} r_{2}=G m_{1} m_{2} \frac{r_{1}-r_{2}}{\left|r_{1}-r_{2}\right|^{3}} \tag{3.11}
\end{equation*}
$$

Combining the equations (3.10) \& (3.11) gives: $F_{1}=m_{1} r_{1}=$
$G m_{1} m_{2} \frac{r_{2}-r_{1}}{\left|r_{2}-r_{1}\right|^{3}}$

And with $\mathrm{r}=\mathrm{r}_{2}-\mathrm{r}_{1} ; \quad \ddot{r_{2}}-\ddot{r_{1}}=-G\left(m_{1}+m_{2}\right) \frac{r_{1}-r_{2}}{r^{3}}$
Hence the equation (3.13) is the satellite motion equations of two body.( Primary body, Earth - Secondary body, Satellite) problems [24].

### 3.2.1. Geometry of Satellite Orbits

The orbit of the satellite follows the geometry conic section. The conic sections consist different joint of plan and cone. Sections are circle, ellipse, parabola, Hyperbola. Circle connects horizontal line, ellipse joins incline slope in the cone, see Figure 3.9. Both curves intersect and make a closed path. The consequential path joins/connect the Hyperbola [57].


Figure (3.9) Conic Section (Circle, Ellipse, Parabola, Hyperbola) [57]

The parabolic curve is single margin as it creates the elliptical path from the hyperbolic path with parallels to the conic section. The peri-center \& apo-center, two points connect the orbit, peri-center is a point where satellite is closest to the object orbiting the Earth's surface.

In Figure 3.10 (a) shows the distance from apogee to perigee point, (SMA) a, Size of the orbit (Ecc) e, apo-center it's a point where satellite furthest away from the object orbiting the earth's surface. In Figure 3.10 (b) shows the true anomaly, angle between vernal equinoxes to ascending node $(\Omega)$, angle between perigee to line of nodes from south to north pole, $(\omega)$ ae is the distance between the centers of earth with respect to center of the focal point. It is determined by the eccentricity of the orbit with respect to the conic section. The eccentricity determines the type of orbit obtained [24]. In Table 3.2 shows the Keplerian parameters.


Figure (3.10) (a) Satellite (major/Minor) axis [57]

(b) Satellite position [26]

The orbital eccentricity to determine the shape of the orbit, also other Keplerian element is mentioned in the Figure (3.10). To find the position of the satellite by the knowledge of the angular shift from perigee point to satellite velocity vector direction.

Table 3.2: Description of Keplerian Elements

| Elements | Name | Description |
| :--- | :--- | :--- |


| A | Semi-major axis | See Figure 3.10 (a) |
| :---: | :---: | :---: |
| E | Eccentricity | When multiplied with a, it gives the distance from the epicenter of the orbit to the principal point |
| I | Inclination | The angle between the equator plane and the orbit plane |
| $\Omega$ | Vernal Equinox <br> (VE) w.r.t (RAAN) <br> Ascending node | Satellite moving from south to north path crosses to the Equator with VE. |
| $\Omega$ | Argument of perigee | Describes the orientation of the orbit |
| N | True anomaly | The angle between the satellite position to the perigee position |

The size of the orbit determines by semi-major axis (SMA) [24]. The size and shape of the orbits represent by Keplerian orbital elements. These elements are mostly used to design the specific characteristics of the orbits (Circular or Elliptical). Also, it describes the orbital motion include mean orbital rates of the satellite

## Table 3.3 Orbits with corresponding eccentricities [24]

| Eccentricity | Orbit |
| :--- | :--- |
| $\mathrm{e}=$ zero | Circular path |
| The value 'e' |  |
| between 0 to 1 |  |
| $\mathrm{e}=$ one |  |
| e greater than 1 |  |$\quad$| Elliptical path |
| :--- |
| Parabolic path |
| Hyperbolic path |

In Table 3.3 describe the path of an orbit accurately. The six orbital elements used to fully define an orbital motion of the satellite. The proposed methods discuss the Runge - Kutta Numerical integration to solve ordinary differential equations (ODE) of the orbital motion.

### 3.2.2. Satellite coordinates and Keplerian equations

In this section (3.2.2) discusses the, classical orbital parameter that is often used is mean anomaly M. In this thesis 'M' used instead of true anomaly $v$ (Low Earth Circular Orbits). M is defined by equation [59]

$$
\begin{equation*}
M=\epsilon-e * \sin \epsilon,(3.14) \quad \cos \epsilon=\frac{e+\cos * v}{1+e * \cos * v} \tag{3.15}
\end{equation*}
$$

The velocity vectors

$$
\begin{equation*}
\overrightarrow{\text { velocıty }}=\dot{r} \overrightarrow{e_{r}}+r \theta \overrightarrow{e_{\theta}} \tag{3.16}
\end{equation*}
$$

Acceleration vector $\quad \overrightarrow{\text { acceleration }}=\left(\ddot{r}-r \dot{\theta}^{2}\right) \overrightarrow{e_{r}}+(r \ddot{\theta}+2 \dot{r} \dot{\theta}) \overrightarrow{e_{\theta}}$
Hence, the motion equation (3.16) \& (3.17) into circular and diagonal direction


Figure (3.11) Polar coordinates (r, $\boldsymbol{\theta}$ ) vehicle equation [59]

There are many coordinate systems which explain the satellite motion in the body frame w. r .t inertial frame of reference. Polar coordinate system is easy to describe the orbital mechanics [37]. It has two quantities ( $r, \theta$ ) unit vectors are $e_{r}$ and $\mathrm{e}_{\theta}$.

In the circular direction the motion equation is

$$
\begin{equation*}
\ddot{\boldsymbol{r}}-\boldsymbol{r} \dot{\boldsymbol{\theta}}^{2}=-\frac{\mu}{r^{2}} \tag{3.18}
\end{equation*}
$$

Considered for the simulation's Gravitational parameter, $\mathbf{G}$ and Mass of the satellite $\mathbf{m}$ then constant $\mu$ is a product of the gravitational constant and the mass of the vehicle Let, $\mu=G * m$ this equation expresses mainly acceleration in circular direction [25]. In motion equation diagonal direction

$$
r \ddot{\theta}-2 \dot{r} \dot{\theta}=0
$$

This can be restated as

$$
\frac{1}{r} \frac{d}{d t}\left(r^{2} \dot{\theta}\right)=0
$$

$r^{2} \theta^{\prime}$ is constant. $r^{2} \theta$ is equal to the angular momentum per unit mass; $h$

$$
\begin{equation*}
\ddot{r}-\frac{h^{2}}{r^{3}}=-\frac{\mu}{r^{2}} \quad r(\theta)=\frac{\frac{h^{2}}{\mu}}{1+\frac{A h^{2}}{\mu} \cos \left(\theta-\theta_{0}\right)} \tag{3.19}
\end{equation*}
$$

Where, A and $\theta_{0}$ are constants. Hence, the equation (3.19) has denoted the polar coordinates of satellite motion [25].

## Vectors \& Matrices

Notation: Vectors are used to represent the magnitude and direction in 3D space. Vector is connected to the coordinate frame of reference to the Matrix. Vector is the value and its quantity which is used to describe the frame. Matrixes represent the axis ( $\mathrm{x}, \mathrm{y}, \mathrm{z}$ )

## The Matrix Riccati equations

$C=P \dot{C}+C P^{T}-C M^{T} R^{-1} M C+P n$
Which relates to the state covariance C ; to the plant model P; plant spectral density matrix Pn ; the measurement matrix M ;

The equation (3.20) of an ellipse is $r(v)=\frac{a *\left(1-e^{2}\right)}{1+e * \cos * v}$

The satellite position with respect to center of object is orbited to Earth considers as $\mathbf{r}$, semi-major axis a, See Figure 3.12, orbital eccentricity e \& true anomaly is $\mathbf{v}$. It is simplified at the peri-center and apocentre. [24]


Figure (3.12) Keplerian elements with satellite position in the orbit [24]

At the peri-center $v=0 \& r_{p}=a(1-e)$
At the apo-center $v=\pi \& r_{a}=a(1+e)$
The orbit is found by identifying

$$
\begin{equation*}
h^{2}=\mu\left(1-e^{2}\right) \tag{3.21}
\end{equation*}
$$

At the pericentre, angular momentum per unit mass, $h$, is $r_{p} v_{p}$. As $h$ is conserved,

$$
\begin{equation*}
a^{2}(1-e)^{2} v_{p}^{2}=\mu a\left(1-e^{2}\right) \tag{3.22}
\end{equation*}
$$

This results in the equation (3.22) for peri-center velocity (3.23)

$$
\begin{equation*}
v_{p}=\sqrt{\frac{\mu 1+e}{a 1-e}} \tag{3.23}
\end{equation*}
$$

Similarly, speed in the apocentre is found to be total energies (See equation 3.24)) is expressed as combining potential energy (P) and kinetic energy (K); $\mathbf{E}=\mathbf{P}+\mathbf{K}$, the total energy E per unit mass is [24]

$$
\begin{equation*}
E=\frac{1}{2} v^{2}-\frac{\mu}{r_{\text {earth }}} \tag{3.24}
\end{equation*}
$$

At the peri-center

$$
E=\frac{1}{2}\left(\frac{\mu 1+e}{a 1-e}\right)-\frac{\mu}{a(1-e)}
$$

The energy equation (3.25) can be restated as

$$
\begin{equation*}
E=-\frac{\mu}{2 a} \tag{3.25}
\end{equation*}
$$

Substitute equation (3.25) to (3.24)

$$
\frac{1}{2} v^{2}-\frac{\mu}{r_{\text {earth }}}=-\frac{\mu}{2 a}
$$

Hence, the equation (3.26) of velocity on an elliptical orbit

$$
\begin{equation*}
v^{2}=\mu\left(\frac{2}{r_{\text {earth }}}-\frac{1}{a}\right) \tag{3.26}
\end{equation*}
$$

The time periods of the orbit (3.27) calculated from the cube of the semi-major axis and gravitational parameter and its results

$$
\begin{equation*}
\text { Time periods }=2 \pi \sqrt{\frac{a^{3}}{\mu}} \tag{3.27}
\end{equation*}
$$

## Circular Orbits, Parabolic Orbits \& Hyperbolic Orbits:

For a circular orbit, Radius $\mathrm{r}_{\text {earth }}$ is constant for all; the point in the orbit, Eccentricity of the orbit is remaining zero, the circular velocity (3.28) and time periods (3.29) equations of the circular orbit given by [25]

$$
\begin{align*}
& \text { Circular velocity }=\sqrt{\frac{\mu}{\mathrm{r}_{\text {earth }}}}  \tag{3.28}\\
& \text { Time Period }=2 \pi \sqrt{\frac{\mathrm{r}_{\text {earth }}{ }^{3}}{\mu}} \tag{3.29}
\end{align*}
$$

For parabolic orbit eccentricity is one, Velocity equation (3.30) is given by

$$
\begin{equation*}
\text { velocity }=\sqrt{\frac{2 \mu}{r_{\text {earth }}}} \tag{3.30}
\end{equation*}
$$

The time taken to complete one orbit, $\mathrm{T} \rightarrow$ infinity since the semi major axis are $\rightarrow$ infinity [39].
For, eccentricity in the orbit is (hyperbolic path) e $>1$. The velocity equation (3.31) is given by

$$
\begin{equation*}
v^{2}=2 \frac{\mu}{r_{\text {earth }}}+V_{\infty}^{2} \tag{3.31}
\end{equation*}
$$

Where, free stream velocity (3.32) expressed as

$$
\begin{equation*}
\mathrm{V}_{\infty}=\sqrt{\frac{\mathrm{GM}}{\mathrm{a}}} \tag{3.32}
\end{equation*}
$$

Locating/Position of the Satellite Orbit:
To find the position vectors (3.33) of object (satellite) with respect to perigee [24].

$$
\begin{equation*}
\mathbf{r}_{\text {initial }}=\frac{\mathrm{a}\left(1-\mathrm{ecc}^{2}\right)}{1+\mathrm{ecc} * \cos \mathrm{~T}_{\text {anomaly }}} \tag{3.33}
\end{equation*}
$$

Where:

- ecc: eccentricity of the orbit
- a: line connecting from apogee and perigee (Semi-major axis)
- $\mathrm{r}_{\text {initial }}$ : radius from the foci of the planet
- $\mathrm{T}_{\text {anomaly }}$ : True anomaly (measure of the angle from the perigee to the position of the satellite

To Calculates the flight path angle (3.34) and velocity (3.35) of spacecraft by following relations [24]:

$$
\begin{align*}
& \text { Flight path angle }=\arctan \left(\frac{\mathrm{ecc} * \sin \mathrm{~T}_{\text {anomaly }}}{1+\mathrm{ecc} * \cos \mathrm{~T}_{\text {anomaly }}}\right)  \tag{3.34}\\
& \text { Velocity }=\sqrt{G * \mathbf{m}_{\mathbf{1}}\left(\frac{2}{r_{\text {earth }}}-\frac{1}{\mathrm{a}}\right)} \tag{3.35}
\end{align*}
$$

Energy equation for satellite:

The Energy equation (3.36) of the satellite is calculated by the difference between P and $K$. From the equation mass of the vehicle/satellite $\mathbf{m}_{\mathbf{2}}$, and satellite velocity as $\mathbf{V}$, and Semi-major axis as, a [25]

$$
\begin{equation*}
E=\frac{m_{2} V^{2}}{2}-\frac{G m_{2}}{a} \tag{3.36}
\end{equation*}
$$

The circular velocity (3.37) of an orbit around an object is defined as [25]:

$$
\begin{equation*}
\mathbf{V}_{\text {circular }}(\mathbf{V c})=\sqrt{\frac{\mathbf{K}^{2}}{\mathrm{r}_{\text {earth }}}}=\sqrt{\frac{\mathbf{G} * \mathrm{~m}_{1}}{\mathrm{r}_{\text {earth }}}} \tag{3.37}
\end{equation*}
$$

Where, for Earth
$\mathrm{K}^{2}=$ Gravitation constant $*$ Mass of the earth $=3.98 * 10^{14} \mathrm{~m}^{3} / \mathrm{s}^{2}$, Earth radius $\left(\mathbf{r}_{\text {earth }}\right)$ $=6.37 * 10^{6} \mathrm{~m}$, Hence, for escape from earth into circular orbit need velocity of 7.9

## km / Sec

Time Periods Calculation of a Satellite given in equation (3.38)

$$
\begin{equation*}
\mathrm{T}=\left(2 \Pi \mathrm{r}_{\mathrm{earth}}{ }^{\frac{3}{2}}\right) /(\mathrm{K}) \tag{3.38}
\end{equation*}
$$

### 3.2.3. The Three-Body Problem

The controlled 3-body problem is a very good way to describe the forces between Earth, the Moon and a satellite. It consists of a system that includes three masses moving in a plane. Let us assume that Earth has mass $\mathrm{m}_{1}$, the Moon mass $\mathrm{m}_{2}$ and the satellite mass $m_{3}$. Mass $m_{3}$ is a lot smaller than $m_{1}$ and $m_{2}$, so it can be neglected [24]. The law of gravitation gives gravity force $F^{\sim 1}$ on Earth from the Moon and gravity force $\mathrm{F}^{\sim 2}$ the opposite way. They are given in (equation 3.39)

$$
\begin{equation*}
F_{1}=-F_{2}=k^{2} \frac{m_{1} m_{2}}{L^{2}} b_{1} \tag{3.39}
\end{equation*}
$$

Let $\mathrm{k}=$ Gaussian parameter of gravitation $\& \mathrm{~L}$ is the distance between body 1 and 2 . The vector from Earth' center to the Moon center of rotates with an angular velocity $\omega_{\text {earth } / \text { moon }}=\omega_{b 3}$ Earth has position $R_{1}=-x_{1} b_{1}$ And $R_{2}=-x_{2} b_{2}$ The Moon has Position $L$ Vector is given by $\mathrm{L}=\mathrm{x} 1+\mathrm{x} 2$. [24]

The acceleration becomes in equations (3.40) and (3.41)

$$
\begin{align*}
& \ddot{a}_{1}=\omega_{\text {earth } / \text { moon }} X\left(\overrightarrow{\omega_{\text {earth } / \text { moon }}} X \vec{R}_{1}\right)=\omega^{2} x_{1} \overrightarrow{b_{1}}  \tag{3.40}\\
& \ddot{a}_{2}=\omega_{\text {earth } / \text { moon }} X\left(\overrightarrow{\omega_{\text {earth } / \text { moon }}} X \vec{R}_{2}\right)=\omega^{2} x_{2} \overrightarrow{b_{1}} \tag{3.41}
\end{align*}
$$

The gravitational and centrifugal forces are in balance. This gives

$$
k^{2} \frac{m_{1} m_{2}}{L^{2}}=m_{1} x_{1} \omega^{2}=m_{2} x_{2} \omega^{2}
$$

And from this Kepler's third law is found at

$$
\omega^{2}=\frac{k^{2} M}{L^{3}}
$$

Where $\mathrm{M}=\mathrm{m} 1+\mathrm{m} 2$.

The position (3.42) of satellite is

$$
\begin{equation*}
r=x \overrightarrow{b_{1}}=y \vec{b}_{2} \tag{3.42}
\end{equation*}
$$

The velocity (3.43) of satellite is

$$
\begin{equation*}
\vec{v}=\frac{d \vec{r}}{d t}+\vec{\omega}_{i b} X \vec{r}=\dot{x} \overrightarrow{b_{1}}+\dot{y} b_{2}+\omega\left(x \overrightarrow{b_{2}}-y \overrightarrow{b_{1}}\right) \tag{3.43}
\end{equation*}
$$

And the acceleration becomes (3.44)

$$
\begin{equation*}
\vec{a}=\frac{d^{2}}{d^{2} t} \vec{r}+2 \overrightarrow{\omega_{l b}} X \frac{d}{d t} \vec{r}+\vec{\alpha}_{i b} X \vec{r}+\overrightarrow{\omega_{l b}} X\left(\overrightarrow{\omega_{l b}} X \vec{r}\right) \tag{3.44}
\end{equation*}
$$

The motion (3.45) of the moving vehicle can be described as [24]

$$
\begin{equation*}
\overrightarrow{F_{3}}=-k^{2} \frac{m_{1} m_{3}}{r_{1}^{3}}\left\lfloor\left(x+x_{1}\right) \overrightarrow{b_{1}}+y \overrightarrow{b_{2}}\right\rfloor-k^{2} \frac{m_{2} m_{3}}{r_{2}{ }^{3}}\left\lfloor\left(x+x_{2}\right) \overrightarrow{b_{1}}+y \overrightarrow{b_{2}}\right\rfloor, \tag{3.45}
\end{equation*}
$$

Where,

$$
\overrightarrow{r_{v 1}}=\sqrt{\left(x+x_{1}\right)^{2}+y^{2}}, \quad \overrightarrow{r_{v 2}}=\sqrt{\left(x+x_{2}\right)^{2}+y^{2}}
$$

In x and y direction this results in

$$
\begin{gathered}
\ddot{x}-2 \omega \dot{y}-\omega^{2} x=-k^{2}\left[\frac{m_{1}}{r_{1}{ }^{3}}\left(\left(x+x_{1}\right)+\frac{m_{2}}{r^{3}{ }_{2}}\left(x-x_{2}\right)\right]\right. \\
\ddot{y}+2 \omega \dot{x}-\omega^{2} y=-k^{2}\left(\frac{m_{1}}{r_{1}{ }^{3}}+\frac{m_{2}}{r_{2}{ }^{3}}\right)
\end{gathered}
$$

This model is usually presented in normalized form where the distances are divided by L and $\tau=\omega \mathrm{t}$. [24]

The General N-Body Problem:
The system involving of many bodies, the summation of all forces acting on the $\mathrm{i}^{\text {th }}$ body (3.46) [24]

$$
\begin{equation*}
F_{i}=G \sum_{j=1}^{j=n} \frac{m_{i} m_{j}}{r_{i j}^{3}}\left(r_{j-} r_{i}\right), \quad \text { i not equal to } j ; \tag{3.46}
\end{equation*}
$$

It follows from Newton's $2^{\text {nd }}$ law of motion in equation (3.47)

$$
\begin{equation*}
\frac{d^{2} r}{d x^{2}}=G \sum_{j=1}^{j=n} \frac{m_{i}}{r_{i j}^{3}}\left(r_{j-} r_{i}\right), \quad i \neq j \tag{3.47}
\end{equation*}
$$

### 3.2.4. Equations of satellite Euler rates

The motion equation of the vehicle or satellite obtained from the total angular (3.48) momentum [60].

$$
\begin{equation*}
\mathrm{H}_{\text {total }}=\mathrm{R}_{\max } * \mathrm{H}_{\text {satellite }} \tag{3.48}
\end{equation*}
$$

$\mathrm{H}_{\text {satellite }}$ is the angular momentum for the satellite body frame; $\mathrm{R}_{\max }$ is the transformation from one frame to another frame (Satellite frame to an inertial frame of reference) [60]

The torque (3.49) on the satellite, $\mathrm{T}_{\text {Satellite }}=\dot{H}_{\text {Satellite }}+\omega x$ Hsatellite

For, Rigid body $\mathrm{H}_{\text {satellite }}=\mathrm{I} * \omega$


Figure (3.13) Angular velocity and Rates in body frame [60]

In Figure (3.13) illustrate the motion of the satellite under the influence of gravitational force. The variables are defined as the position vector $\mathbf{r}$, Velocity vector $\mathbf{v}$, angular rate $\boldsymbol{\omega}$. The satellite velocity changes due to the angular velocity $\mathrm{V}_{\theta}$ (kinetic motion of the vehicle). The angular momentum of the satellite calculates in the satellite body with respect to the inertial reference frame [62]. The orbit frame (Local Vertical Local Horizontal frame-LVLH), satellite body frame, inertial frame one more frame is referred as earth frame. The body frame is fixed with satellite and principle moment of inertia. The proposed analysis considers the orbit frame (3.51) and earth frame (3.52) angular velocity in the inertial frame [63]

$$
\begin{align*}
& \omega\left(\text { (orbit-Inertial) }=\mathrm{R}_{\text {mat(Orbit-Body) }} \cdot \Omega_{\mathrm{o}}(\text { Orbit })\right.  \tag{3.51}\\
& \omega_{(\text {Earth-Inertial })}=\mathrm{R}_{\text {mat(Earth }- \text { Body })} \cdot \Omega_{\mathrm{e}}(\text { Earth }) \tag{3.52}
\end{align*}
$$

In this thesis considered the attitude sensor is (INS/GPS and IMU) used to measure the angular rates (Pitch $\boldsymbol{\theta}$, Yaw $\boldsymbol{\psi}$, and Roll $\boldsymbol{\phi}$ ) information in satellite frame. Then, it converts to angular rates of the vehicle body and the orbit frame by using Direct cosine matrix (DCM) the rotation 3-2-1 as equation (3.53) given below [62]
$R_{\text {mat }_{(\text {orbit to body })}}=$

$$
\left[\begin{array}{ccc}
c \theta c \psi & c \theta s \psi & -s \theta  \tag{3.53}\\
-c \phi s \psi+s \phi s \theta c \psi & c \phi c \psi+s \phi s \theta s \psi & s \phi c \theta \\
s \phi s \psi+c \phi s \theta c \psi & -s \phi c \psi+c \phi s \theta s \psi & c \theta c \phi
\end{array}\right]
$$

$\mathrm{C}=\cos$ and $\mathrm{S}=\sin$ terms,

To find the angular velocity (3.54) in the body to orbital frame [62]
$\omega_{\text {(Body-Orbital) }}=(\dot{\phi}-\psi S \theta) \overrightarrow{S_{b_{1}}}+(\dot{\theta} \mathrm{C} \phi+\dot{\psi} \mathrm{S} \phi \cos \theta) \overrightarrow{S_{b_{2}}}+(\dot{\psi} \mathrm{C} \theta \mathrm{C} \psi-\dot{\theta} \mathrm{C} \phi) \overrightarrow{S_{b_{3}}}$

Unit vectors are $=\overrightarrow{S_{b_{1}}}, \overrightarrow{S_{b_{2}}}, \overrightarrow{S_{b_{3}}}$, From equation (3.54) find the Euler rates as [63]

$$
\begin{gather*}
\dot{\phi}=\omega_{1}+\omega_{2} \tan \theta \sin \phi+\omega_{3} \tan \theta \cos \phi, \\
\dot{\theta}=\omega_{2} \cos \phi-\omega_{3} \sin \phi \\
\dot{\psi}=\frac{\omega_{2} \sin \phi+\omega_{3} \cos \phi}{\cos \theta} . \tag{3.54}
\end{gather*}
$$

In this section (3.2.3) for integrating the rate equation found the Euler angles. This angle represents a satellite reference to orbit reference. The magnetometer is used to detect the attitude of satellite. This signal compare with reference orbit trajectory produces the errors to the PD controllers. The actuator generates the control voltage to trigger the dynamic of satellite at low earth orbit NANO Satellite. The magnetic torques produces the control torque to the satellite [64]. Also correct the attitude errors in the body coordinates. The major role of attitude determination and control system is to bring from satellite perturbed path into an actual orbit path.

## CHAPTER 4

## MATHEMATICAL MODELING OF SATELLITE DYNAMICS

The proper attitude determination and control (ADCS) system used to stabilize the satellite into pre-determined attitude. The design requirements included this thesis, types of orbit, perturbation forces, types of satellite, and types of space environments. The satellite attitude sensors provide the rate information such as Rate Gyro's, Sun sensor, Star Sensor, Magnetometer [65], [66]. The control is delivered by suitable actuators Momentum Wheels or Magnetic Torquers. The Attitude control algorithm based upon differences between the original attitude signal and feedback signal measured from attitude sensor. The changes in attitude due to the perturbation force it generates the errors in the actuator for maintaining the desired attitude for design the suitable controllers to actuate the actuator to the required attitudes [67], [68].

$$
\begin{aligned}
& \text { Angular Momentum = Spacecraft Moment of Inertia * Angular Velocity } \\
& \mathrm{H}=\mathrm{I}_{\mathrm{sc}} \grave{\omega}_{(/ \mathrm{IB})} \\
& \mathrm{I}_{\mathrm{sc}}= {\left[\mathrm{I}_{\mathrm{xx}} \mathrm{I}_{\mathrm{yy}} \mathrm{I}_{\mathrm{zz}}\right]=\text { Spacecraft Moment of Inertia } }
\end{aligned}
$$

### 4.1. Dynamics of the Satellite

The dynamics of satellite derived from angular momentum (4.1) equation. Angular velocity is considered for the simulation, Inertial Reference Frame must be expressed in Body Frame [69].

$$
\begin{array}{r}
\mathrm{H}=\left[\begin{array}{ccc}
\mathrm{Ixx} & 0 & 0 \\
0 & \text { Iyy } & 0 \\
0 & 0 & \text { Izz }
\end{array}\right] \grave{\omega}_{(/ \mathrm{IB})} \\
\grave{\omega}_{(/ \mathrm{B})}=\left(\dot{\phi}-\Omega_{0} \psi\right) \widehat{b_{1}}+\left(\dot{\theta}-\Omega_{0}\right) \widehat{b_{2}}+\left(\dot{\psi}-\Omega_{0} \phi\right) \widehat{b_{3}} \tag{4.3}
\end{array}
$$

The angular momentum in the satellite having two parts, one is Angular momentum in the satellite body $\left(\mathrm{H}_{\mathrm{s}}\right)$, and another angular momentum in the momentum wheel $\left(\mathrm{H}_{\mathrm{w}}\right)$.

$$
\mathrm{H}=H_{\text {satellite body }}+H_{\text {Momentum Wheel }}
$$

It rotates Vehicle/Satellite body with respect to the center of mass. The rate of change of angular momentum is called as external moments [70]

$$
\begin{equation*}
\mathrm{M}=\left(\frac{d H}{d t}\right)_{\text {Inertial }}=\left(\frac{d H}{d t}\right)_{\text {Body }}+\grave{\omega}_{(I / \mathrm{B})} \mathrm{H} \tag{4.4}
\end{equation*}
$$

To design the attitude determination and control system (ADCS) considered the external moments, including the perturbation forces such as Aerodynamic force, Solar force, Gravitational attraction of the body.

Euler angle and angular rates measured from the torque equation, the attitude dynamics of the satellite equation (4.5), (4.6), (4.7) as given below (pitch, roll, and yaw) [70]

$$
\begin{align*}
& \frac{\phi(s)}{T_{x}(s)}=\frac{\frac{1}{I_{x}}}{S^{2}+\frac{K_{v x}}{I_{x}} S+\frac{4 \Omega^{2}\left(I_{y}-I_{x}\right)-\Omega h_{y}+k_{x}}{I_{x}}}  \tag{4.5}\\
& \frac{\theta(s)}{T_{y}(s)}=\frac{\frac{1}{I_{y}}}{S^{2}+\frac{K_{v y}}{I_{y}} S+\frac{3 \Omega^{2}\left(I_{x}-I_{z}\right)+k_{y}}{I_{y}}}  \tag{4.6}\\
& \frac{\psi(s)}{T_{Z}(s)}=\frac{\frac{1}{I_{z}}}{S^{2}+\frac{K_{v z}}{I_{z}} S+\frac{\Omega^{2}\left(-I_{x}+I_{y}\right)-\Omega h_{y}+k_{z}}{I_{z}}} \tag{4.7}
\end{align*}
$$

Orbital angular velocity, $\Omega$ is constant. The denominator equation of the second order transfer function is denoted as characteristic equation (4.8) is given by

$$
\begin{equation*}
S^{2}+2 \delta \omega_{n}+\omega_{n}^{2} \tag{4.8}
\end{equation*}
$$

$\boldsymbol{\omega}_{\boldsymbol{n}}$ is Undamped natural frequency, $\boldsymbol{\delta}$ is damping ratio. In the satellite system the type of damping or oscillation desired by natural frequency and damping ratio [71].

Find the pointing accuracy of satellite 0.1 degree from the final value theorem vehicle steady states in each axis [72]

$$
\begin{equation*}
\mathrm{f}(\infty)=\lim _{t \rightarrow \infty} f(t)=\lim _{s \rightarrow 0} S F(s) \tag{4.9}
\end{equation*}
$$

This is helping to determine the satellite dynamic motion and attitude in satellite frame. Attitude sensor detects the errors in the body coordinates. Euler equation of the satellite is given by [70]
$T_{x}=I_{x} \ddot{\phi}+4 \Omega^{2}\left(I_{y}-I_{z}\right) \phi-\Omega h_{y} \phi-\Omega h_{z}-\Omega\left(-I_{x}+I_{y}-I_{z}\right) \dot{\psi}-h_{y} \dot{\psi}-h_{z} \dot{\theta}-$ $I_{x} \Omega \dot{\psi}+\dot{h}_{x}$
$T_{y}=I_{y} \ddot{\theta}+3 \Omega^{2}\left(I_{x}-I_{z}\right) \theta+h_{x} \dot{\psi}+\Omega h_{z} \psi+\Omega h_{x} \phi-h_{z} \dot{\phi}-I_{y} \dot{\Omega}+\dot{h}_{y}$
$T_{z}=I_{z} \ddot{\psi}+\Omega^{2}\left(-I_{x}+I_{y}\right) \psi-\Omega h_{y} \psi+\Omega h_{x}+\Omega\left(I_{x}-I_{y}+I_{z}\right) \dot{\phi}-h_{x} \dot{\theta}+h_{y} \dot{\phi}-$ $I_{z} \dot{\Omega} \phi+\dot{h}_{z}$

From the equation (4.10), (4.11), (4.12) describes the disturbance torque subjected to the satellite. The angular velocity of the satellite is $\Omega$, the angular moment of momentum wheel is $h(x, y, z)$, Moment of inertia of satellite $\left(I_{x}, I_{y}, I_{z}\right)$, Angular acceleration of body $(\ddot{\phi}, \ddot{\theta}, \ddot{\psi})$ Angular velocity $(\dot{\phi} \dot{\theta}, \dot{\psi})$. The torques $\mathrm{T}(\mathrm{x}, \mathrm{y}$, and z ) directions, whenever any perturbation forces acting on a satellite it creates the counter moments because of conservation of angular momentum [70].

### 4.1. Perturbations in Low Earth Orbit

The perturbations (or) disturbances are orbital variation in the nominal orbit. This variation is periodic nature. These perturbations affect the Keplerian elements such as angular motion of perigee with respect to ascending node, true anomaly, mean anomaly (For Circular Orbits), and right assertion to ascending node (RAAN). The linear changes in the orbit referred as secular variation. It might be small variation or large variation [73]. The variation is less than the orbital period to complete one cycle referred as linear small variation in orbital elements, and large variation the orbital period is greater than the orbital period. For, low earth orbiting (LEO) satellite
reflected as two body problems. In this thesis considered the aerodynamic drag, the gravitational attraction of the earth's surface, solar radiation. For, High earth orbits (HEO) is referred as three body problems such as moon attraction, the sun's attraction is mainly affecting the ascending node with respect to the perigee position. This perturbation occurs in the orbit because of ecliptic pole and gyroscopic precision [74]. The planets disturb the earth orbital plane such as Saturn, Jupiter, and Venus. These planets and SUN perturb the lunar orbit around the earth's surfaces. Mostly high earth orbits have little or no aerodynamic drag. HEO satellite travels with less speed compares to the low earth orbiting satellite. The gravitational attractions/pulling near to the atmospheric region having enormous amount of aerodynamic drag acting on the body. At, LEO satellite travels very high speed near to perigee and less speed at apogee [75].

The trajectories analysis Runge- Kutta methods used to compute of orbit multiple body system presented in this thesis. The orbital position calculated by using high performance computing methods in the orbit. The artificial satellite bodies whose orbit enters the atmosphere under the influence of atmospheric drag force and gravitational attractions [76]. Where it is either dis-integrate the satellite in the atmosphere due to the gravitational force acting on a body. To avoid this need to maintain the orbital plane, it requires the frequent firing from the thruster to maintain the satellite as per the required orbital plane from perturbed plane [77]. The shape of the Earth is not perfectly spherical. The homogeneity of earth will cause the perturbation in the orbit. The bulge in near equator and $\mathrm{J}_{2}$ perturbation forces will make the changes in the orbital elements. When satellite crosses to the equator, it may affect the orbital elements pull towards to the earth's surface [78].

The artificial satellite in the low earth orbit increase day by day. It makes space debris in the near earth's atmosphere. There is the possibility to colloid each satellite changes the orientation and Keplerian elements. To avoid such collation continually monitor the space debris and to keep the satellite in the same orbit requires the attitude stabilization and control mechanism [79]. The satellite position/orientation measured from star sensor, sun sensor, and earth sensor. The feedback signals from attitude sensors compare with references attitude signal based
on that dynamic of the actuation system like control momentum gyroscope, fly wheels, magnetic torque generates the control torque to the satellite system [80]. The perturbation equation is very useful to design the Attitude determination and control system (ADCS).

## Types of Perturbations

The different types of perturbations

- Atmospheric drag
- Lunar and solar gravity
- Shape of the earth
- Solar radiation


### 4.2.1 Atmospheric Drag Effects

Atmospheric drag forces vary with altitude at low earth orbit having more aerodynamic drag between 120 KM 160 KM . It reduces the satellite altitude in the earth's atmosphere. When satellite reaches near to Roche limit focal point approximately 80 KM altitudes very rapidly decay the orbital elements. For, above 1000 KM altitudes drag is less changes of satellite orbital elements. This is called as orbital decay [81]. The low earth orbit satellite to decrease the altitude due to large drag affects the satellite body, also decrease the mission lifetime. The drag forces $\mathbf{F}_{\text {drag }}$ is increased with velocity $\mathbf{V} ; \mathrm{F}_{\text {drag }} \perp \mathrm{V}^{2}$. The satellite/vehicle velocity depends upon the different altitudes, also its proportional with Area of the satellite and air density; if the surface area of the satellite increases the drag will also increase. The large amount of aerodynamic drag will strike satellite during launch and Re-Entry period. The Ballistic constant $\frac{m 2}{C_{d} A}$ depends upon the satellite mass, coefficient of drag (For flat plate it is 2), area of the satellite. Mostly low earth orbits the satellite mass is considered as constant [82].

The aerodynamic drag effect is the major force affecting the object or satellite in the low earth orbit (LEO). When the vehicle (satellite) body moving in the
atmosphere by GAS particle and liquid particle produces the resistance offered due to drag force. The drag force is more effects during launch the satellite and re-entry the space vehicle [83]. See figure (4.1) when the satellite has entered the upper atmosphere because of the gravitational attraction and aerodynamic drag effect are considered more altitude. Finally, near to earth atmosphere to dis-integrate the satellite parts and further enters because of orbital decay. The satellite attitude reduces because of these perturbation forces. This effect decreases the lifetime of the satellite.

The Cowell's perturbation equations used to design and analysis the perturbations in ideal conditions due to atmospheric drag and solar radiation Pressure in Low Earth Orbit satellite [84]. Their causes and how they affect the spacecraft in the Low Earth Orbit are explained shown in figure (4.2). The perturbation algorithms developed by MATLAB environment.


Figure (4.1) Perturbation forces in LEO Satellite [84]
The aerodynamic force of the satellite is given in the equation (4.13)

$$
\begin{equation*}
F_{\text {drag }}=\frac{1}{2} \rho v^{2} A c_{d} \tag{4.13}
\end{equation*}
$$

Let, $F_{\text {drag }}$ aerodynamic parameters of the vehicle body, $\rho$ air density, $v$ is velocity acting on a satellite, A is the Surface Area of the object or Satellite, $C_{d}$ is the drag coefficient [84].

Typically, for earth's approaching satellites having a more coefficient of drag. It is mainly depending upon the changes in the altitude [85]. The altitude above 90 km having extreme ultraviolet radiation and more temperature because of SUN effects with respect to altitude. The altitude between 200 km to 260 km temperature is reduced from 1150 K to 600 K because of the solar activity due to high density. Due to SUN effects the solar radiation satellite has maximum decay in day times and minimum solar perturbation decay during eclipse period [85].


Figure (4.2) Satellite orbit trajectory in drag region [85]

The ADCS of Hubble Space Telescope design specification is 7/1000th point accuracy of arcsecond. The proposed work the accuracy of satellite dependent on disturbance effects accumulating with time. The disturbance/Perturbation arises from internal factors or external factors. The sensor calibration and alignment error create the internal noises and due to the environmental disturbances are external factors. The internal disturbances are closely tied witch spacecraft structure, in particular: internal moving parts and mass or radiation being emitted [86].

For circular orbit the changes in (period, velocity, acceleration) a, T, v per revolution as is given equation (4.14), (4.15), (4.16), (4.17)
$\Delta_{a_{r e v}}=-\frac{2 \pi C_{d A} A a^{2}}{m_{2}}$
$\Delta P_{\text {rev }}=-\frac{6 \pi^{2} C_{d} A \rho a^{2}}{m_{2} V}$
$\Delta V_{\text {rev }}=\frac{\pi C_{d} A \rho a V}{m_{2}}$
$a_{r}=-\frac{4.5 \times 10^{-8} A}{m_{2}}$
The Ballistic coefficient ( BC ) depends upon the mass of the satellite and area of satellite it is described by $m_{2} /\left(C_{d} A\right), \mathrm{BC}$ is constant for most of the satellite. The drag effects are large when the lower valve of ballistic constant and drag is less when high value of ballistic constant [87].


Figure (4.3) Direction of aerodynamic drag force and atmospheric torque [87]
To estimate the lifetime of the satellite life cycle as is given equation

$$
\begin{equation*}
L=\frac{-H}{\Delta a_{r e v}} \tag{4.18}
\end{equation*}
$$

Hence, the equation (4.18) describes the changes in atmospheric density due to altitude variation and solar changes. H is mean by the atmospheric scale height of density [87].

### 4.2.2. Solar and Lunar Gravity Perturbation

The three-body problems such as SUN attraction and MOON attraction perturb the satellite into normal orbit to perturbed orbit due to gravitational effects. These types of perturbation due to other body except earth disturb the orbital elements [88]. In High earth orbit satellite attitude is change periodically s due to the ecliptic pole and precision of the gyroscope. The SUN and lunar attractions mainly change the argument of perigee. The vehicle moves in orbit intersecting the equatorial orbit with satellite orbit from South Pole to the North Pole [88]. The equation of perturbation given by the following

$$
\begin{aligned}
& \Omega \text { moon }=-\frac{\operatorname{Cos}(i n c)}{n} *(0.00338), \quad \Omega \operatorname{sun}=-\frac{\operatorname{Cos}(\text { inc })}{n} *(0.00154) \\
& \omega \text { moon }=\frac{\left(4-5 * \operatorname{Sin}^{2}(\text { inc })\right)}{n} *(0.00169), \omega \operatorname{sun}=\frac{\left(4-5 * \sin ^{2}(\text { inc })\right)}{n} *(0.00074)
\end{aligned}
$$

Where, the orbit inclination (inc) is i, orbital revolutions per day is n, Both Argument of perigee $\&$ longitude of ascending nodes defined as degree/day.

### 4.2.3. The Flattening or Non-Homogeneity of the Earth

The Earth is not a perfect sphere; it is some extent flattening in the top surface and bottom surface [89]. The Non-homogeneity of earth See. Figure 4.4 causes the harmonics such as $\mathrm{J}_{2}$ perturbations.


Figure (4.4) Flattening of the Earth's surface with $\mathbf{J}_{2}$ perturbations [57]

The Earth is slightly bulging on near the equator. Because of the non-homogeneity of the earth's surface at the top and bottom create more change in the high earth orbit especially Geosynchronous Earth Orbits, GEO. The low earth orbit, it is not taking any variation in the orbit after many revolutions [90]. Earth harmonics, $\mathbf{J}_{2}$ changes the satellite position with respect to the perigee. These variations are periodic nature change the orbital elements.

The equations for this perturbation are [90]

$$
\begin{aligned}
& \vec{a} \mathrm{E}_{\mathrm{x}}=\mathrm{GA}_{\mathrm{J} 2}\left(15 \frac{x z^{2}}{r_{\text {earth }}{ }^{-}}-3 \frac{x}{r_{\text {earth }}{ }^{5}}\right) \\
& \vec{a} \mathrm{E}_{\mathrm{y}}=\mathrm{GA}_{\mathrm{J} 2}\left(15 \frac{y z^{2}}{r_{\text {earth }}{ }^{7}}-3 \frac{y}{r_{\text {earth } h^{5}}}\right) \\
& \vec{a} \mathrm{E}_{2}=\mathrm{GA}_{\mathrm{J} 2}\left(15 \frac{z^{3}}{r_{\text {earth }}{ }^{7}}-3 \frac{z}{r_{\text {earth }}{ }^{5}}\right)
\end{aligned}
$$

## Shape of the Earth

When designing the Low Earth Orbits satellite, assuming the earth is purely spherical or symmetrical mass. But most variation in the satellite in the bulge nears the equator. These perturbations are referred as Zonal perturbation depends upon the Geo-potential coefficients and zonal coefficients [91]. The flattening of the earth pole earth potential function, acceleration of the satellite body found from the gradient of potential function depends upon the latitude [91]. The potential function changes the orbital elements because of the non-homogeneity of the Earth's surface from the equation. The variations in $\Omega_{\text {RaAN }} \& \omega_{J 2}$ due to periodic/secular variation of Earth's Oblateness ( $\mathbf{J}_{2}$ ) equations (4.19), (4.20) as follows

$$
\begin{align*}
& \Omega_{J 2}=\frac{-1.15 n J_{2}\left(\frac{r_{\text {earth }}}{\mathrm{a}}\right)^{2} \cos *(i)}{\left(1-e c c^{2}\right)^{2}}  \tag{4.19}\\
& \omega_{J 2}=\frac{-1.15 n J_{2}\left(\frac{\left.r_{\text {earth }}\right)^{2} \cos * i}{2}\right.}{\left(1-e c c^{2}\right)^{2}} \tag{4.20}
\end{align*}
$$

Where, n is the number of revolutions/days
$\mathrm{J}_{2}$ - Zonal Coefficient
$r_{\text {earth }}$ - The earth radius
a -Semi Major axis
i -Inclination of orbital Plane
ecc - eccentricity of orbit

### 4.2.4. Solar Radiation

When design the LEO Nano satellite attitude control system, the major perturbation is solar perturbation periodic variation in the orbital components [92]. The changes in attitudes of LEO satellite due to solar radiation pressure.

The solar radiation pressure force is given in the equation (4.21):

$$
\begin{equation*}
p_{\mathrm{sr}}=\frac{S F}{c}=\frac{1353}{3 \times 10^{8}} \frac{\mathrm{~W} / \mathrm{m}^{2}}{\mathrm{~m} / \mathrm{s}}=4.51 \times 10^{-6} \mathrm{~N} / \mathrm{m}^{2} \tag{4.21}
\end{equation*}
$$

Where,

$$
\begin{aligned}
& \text { Solar Flux }(\mathrm{SF})=1353 \mathrm{~W} / \mathrm{m}^{2} \text { (Radiation constant) } \\
& \text { Velocity of light }(\mathrm{c})=3 \times 10^{8} \mathrm{~m} / \mathrm{s}
\end{aligned}
$$

The solar force is directly proportional to the SF and inversely proportional to the speed of the light. The disturbances analysis at LEO consider for the simulations torque generated (See Equation 4.22) due to solar radiation is given by: [92]

$$
\begin{equation*}
\tau_{\text {Solar-radiation }}=-p_{\mathrm{sr}} \times \mathrm{C}_{\mathrm{r}} \times \mathrm{A} \Theta \times\left(\mathrm{C}_{\mathrm{psr}}-\mathrm{C}_{\mathrm{g}}\right) \tag{4.22}
\end{equation*}
$$

Where,

Visible area of the SUN $\mathrm{A}_{\Theta}$

Reflectivity $\mathrm{c}_{\mathrm{r}}$
Center of pressure $\mathrm{c}_{\mathrm{psr}}$

Center of gravity $\mathrm{c}_{\mathrm{g}}$

$$
\begin{equation*}
a_{r}=-\frac{4.5 \times 10^{-8} A}{m 2} \tag{4.23}
\end{equation*}
$$

The cross-section surface area of vehicle/satellite is defined by ' A ' which observable to the SUN. The total weight (mass) of the satellite is ' m ' expressed in kg .


Figure (4.5) Change in satellite cross sectional area due to Solar Radiation [87]
An altitude below 800 km having more aerodynamic drag acceleration and less solar radiation force or pressure acting on a satellite [93]. An altitude above 800 km having solar radiation pressure (See Figure 4.5) is greater than the aerodynamic drag force. In this region, the predominant effects of perturbation due to the solar activity. The acceleration due to atmosphere drags is $\mathrm{a}_{\mathrm{ad}}$ and the solar radiation pressure is $\mathrm{a}_{\mathrm{r}}$; An altitude below 800 KM distance $\mathrm{a}_{\mathrm{ad}}>\mathrm{a}_{\mathrm{r}}$, An altitude above $800 \mathrm{KM} \mathrm{a}_{\mathrm{r}}>\mathrm{a}_{\mathrm{ad}}$. It's clearly understood when designing the LEO satellite having more acceleration from the atmosphere drag and less acceleration from solar radiation force. The Keplerian elements change due to solar radiation is the most important factor to perturb the orbit from actual path [92].

### 4.3. Periodic and Secular Variation

Periodic and secular variation in the Keplerian elements due to atmosphere molecules changes the satellite orientation [73]. These generate the decay function to decrease the orbit life time. In the lower altitudes having dense atmosphere and more aerodynamic drags \& more heat disintegrates the satellite, to burn the satellite during the re-entry.


Figure (4.6) various sources of orbital perturbation [73]
The preceding example illustrated the effect of periodic variation. In Figure 4.6 shown the deceleration of vehicles varies with altitude under the influence of various sources of perturbation forces.


Figure (4.7) Secular variation vs. short \& long period variation [73]

There are three types of disturbances from Figure 4.7. [73]

- Short Periodic - Cycles every orbital period.
- Long Periodic - Cycles last longer than one orbital period.
- Secular - Does not cycle. Disturbances mount over time.

Table (4.1) Keplerian Elements vs. Orbital Periods [73]

| Orbital Elements | Perturbations with <br> time | Perturbation > <br> Orbital period <br> (Long cycle) | Perturbation < <br> Orbital period <br> (Short cycle) |
| :--- | :--- | :--- | :--- |
| Semi major axis | $X$ | $X$ | $\uparrow$ |
| Eccentricity | $X$ | $\uparrow$ | $\uparrow$ |
| Inclination | $X$ | $\uparrow$ | $\uparrow$ |
| RA of Ascending <br> Node | $\uparrow$ | $\uparrow$ | $\uparrow$ |
| Argument of <br> perigee | $\uparrow$ | $\uparrow$ | $\uparrow$ |
| Mean anomaly | $\uparrow$ | $\uparrow$ | $\uparrow$ |

$\uparrow$ Means YES, X Means NO
In table (4.1) summarizes the perturbation in the orbit short period and long period with Keplerian elements

### 4.3.1 Decay of Eccentricity

The drag occurs at perigee is more than apogee. See Figure 4.8 shows the decay the orbit eccentricity from apogee and perigee with time in days from Epoch. This indicates the lifetime of the satellite decreases with time [94].


Figure (4.8) Variations in the altitudes between perigee/apogee since epoch [94]

### 4.3.2. Drag Effects on Eccentric Orbits

When the satellite moves at perigee have constant eccentricity and density with same altitude is illustrate in Figure. 4.9. The altitude of apogee having variation in size of the orbit eccentricity, density with changes in altitude [95].


Figure (4.9) Variation in eccentricity, density with different altitudes [95]
ISS Shows the Figure 4.10. Orbital decay with altitude vs years [96].


Figure (4.10) Orbit Decay of the International Space Station (ISS) [96]

### 4.3.3. Spacecraft Lifetime Solar Activity Effect

In Figure (4.11), (a) and (b) show the mean value of a lifetime ( $1 \mathrm{~b} / \mathrm{ft}^{2}$ ) and $\left(\mathrm{m}^{2} / \mathrm{kg}\right)$. Actual values will depend on ballistic constant [97].


Figure (4.11) Spacecraft Lifetime Solar Activity (a) Altitudes vs. Lifetime ( $\mathbf{l b / f t} \mathbf{t}^{2}$ ) [97]

(b) Graph between Altitudes vs. Lifetime ( $\mathbf{m}^{2} / \mathbf{k g}$ ) [97]

### 4.4. Perturbation Formulations for Numerical Solutions

The perturbations prediction and analyze with the help of numerical equation to determine the orbit position. There are two types of numerical integration used to analyze the perturbation model one is Cowell formation another one Encke's formation. In this thesis consider for Cowells algorithm used to design \& simulate the orbit with Keplerian elements. The Cowell methods mainly used to find the changes in acceleration with maintain magnitude of primary acceleration [36]. The Encke's method is very much useful for design high earth orbit (HEO) satellite and interplants operation. These methods are used to predicting the satellite position of in the future (or) next state. The design simulation of orbit is analyzed by a differential equation to identify the disturbed orbit from normal orbit. The differential ordinary integration methods widely used to find the motion of satellite due to perturbations forces [36].

The differential equations used to find the motion equations of the satellite with perturbation in the orbit. At LEO satellite revolves the earth's surface in the form of circular motion [84]. The motion equations are practically $2^{\text {nd }}$ order ordinary differential equations (ODE). These formulated equations are used to predict a satellite's upcoming position, velocity, and Keplerian elements. The proposed perturbation design analysis in International space station (ISS), SRM satellite, Pratham satellite implemented with Runge-Kutta numerical methods [98]. This method using ODE Equations to find perturbed trajectory by the knowledge of initial condition and time interval get the actual trajectory. Also, another method Euler method used to predict the oscillating orbit. The Modified Euler methods find the disturbed path from an initial condition of pre-determined path of the orbit. The simulation methods are developed by MATLAB or Global Mission Analysis Tool (GMAT), out of which the results are more accurate, even for larger values of step size (h). These two methods that can be employed in the thesis, formulation of differential equations termed as Cowell's Method at LEO orbits and Encke's Method at HEO orbits [99].

The Cowell's method involves the direct integration of the full equation of motion in rectangular coordinates. Therefore, this method is also known as a direct numerical integration method. This method is mostly used if the values of
perturbation are comparable to or greater than the primary body gravity acceleration [84]. The Encke's method involves the changes in the satellite normal orbit and disturbed/Perturbed orbit. The accelerations different between both orbits are integrated and adding in to initial orbit and perturbed orbit. It is, finally results in perturbed satellite state vectors. This method is a more efficient for long interplanetary mission spacecraft, because of the built-in rectification procedure.

The method which can be used for the formulation of a differential equation is a Cowell's method, because it is much more accurate for the Low Earth Orbit (LEO) satellite when compared to Encke's method [36]. Cowell's method is also less complex, i.e., easy to formulate and program. This research work included is aerodynamic drag, solar radiation pressure after the formulation of a differential equation by Cowell's method with orbital perturbation for the simulation

### 4.4.1. Cowell's construction method

The Cowell's constructions used to find the equations of motion of the satellite and integrating the rectangular coordinate. Illustrate in Figure 4.12, body ' i ' moving into the body 1 because of gravitational attraction perturbed orbit to osculate orbit. This method is broadly used to integrate the perturbation acceleration equal to zero [84].

The Cowell's construction equation (4.24) is given below

$$
\begin{equation*}
\ddot{r}=-\frac{\mu}{r^{3}} r+a_{\text {perturbed }} \tag{4.24}
\end{equation*}
$$

The Position equation (4.25) and velocity equation (4.26) given by

$$
\begin{equation*}
x=\binom{r}{\frac{d r}{d t}} \tag{4.25}
\end{equation*}
$$

And

$$
\begin{equation*}
\frac{d x}{d t}=\binom{\frac{d r}{d t}}{-\frac{\mu}{r^{3}} r+a_{p}} \tag{4.26}
\end{equation*}
$$



Figure (4.12) Cowell's Method from perturbed orbit to osculate orbit
A Cowell construction is expanded by follows:

1. Set Initial condition, position $r=r_{0}$; velocity $\mathrm{v}=\mathrm{v}_{0} ;$ time $\mathrm{t}=\mathrm{t}_{0}$.
2. Govern changes in time $\Delta t / t_{f}$ by n steps
3. Integrate the vector from zero to n steps
a. To Integrate $\frac{d r}{d t} \& \frac{d r}{d t}$ To find position vector and velocity vector.
b. Add time, $\mathrm{t}=\mathrm{t}$ and $\Delta t$.
4. Calculated final position, velocity, time

As we know Aerodynamic Force given as

$$
\begin{equation*}
F_{A D}=\frac{1}{2} \rho v^{2} C_{d} A \tag{4.27}
\end{equation*}
$$

Where, $\rho=$ free-stream density
$v=$ free-stream velocity
$C_{d}=$ Coefficient of Drag $=2.0$
$A=$ Area of the satellite body

Aerodynamic acceleration (4.28) or total perturbing acceleration can be calculated as below

$$
\begin{equation*}
a_{P}=a_{A D}=F_{A D} / m_{2} \tag{4.28}
\end{equation*}
$$

Where, $m_{2}=$ Mass of the satellite body

$$
\begin{gather*}
\dot{r}=v  \tag{4.29}\\
\dot{v}=-\frac{\mu}{r^{3}} r+a_{A D} \tag{4.30}
\end{gather*}
$$

Hence, the equations (4.29) \& (4.30) satellite velocity \& satellite acceleration can be solved with the help of the Runge- kutta method [84].

### 4.2.2. Encke's construction method

The Encke's construction method used to find the perturbed orbit position, velocity from osculating orbit (2-body) by the difference in the acceleration in both the orbit [36].


Figure (4.13) Finding the satellite position and velocity of osculating orbit

To integrating this difference and summation of the result gives the position and velocity of osculating orbits (See Figure 4.13) continues till maximum amount of magnitude of position difference some tolerance limit. Osculating orbit position adjusted from the current disturbance state [36].

### 4.5. Attitude controllers for NANO satellite in low earth orbits

The Magnetic Torquers or Magnetic actuator is used to generate the magnetic flux which interacts with the Earth magnetic flux (GEO Magnetic field). This produces control torque or control forces to satellite. For three axis control, (CMG) Control momentum gyros are used for stabilizing the satellite [100], [101]. The magnetic thruster or armature control DC motor used in in low earth orbits satellite. In this thesis included the dynamics of armature control DC motor with Nano satellite inertia is considered. The Earth's magnetic fields change the satellites attitudes from pre-determined attitude. An actuator used to correct the actual path of satellite orbit even with perturbation. The magnetometer is used to measure the Earth's field (magnetic). The momentum wheel (or) Flywheels used to control the periodic disturbances in satellite [102]. The magnetic control system consists of solenoid coil and Geo-Magnetic field. The NANO satellite weights (mass) considered between 1 kg to 10 kg . It carries smaller payloads for earth observation or navigation application. The NANO satellite either controlled by active control or passive control techniques. An active control method using Armature controlled DC motor to correct the attitudes of NANO satellites in LEO orbiting satellite [103], [104]. The Control momentum Gyroscope (CMG), Momentum Wheel or Flywheels is used to control the High earth orbiting satellite. The CMG consist of small motor attached with wheels on satellite body [105].

When ADCS design requires the torques on a Di-pole (2-pole magnets) with uniform or constant magnetic field. This field is required for keeping the dipole perpendicular to the constant magnetic field. The Dipole moment is a product of the strength of the pole and it is separated by two plates. It has North Pole (N) and South Pole (S) where both force magnitude is same but different direction. The Magnetic Dipole Moment (MDM) is expressed in NM/T [106], [107]

$$
\begin{equation*}
\text { Torque }(\mathbf{T})=\mathbf{M} * \mathbf{B} \tag{4.31}
\end{equation*}
$$

The Magnetic fields (B) and dipole (M) of the torque (4.31). For air armature (or) conductor, the dipole is

$$
\begin{equation*}
\mathbf{M}=\mathbf{N} * \mathbf{I} * \mathbf{A} \tag{4.32}
\end{equation*}
$$

Where (4.32) Coil (or) No of turns in the winding is N , the current in the coil is I, and, the area of the conductor is A, Characteristically, the coil is wounded from place to place of the satellite structure or body [108].

The magnetic fields can be taken as a periodic function for all times, the dynamics of satellite/spacecraft control mechanism and stabilization is achieved by suitable actuator or controllers [109]. By the application of the above-mentioned theory on a satellite actuation system like the magnetic torque control the complete orbit in a fixed dipole and GEO magnetic field. In Figure (4.14) illustrate how to generate the torques or force from magnetic actuator by the help of Dipole and Geo magnetic (earth) fields [110].


Fig (4.14) Magnetic Actuator [110]

The magnetic torquers are generating the control torque, $\boldsymbol{T}$ control for re-orienting the satellite to prescribed orbits. In the Figure 4.15 shows the magnetic moments, required control torque, $\boldsymbol{T}_{\text {Required }}$ and magnetic-flux vector, $\mathbf{B}$ in the satellite frame from the magnetometer measurements. The control torques requirement should be regular to the magnetic flux, we have $\boldsymbol{B T}$ control $=0$ [110]. The most efficient magnetic moment is the one that always points normal to the $\boldsymbol{B}$ vector, so that $\boldsymbol{M T}$ control = zero. The control torques can be calculated from Magnetic field (B) and Regulated torque ( $\mathbf{T}_{\text {Required }}$ ). The magnetic torquer generates the control displacement in the satellite, the torquers generates $0.05 \mathrm{~N}-\mathrm{M}$ at low earth orbit altitude of 1000 KM [111], [112].


Figure (4.15) Regulator Torque $\mathrm{T}_{\mathrm{R}}$ and Magnetic Di-Pole Moment [113]
The respective unit vectors are: $\left(\mathrm{e}_{\mathrm{x}}, \mathrm{e}_{\mathrm{y}}, \mathrm{e}_{\mathrm{z}}\right)=\left(\right.$ Tcontrol $\left./ \mathrm{T}_{\text {control, }, \mathrm{M}} / \mathrm{M}, \mathrm{B} / \mathrm{B}\right)$
Hence, the Control torque (4.33) vector is: (Tc) [113]
Tcontrol $=\left(T_{\text {Required }} * s * \beta\right) e_{x}=T_{\text {Required }}-\left(T_{\text {Required }} * c * \beta\right) e_{z}=T_{\text {Required }}-\left(T_{\text {Required }} *\right.$ $c^{*} \beta$ ) $\mathbf{B} / \mathbf{B}$
$\mathrm{s}=\sin$ and $\mathrm{c}=\cos$ terms
$\mathbf{T}_{\text {control }}=$ Moment (M) $*$ Magnetic field (B)
The Magnetic Diploe Moment (4.34), MDM is the moment of magnetic dipole or solenoid moment; the changes in ' $\mathbf{M}$ ' directly affect the control torques and indirectly affect the magnetic field [114], [115]. The pole strength and distance between two magnetic poles are well-defined; M is inversely proportional to the Field vector (Magnetic) B and required torque vectors $\mathbf{T}_{\text {Required }}$

$$
\begin{align*}
& \mathbf{M} \boldsymbol{\alpha} \text { Magnetic field } * \mathrm{~T}_{\text {Required }} \\
& \mathbf{M}=\frac{\text { Magnetic field } * \text { TRequired }}{B^{2}} \tag{4.35}
\end{align*}
$$

The Nadir is defining the line of observation from satellite looking at string line to the earth's surface; Satellite and earth have the same line of sight [116]. The Zenith is exactly opposite to the Nadir and the satellite moment around Nadir/Zenith direction is expressed by Yaw angle and velocity direction is expressed by Roll angle, orbit direction (or) movement is expressed by Pitch angle [117].

## Solenoid coils

Solenoid is bar magnets and current caring conductor/coil creates the Di-Pole moments. It is an electromagnetic armature used in satellite system. The interaction between the earth magnetic flux vs. Magnetic torquer derives this section [118]

Motion equations (EOM) for twin coil BAR magnet and coil
In free space, a Magnetic field plays the major roles in the permeability. And if we consider closed surface area or nozzle area, permeability of median it is assumed as infinity [118]. Therefore, the magnetic fields are considered zero ( $=0$ )

$$
\oint_{C}^{1} \text { Magnetic field } . d l=\int_{s}^{1} J n d a
$$

Hence, the above equation states, to integrate the magnetic field in the closed path in the surface area $\mathbf{S}$. Surface area in the magnetic field is mostly same of the integration of electric current, $\mathbf{I}$ on the surface path.

### 4.5.1. IGRF: International Geomagnetic Reference Field

IGRF model is the typical numerical model of GEO magnetic field under changes in the orbit due to perturbation variation [119]. The data are collected from the "International Association of Geomagnetism and Aeronomy" (IAGA). This is an international association to study and observation of magnetism in planetary and terrestrial field. The scalar potential model (4.36) of earth's magnetic field consists of Gauss coefficient and Spherical harmonics as given below [120]

$$
\begin{equation*}
V(r, \phi, \theta, t)=a \sum_{\ell=1}^{L} \sum_{m=0}^{\ell}\left(\frac{a}{r}\right)^{\ell+1}\left(s_{\ell}^{m}(t) \cos m \phi+h_{\ell}^{m}(t) \sin m \phi\right) P_{\ell}^{m}(\cos \theta) \tag{4.36}
\end{equation*}
$$

$r=$ Radial distance from the earth center
$\phi=$ East longitude
$\theta=$ latitude (polar angle)
$\mathrm{a}=$ Radius of the earth's surface ( $6,371 \mathrm{~km}$ or 3,959 miles)
$g_{l}^{m} \& h_{l}^{m}=$ Gauss coefficients
$p_{l}^{m}=$ Normalized Legendre function (canonical form) of (1 is the degree; $m$ is the order)
$\mathrm{L}=$ Expansion degree (Max)

The Gauss coefficients may vary linearly with respect to times. IGRF models provide the accurate measurements in the earth magnetic field over a period of times [120].

### 4.5.2. Attitude control with control momentum gyroscope

The Attitude Control system (ACS) of satellite with control momentum gyroscope attached with a stable platform. It is called as Gimbal (single gimbal or dual gimbals) [121]. The design of the ACS with momentum wheel is shown in Figure (4.16)


Fig (4.16) Two-axis Attitudes control of satellite with momentum wheels [121]

It is very important to analysis the less weight and low-cost sensors models. The design simulation of ADCS requires the single gimbal variable speed control momentum gyroscopes (SGVSCMG). [122].


Fig (4.17) Satellite positions and orientations with Gyroscope frame [122]

The Figure (4.17) describes the inertial frame $\left(\mathrm{X}_{\mathrm{I}}, \mathrm{Y}_{\mathrm{I}}, \mathrm{Z}_{\mathrm{I}}\right)$ rotations about body frame $\left(X_{B} / B_{1}, Y_{B} / B_{2}, Z_{B} / B_{3}\right)$ about the roll axis. The reaction wheels/momentum wheel is attached to satellite body. The Gyroscopes consists of inner gimbals and outer gimbals to calculate angular motions in spin axis. The Gimbal is attached to the rotor or small motor to the gyroscope frame [123]. The Unit vectors are ( $\mathbf{u v g}^{1}$; $\mathbf{u v g}^{2}$; $\mathbf{u v g}^{2}$ ) respectively. The quaternion parameter represents the attitude orientation of the satellite. These parameters consist of one real quantity and three imaginary quantities used to find the vectors in one frame to another frame vectors. The Quaternion parameters (4.37) relates with the angular frequency defined by [124]

$$
\begin{align*}
& \text { quaternion } \left.=\llbracket \frac{q 0}{q v} \rrbracket=\llbracket \frac{\frac{q 0}{q x}}{\frac{q y}{q z}}\right\rfloor  \tag{4.37}\\
& \left.\llbracket \frac{\frac{q_{0}}{q_{1}}}{\frac{q^{2}}{q_{3}}}\right\rceil=0.5 * \text { [quaternion parameters] } * \text { [angular velocity] } \tag{4.38}
\end{align*}
$$

In the section discuss the quaternion method is described angular momentum of the satellite. Where $q 0, q \mathrm{x}, q \mathrm{y}$, and $q \mathrm{z}$ are the parameter quaternion, and angular velocity $=\left[\omega_{\text {satellite }}^{1}, \omega_{\text {satellite }}^{2}, \omega_{\text {satellite }}^{3}\right]^{T}$ are angular vector (4.38) of satellite about the satellite body reference coordinates. [124]

$$
\begin{gathered}
C v_{\text {Body }}^{\text {Inertial }}=\llbracket C v_{j k} \rrbracket \mathrm{j}, \mathrm{k}=1,2,3 \quad C_{\text {Body }}^{\text {Inertial }}=C_{B}^{I} \\
C v \mathrm{jk}=\cos * \beta \mathrm{jk} \\
C v_{\text {Body }}^{\text {Inertial }}= \\
C v_{B}^{I}=\left[0 \quad \omega_{\text {satellite }}^{3} \omega_{\text {satellite }}^{2} ; \omega_{\text {satellite }}^{3} 0-\omega_{\text {satellite }}^{1} ;-\omega_{\text {satellite }}^{2} \omega_{\text {satellite }}^{1} 0\right]
\end{gathered}
$$

Direct cosine matrix included in quaternion, because it is an easy way to describe the satellite axis into the inertial reference axis. The DCM is described the control of $Z_{B}$ axis direction by $C v_{\text {Body }}^{\text {Inertial }}$. Hence the equation (4.39) given by

$$
\left[\begin{array}{l}
C v 13  \tag{4.39}\\
C v 23 \\
C \dot{v} 33
\end{array}\right]=\left[\begin{array}{lll}
C v 11 & C v 12 & C v 13 \\
C v 21 & C v 22 & C v 23 \\
C v 31 & C v 32 & C v 33
\end{array}\right]\left[\begin{array}{c}
\omega 2 \\
-\omega 1 \\
0
\end{array}\right]=\left[\begin{array}{lll}
-C v 12 & C v 11 \\
-C v 22 & C v 21 \\
-C v 32 & C v 31
\end{array}\right]\left[\begin{array}{l}
\omega_{\text {satellite }}^{1} \\
\omega_{\text {satellite }}^{2}
\end{array}\right]
$$

In the design analysis considered the net angular momentum (4.40) of the vehicle/satellite is not equal to zero. It is constant with respect to a fixed reference frame [125]
$\mathrm{h}_{\text {momentum (h) }}=\mathbf{J} \boldsymbol{\omega}+\boldsymbol{J}_{\boldsymbol{G}(\text { Gimbal })} \dot{\boldsymbol{\gamma}} \boldsymbol{g}_{\mathbf{2}}+\boldsymbol{J}_{\boldsymbol{W}(\text { Wheel })} \Omega \boldsymbol{g}_{\mathbf{3}} \neq 0$

Where, $J=$ diag [J1; $J 2 ; J 3$ ] the satellite inertia matrix with gyroscope in the inner gimbal and wheel attached to it. The Wheel moment of inertia, Jwheel is defecting about $Z_{G}$ axis and Gimbal moment of inertia, $\mathbf{J G i m b a l}$ is rotation about $\mathrm{Y}_{\mathrm{G}}$. The revolution of the Wheel is $\Omega$
M.O.I is moment of inertia:

The equation (4.41) of satellite as given below [110]
Angular momentum, (h) $=\left[\begin{array}{c}\text { hmoment } 1 \\ \text { hmoment } 2 \\ \text { hmoment } 3\end{array}\right]=\left[\begin{array}{cc}J 1 \boldsymbol{\omega 1}+J_{\boldsymbol{W}(\text { Wheel })} & \Omega \sin \gamma \\ j 2 \boldsymbol{\omega} \mathbf{2} & \\ J 3 \boldsymbol{\omega} 3+\underset{\boldsymbol{W}(\boldsymbol{W h e e l})}{ } & \Omega \cos \gamma\end{array}\right]$

$$
\begin{equation*}
\dot{h}+\boldsymbol{\omega} \times \mathrm{h} \tag{4.42}
\end{equation*}
$$

The angle of gimbal is $\boldsymbol{\gamma}$ and the acceleration of the wheel is $\dot{\Omega}$ where both are used to produce the feedback torques needed to control the actuator input of the satellite [125].

$$
\dot{\gamma}=u 1, \quad \dot{\Omega}=u 2
$$

Gradient (Gravity) Torque: (Gt)
Assumptions: Upper-order relations ignored.
The Rigid Satellite body assumed (4.43).
$\mathrm{T}_{\text {gravity gradient }}=3\left[\left(\mathrm{I}_{3}-\mathrm{I}_{2}\right) \mathrm{mni}+\left(\mathrm{I}_{1}-\mathrm{I}_{3}\right) 1 \mathrm{ni}+\left(\mathrm{I}_{2}-\mathrm{I}_{1}\right) 1 \mathrm{mk}\right] \dot{\theta}^{2} /(1+\mathrm{e} \cos v \mathrm{v})(4.43)$
Variables: $\mathrm{T}_{\text {gravity gradient }}=$ gravity gradient torque [126]

$$
\begin{gathered}
v=\text { true anomaly } \\
e=\text { orbital eccentricity }
\end{gathered}
$$

$\mathrm{I}_{1}, \mathrm{I}_{2}, \mathrm{I}_{3}=$ moments of inertia of principle axis of body
$\mathrm{l}, \mathrm{m}, \mathrm{n},=($ satellite with principle axis of body $)$ - direction cosines
Gravity gradient stabilization:
Gravity gradient torque disappears:

- An angle between 2 vectors, cosine is zero.
- $\mathrm{I} 1=\mathrm{I} 2=\mathrm{I} 3$
- In the axis one cosine is zero other two moments of inertia equal

The Gravity Gradient Torque ( $\mathrm{T}_{\text {gravity }}$ ) is referred as a perturbation in the orbit [126]. The maximum moments of inactivity in the orbital plane \& minimum moment of inactivity in the local frame this configuration used to re-orient the stabilized satellite. A small mass injected from the satellite body for the de-saturation to stabilize the attitudes of satellite in gravity gradient method (GGM). The GGM method is used for low earth orbit satellite. For, High earth orbit such as geostationary earth orbit (GEO) it has less pointing accuracy to stabilize the satellite [127]. It requires the more effective damping elements/components. The eccentricity decides the size of the orbit. For, Elliptical orbit gravity gradient torques periodically changes the orbital elements.

Motion equation of satellite in the body frame

$$
\left[\begin{array}{l}
v 1  \tag{4.44}\\
v 2
\end{array}\right]=\left[\begin{array}{cc}
-\Omega * C * \gamma & -S * \gamma \\
\Omega * S * V & -C * \gamma
\end{array}\right]\left[\begin{array}{l}
u 1 \\
u 2
\end{array}\right] \quad \text { C }=\cos , S=\sin
$$

This is dynamic equation (4.44) of the satellite can be used to design the control technique of the satellite model. [127]

### 4.6. Perturbation Analysis of at Low Earth Orbiting satellite:

The satellite attitude governor/control is very important to stabilize the satellite along with its predetermined orientation. The perturbation forces or environmental disturbances may affect the satellites and the original orbit may get changed due to these perturbations. Thus, it is very important to reduce these perturbation forces as they can cause the life time of the satellite to be significantly reduced [128]. The environmental force differs at various altitudes. The proposed analysis considered for the simulations of LEO satellites affected more by aerodynamic drag and gravitational attraction due to the proximity of earth as compared to other perturbations.

### 4.6.1. Spacecraft considered: International Space Station (ISS)

The purpose of the International Space Station used for observation in Earth surface at low Earth's orbit satellite. The design for predicting the future missions to the planetary surface such as Mars, Moon, etc. First mission launched to the orbit 1998. It is affected by environmental perturbation, so it requires the frequent thruster to maintain the satellite into the actual orbit. The main aims for the mission are earth observation, trajectory transfer, predicts the future mission, Now ISS is used for commercial purpose and educational \& diplomatic conditions [129]. Its perigee is at

403 km and apogee is 406 km and inclination of 51.36 deg. The period of one revolution of this satellite is about 92.5 minutes and it completes about 15.54 orbits per day. It has an orbital velocity of $7.67 \mathrm{~km} / \mathrm{s}$. The initial values of ISS (See table 4.2) that are required for the perturbation calculation are given in Appendix B.

Table 4.2: Initial values of International Space Station (ISS) [129]

| Epoch Time (GMT) | $2018 / 092 / 13: 19: 31.516$ |
| :--- | :--- |
| Initial position [x, y, z] (m) | $[-3754663.80-5060641.17-2517733.24]$ |
| Initial velocity [u, v, w] (m/s) | $[5356.846188-1332.207337-5317.508141]$ |
| Mass of ISS (kg) | 411326.99584 |
| Area of ISS $\left(\mathbf{m}^{2}\right)$ | 1640.6 |

In Figure 4.18; describe the changes in ISS attitude such as position and velocity due to the perturbation forces action on it. This graph shows the change in altitude, earth radius with the altitude of the satellite changes with atmospheric perturbation [130]. The magnitude of position vector (4.45) is:

$$
\begin{equation*}
r=\sqrt{x^{2}+y^{2}+z^{2}} \tag{4.45}
\end{equation*}
$$

Let, Position vector $=\mathrm{x}_{\mathrm{i}}+\mathrm{y}_{\mathrm{j}}+\mathrm{z}_{\mathrm{k}}$

$$
\begin{aligned}
& \mathrm{x}=\text { distance of satellite measured in } \mathrm{x} \text {-direction }(\mathrm{m}) \\
& \mathrm{y}=\text { distance of satellite measured in } \mathrm{y} \text {-direction }(\mathrm{m})
\end{aligned}
$$

And, $\mathrm{z}=$ distance of satellite measured in z -direction (m)
Similarly, the velocity vector (4.46) of satellite is:

$$
\begin{equation*}
V=\sqrt{u^{2}+v^{2}+w^{2}} \tag{4.46}
\end{equation*}
$$

Where, $V=u i+v j+z k$
$u=$ velocity measured in the direction, $x(m / s)$
$v=$ velocity measured in the direction, $y(m / s)$

And, $\mathrm{w}=$ velocity measured in the direction, $\mathrm{z}(\mathrm{m} / \mathrm{s})$

The position and velocity vectors are shown in Figures.4.18. \& 4.19 respectively and their variations with respect to times use referred from appendix B.


Figure (4.18) Changes in position of the ISS with time
Since, the altitude of ISS satellite is about 350 to 420 km from the Earth's surface. Therefore, the Semi-Major axis will be from satellite altitude and earth center distance.

$$
\text { SMA }=R_{e}+\left(R_{\text {apogee }}+R_{\text {perigee }}\right) / 2
$$

Where, SMA $=$ Semi-Major Axis distance
$\operatorname{Re}=$ Radius of Earth $=6378 \mathrm{~km}, \mathrm{R}_{\text {apogee }}=$ Distance between satellite and Earth at apogee point, $\mathrm{R}_{\text {perigee }}=$ Distance between satellite and Earth at perigee point


Figure (4.19) Changes in velocity of the ISS time

The orbital components are the parameters required to remarkably recognize a particular orbits. In celestial mechanics these components are for the most part considered in establishing two-body frameworks, where a Kepler orbit is utilized. There is a wide range of approaches to numerically depict a similar orbit, however, certain plans, each comprising of an arrangement of six parameters, are ordinarily utilized as a part of space science and orbital mechanics. A genuine orbit (and its components) changes after some time because of gravitational irritations by different items and the impacts of relativity. A Keplerian orbit is simply a romanticized, scientific estimate at a specific time. The conventional orbital components are the six Keplerian components, after Johannes Kepler and his laws of planetary movement.

Source of Data: The data of the perturbation analysis are based on the sources mentioned below.

NASA 2-Line code- A two-line component set (TLE) is information arranges encoding a rundown of orbital components of an Earth-orbiting object for a given point in time, utilizing appropriate forecast recipe, the state (position and speed) any time before or future can be assessed with some exactness. TLEs can portray the directions just of the Earth-orbiting object [131]. The way to achieve goals is through a well-defined path. The path includes different stages of analysis and observations to be thoroughly kept in mind during the whole process.

### 4.6.2. Methodology

In flow chart shown in Figure (4.20) describe the Nano satellite perturbations calculation with different atmospheric condition using MATLAB Satellite Control ToolBox (SCT)


Figure (4.20) Flow chart of Nano Satellite Cowell's orbital perturbation
calculations (SCT)

In the Cowell's perturbation algorithms, Nano satellite Keplerian element due to variation in orbital perturbation is estimated. For, simulations considered the suitable design parameters and constants. In the thesis, discusses the perturbation effects in Nano satellites such as SRM satellite, Pratham (IITB) satellite considered. The evaluation of perturbation of satellite is briefly explained. The simulation results illustrate the variation in orbital elements is SMA, orbital inclination, Eccentricity, Argument of perigee, RAAN, True anomaly. For, reference analysis of motion and trajectory equations of satellite considered the International Space station (ISS). The following NORAD data considered for simulation. To implement's the Cowells algorithms using MATLAB environments.

## Satellite Selection:

The research focuses around the investigation of a satellite that has been orbiting around the Earth. This can be satellite from the Low Earth Orbit (LEO). A satellite in the LEO orbit is subjected to different drag powers and thus has expanded perturbations and probabilities along its way. The way of the satellite can be exceedingly circular along with little changes in the way and speed. To maintain a strategic distance from such radical conditions in the counts and keeping the examination up to its stamp, engaged towards LEO orbits.

## Raw Data Collection

The collection of data for simulations includes the Keplerian elements of the satellites which keep on changing every two weeks. The data used is as follows. (Refers Appendix A)

## International Space Station (ISS): Two-Line Elements

125544 U 98067 A
22554451.873 18300.65041667.00001569 00000-0 $351217-409996$
Epoch Format
Epoch

Figure, (4.21) shows the General Mission Analysis Tool describe the orbit with the Epoch formats. The GMAT considers the Universal Time Coordinates (UTC) Gregorian and coordinate frames EarthMJ2000Eq models in Keplerian orbit.


Figure (4.21) Low Earth Orbit analysis of the International Space Station epoch at 03 Apr 2018 13:19:31.516 UTC Gregorian EarthMU2000Eq, General Mission

## Analysis Tool (GMAT) with Keplerian Elements

In Figure (4.22) the propagator considered for the simulation is Runge-Kutta ODE Integrator with initial step size is 60 seconds, primary body an earth's surface. The Gravity model's EGM-96 has been introduced for both degree and order is 4 . The atmospheric model MSISE90 considered for GMAT simulations [132], [133]. The simulation minimum step size 0.001 and maximum step size is 2700 up to 50 attempts.


Figure (4.22) Selection of Propagator of ISS Primary bodies are Earth-Satellite given Gravity model EGM-96 and Atmosphere Model MSISE90

Table (4.3) ISS Keplerian, UTC Gregorian 03 Apr 18, 13:19:31.516 [129]

| Orbital Parameter's | Simulation values |
| :---: | :---: |
| Semi major axis (SMA) | 6786.4253 Km |
| Eccentricity | 0.000336 |
| Inclination | 51.873 Degree |
| RAAN | 35.151Degree |
| Argument of Perigee (AOP) | 135.0151 |
| Degree True Anomaly (TA) | 73.1269 Degree |

The following Keplerian simulations results obtained for International space station (ISS) epoch 03 Apr 2018 13:19:31.516 UTC Gregorian EarthMU2000Eq, the gravitational parameters consider for simulation is $3.98 * 10^{14}\left(\mathrm{~km}^{3} / \mathrm{Sec}^{2}\right)$ [133]. In table (4.3) reflect the design parameter of ISS Keplerian element at UTC Gregorian 03 Apr 2018 13:19:31.516. The following perturbation forces considered for the International space station includes aerodynamic drag, and solar radiation pressure. The validation of Cowell's simulation results compared with known Keplerian elements. The response of the known orbital elements is modeled by General Mission Analysis Tools. In the GMAT, the Gravity model EGM-96 and Atmosphere Model MSISE90 at Low Earth Orbit satellite has been introduced. The atmospheric model considered the valid assumption used for standard/ accurate results [132].


Figure (4.23) Variations of Argument of Perigee with time of International Space Station (ISS) Epoch 03 Apr 2018 13:19:31.516 UTC Gregorian

The argument of perigee usually varies between 0 to 360 degrees in the entire orbit. In perturbation analysis the argument of perigee changes periodically due to orbital perturbations. A large variation can be seen in Figure (4.23) due to proximity of Earth surface. During initial simulation time the value reach up 62 degree and at final time reach 140 degrees. The simulation error tolerance is considered as $\mathrm{e}^{-8}$ and step size is 30 Graphics simulations. Figure 4.23 shows increase in argument of perigee from initial to final time values. The result shown in orange color indicates the actual orbital trajectory and BLUE color indicates the predicted trajectory of the International space station modeled by GMAT initial step size 0.001 maximum step size 2700 for 50 attempts has been introduced the coordinate system EarthMU2000Eq. The simulation period of the ISS includes 0.07 times (days). In the simulation, results indicate variations in argument of perigee 80 degrees to 120 degrees between period 0.03 to 0.04 times (days). So, it is desired to model the disturbance equation using Cowell's perturbation ODE solver. The simulation of ODE obtained with minimum errors percentage is $2.67 \%$. The RA of Ascending Node (RAAN) has a secular behavior as well as a periodic variation with low amplitude. The variation is approximately 4.8 to 5.2 degrees between initial and final time.


Figure (4.24) Variations of RAAN with respect to time of International Space Station (ISS) Epoch 03 Apr 2018 13:19:31.516 UTC Gregorian

In the simulation, results indicate the variation in an argument of perigee 36 degrees to 30 degrees between the period 0 to 1.2 times (days). So, it is desired to model the disturbance equation using Cowell's perturbation ODE solver. The simulation of ODE obtained with minimum errors percentage is $0.87 \%$. The variations of right assertion ascending node (RAAN) with respect to the vernal equinox (VE) are shown in Figure 4.24. The angular deviation from perigee 403 Km to vernal equinox is $30^{\circ}$ of the International Space station. In the simulation the Right Ascension of Ascending Node (RAAN) from intial period 36 deg to 30 deg changes during time 1 (days).

In the simulation, results shows the position of the ISS with respect to perigee. Hence, the variation is very large with respect to initial and final time between 0 to 0.045 times (days). The Figure 4.25 shows an increment in True Anomaly with respect to time period. It is clearly observed the position of satellite in orbit continuously varies with respect to time period. The simulation error tolerance is considered as $\mathrm{e}^{-8}$ and step size is 30 Graphics simulations.


Figure (4.25) Variations of True Anomaly with respect to time of International Space Station (ISS) Epoch 03 Apr 2018 13:19:31.516 UTC Gregorian

The variation is approximately 96 to 300 degrees between initial and final time in orbit. Figure 4.25. shows an increment in True Anomaly (RAAN) with respect to time period. The simulation of ODE obtained with minimum 1.02 percentage (\%) of error deviations. Since the apogee and perigee of the ISS satellite are almost equal, therefore, the eccentricity will be small i.e., near to 0 , which means that the orbit is almost circular in nature at Low Earth Orbit. In Figure 4.26 shows the actual trajectory (Orange color) and predicted trajectory (Blue color) variation of eccentricity with respect to time periods.


Figure (4.26) Variations of Eccentricity with respect to the time for International Space Station (ISS) Epoch 03 Apr 2018 13:19:31.516 UTC Gregorian

Eccentricity oscillates in orbit due to perturbation forces. The eccentricity decides the shape of the orbit. The simulation periods considered of ISS is 0.05 times (days). The variation of eccentricity from initial simulation time is 0.00036 to final time is 0.00020 in LEO orbit. The orange colour indicates the actual trajectory and BLUE colour indicates predicted trajectory. The Cowell's perturbation algorithm simulation of ODE obtained with minimum errors percentage is $3.97 \%$.

Figure 4.27 shows the variation between orbital inclination and time periods of orbit. The overall change in orbital inclination is 51.875 degrees to 51.845 degrees from initial to final time. The orbital inclination is angle made by satellite orbital
plane with respect to Earth's equatorial plane. The change in inclination will change the orientation of the transmitting and receiving antennas on satellite which will make communication of satellite very difficult to receive the information. Therefore, the orbital inclination must be maintained at all time under any perturbation condition.


## Figure (4.27) Variations of Orbital Inclination with respect to International Space Station (ISS) Epoch 03 Apr 2018 13:19:31.516 UTC Gregorian

The variations of inclination measured from Epoch 03 Apr 2018 13:19:31.516 UTC Gregorian. The Keplerian EarthMJ2000Eq has been introduced 0.045 times (days). The simulation of actual trajectory to the predicted trajectory of orbit obtained with minimum 0.002 errors percentage (\%) of the orbit. In Figure 4.28 shows the variations of Semi-major axis with respect to time which varies at an equal interval. As due to aerodynamic drag perturbations, the altitude decreases due to which semimajor axis decreases within a time variation from 0 to 0.040 times (days). This decrease can be prevented by using control mechanisms such as thrusters (DC motor) to reorient the satellite whenever its orientation is disturbed due to any perturbations in the orbit. (discussed in chapter 6 attitude control system) The satellite's SemiMajor axis reduces from 6786 to 6777 km i.e., an altitude changes of almost 9 km within 0.040 times (days). The orange colour indicates the actual trajectory and BLUE colour indicates the Cowell's perturbation algorithm (ODE Solver) predicted
trajectory of the orbit. The simulation of ODE obtained with minimum errors percentage is $0.002 \%$ from actual trajectory to the predicted trajectory.


Figure (4.28) Variations of Semi-Major Axis with respect to International Space

## Station (ISS) Epoch 03 Apr 2018 13:19:31.516 UTC Gregorian

The path travelled by satellite can be visualized by the help of GMAT (General Mission Analysis Tool) software in which the Cartesian state vectors, i.e., position and velocity or Keplerian elements are used as the input at a point of time, known as Epoch and the path of the satellite over the Earth's surface is the output shown in the simulations. The ground track plots (Figure 4.29) shows the orbit view of the satellite corresponding to epoch time 03 April 2018 13:19:31.516 with considered atmospheric drag model in GMAT.


Figure (4.29) Ground Track Plot at Epoch - 03 Apr 2018 13:19:31.516


Figure (4.30) Orbit View at Epoch - 03 Apr 2018 13:19:31.516


Figure (4.31) Orbit view at EarthMJ2000Eq UTC
The Figure $4.30 \& 4.31$ show the orbit view at Epoch with corresponding to final values of position and velocity which includes the Atmospheric Drag model in GMAT solved by ordinary differential equation (ODE) with the help of Runge Kutta method through MATLAB Environment.

Table (4.4) Magnitudes of errors of Keplerian elements of International Space Station (ISS) with actual and predicted trajectory in the orbit with error variation

| Semi-Major Axis (Km) |  |  |  | Orbital Inclination (deg) |  |  | Argument of Perigeee (deg) |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $\begin{array}{c}\text { Magnitudes } \\ \text { of Actual } \\ \text { (GMAT) }\end{array}$ | $\begin{array}{c}\text { Magnitudes } \\ \text { of Predicted } \\ \text { (Cowell's) } \\ \text { Simulations }\end{array}$ | $\begin{array}{c}\text { Error } \\ \text { Simulations }\end{array}$ | $\begin{array}{c}\text { Magnitudes } \\ \text { (\%) }\end{array}$ | $\begin{array}{c}\text { Magnitudes of } \\ \text { of Actual } \\ \text { (GMAT) } \\ \text { Simulations }\end{array}$ | $\begin{array}{c}\text { Predicted } \\ \text { (Cowell's) } \\ \text { Simulations }\end{array}$ | $\begin{array}{c}\text { Error } \\ \text { Deviations } \\ \text { (\%) }\end{array}$ | $\begin{array}{c}\text { Magnitudes } \\ \text { of Actual } \\ \text { (GMAT) } \\ \text { Simulations }\end{array}$ | $\begin{array}{c}\text { Magnitudes } \\ \text { of Predicted } \\ \text { (Cowell's) } \\ \text { Simulations }\end{array}$ |  | \(\left.\begin{array}{c}Error <br>

Deviations <br>
(\%)\end{array}\right\}\)

| True Anomaly (deg) |  |  | RAAN (deg) |  |  | Eccentriciy (deg) |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Magnitudes of Actual (GMAT) Simulations | Magnitudes <br> of Predicted <br> (Cowell's) <br> Simulations | Error Deviations (\%) | Magnitudes of Actual (GMAT) <br> Simulations | Magnitudes <br> of Predicted <br> (Cowell's) <br> Simulations | Error Deviations (\%) | Magnitudes of Actual (GMAT) <br> Simulations | Magnitudes <br> of Predicted <br> (Cowell's) <br> Simulations | Error <br> Deviations <br> (\%) |
| 110 | 96 | 14 | 36 | 36 | 0 | 0.00039 | 0.00036 | 0.00003 |
| 155 | 150 | 5 | 35 | 35 | 0 | 0.00039 | 0.00037 | 0.00002 |
| 222 | 222 | 0 | 34 | 34 | 0 | 0.00039 | 0.00037 | 0.00002 |
| 280 | 280 | 0 | 33 | 33 | 0 | 0.000265 | 0.000265 | 0 |
| 306 | 304 | 2 | 32.8 | 32 | 0.8 | 0.0002 | 0.0002 | 0 |

The table 4.4 shows the International space station magnitudes of actual simulation (GMAT) and magnitudes of predicted simulations (Cowell's) with minimum error deviation of Keplerian elements in reference low earth's orbit at UTC 03 Apr 2018 13:19:31.516 Gregorian's. The trajectories of GMAT simulation and the Cowell's simulation are close to the nearest integer. The Cowell's predicts the accurate trajectory the satellite deviates from actual orbit. This data is very much useful to estimate the errors in attitudes used to design the proper attitude control techniques. (discussed in chapter 6)

### 4.6.3. Pratham (IIT- Bombay) Satellite perturbation analysis at LEO.

The Pratham (IIT Bombay) satellite Keplerian elements consider for the simulations is epoch 08 Jul 2018 04:53:04.000 UTC Gregorian EarthMJ2000Eq, the gravitational parameters that has been introduced for simulation is $3.98 * 10^{14}\left(\mathrm{~km}^{3} / \mathrm{Sec}^{2}\right)$ [134]

## Methodology

Flow chart shown in Figure (4.32) discuss the satellite Keplerian elements of Nano satellite perturbation analysis using Global Mission Analysis tools (GMAT).


Figure (4.32) Flow chart of Satellite Keplerian GMAT Simulation
The simulation time is minimum step size 0.001 and maximum step size is 2700 up to 50 attempts. The Keplerian elements are referred from NORAD 2-line elements. The perturbation forces aerodynamic drag, solar pressure considered for the simulations. In Table (4.5) considered the Pratham (IITB) satellite simulation parameters using GMAT. In Global Mission Analysis tools (GMAT) perturbation algorithms, Nano satellite Keplerian element due to variation in orbital perturbation is estimated. For,
simulations considered the suitable design parameters and constants. In the thesis, discusses the perturbation effects in Nano satellites such as SRM satellite, Pratham (IITB) satellite considered. The simulation results illustrate the variation in orbital elements is SMA, orbital inclination, Eccentricity, Argument of perigee, RAAN, True anomaly. The following NORAD data considered for simulation [135]. To implement's the algorithms using Global Mission Analysis tools (GMAT).

Table (4.5) Pratham satellite simulation data [135]

| Epoch Format | UTC Gregorian |
| :--- | :--- |
| Epoch | 08 Jul 2018 04:53:04.000 |
| Coordinate System | EarthMJ2000Eq |
| State Type | Keplerian |

NORAD Two-Line Elements of Pratham (Refer Appendix A)
141783 U 16059A 18295.14940496 .00000065 00000-0 21820-4 09994
24178398.116250 .4222003429931 .8295302 .121014 .62950175110526

The data used for perturbation algorithms is given below
Figure (4.33) shows the Universal Time Coordinates as Epoch at 08 Jul 2018 04:53:04.000 UTC Gregorian with Keplerian coordinates EarthMJ2000Eq.


Figure (4.33) LEO analysis of the Pratham, Epoch at 08 Jul 2018 04:53:04.000

## UTC Gregorian EarthMJ2000Eq GMAT with Keplerian coordinates

The Keplerian simulations results obtained for Pratham satellite (IIT Bombay), Mass, 10 kg , Size $26 \times 26 \times 26 \mathrm{~cm}$, Orbital angular velocity, $\Omega=0.0010346,817 \mathrm{~km}$ altitude, Epoch 08 Jul 2018 04:53:04.000 UTC Gregorian, EarthMJ2000Eq. [134]


Figure (4.34) Selection of Propagator of Pratham Primary bodies are Earth-
Satellite given Gravity model JGM-2 and Atmosphere Model MSISE90
The atmospheric model has been introduced MSISE90 for the GMAT simulations. During simulation, minimum step size 0.001 and maximum step size is 2700 up to 50 attempts. In the perturbation algorithm considered the Runge-Kutta ODE Integrator with initial step size is 60 seconds, primary body an earth's surface. The Gravity models JGM-2 for degree and order is 4 .

Table (4.6) Pratham (IITB) satellite NORAD data, Epoch at 08 Jul 2018
04:53:04.000 UTC Gregorian EarthMJ2000Eq [134]

| Keplerian Details | NORAD Data |
| :---: | :---: |
| NORAD ID | 41783 |
| Int'l Code | $2016-059 \mathrm{~A}$ |
| Perigee | 666.3 km |
| Apogee | 715.3 km |
| Inclination | $98.1^{\circ}$ |
| Period | 98.4 minutes |
| Semi major axis | 7061 km |

In table (4.6) shows the NORAD data of Pratham satellite, Epoch at 08 Jul 2018 04:53:04.000 UTC Gregorian EarthMJ2000Eq


Figure (4.35) Ground Track Plot at Epoch - 08 Jul 2018 04:53:04.000
The path travelled by satellite can be visualized by the help of GMAT (General Mission Analysis Tool) software in which the Cartesian state vectors, i.e., position and velocity or Keplerian elements are used as the input at a point of time, known as Epoch and the path of the satellite over the Earth's surface is the output shown in the simulations. The ground track plots. In Figure 4.35 shows the orbit view of the satellite corresponding to the epoch time 08 Jul 2018 04:53:04.000 with an atmospheric drag model in GMAT.


Figure (4.36) Pratham orbit view at EarthMJ2000Eq UTC
The Figure 4.36 shows the orbit view at Epoch with corresponding final values of position and velocity of satellite. In the simulation parameters included the Atmospheric Drag model using GMAT/ MATLAB solved by the ordinary differential equation (ODE) with the help of Runge - Kutta methods. The gravitational parameters considered for simulation is $3.98 * 10^{14}\left(\mathrm{~km}^{3} / \mathrm{Sec}^{2}\right)$. The perturbation forces for Pratham includes are, aerodynamic drag, and solar pressure. The validation
of Cowell's simulation results compared with known Keplerian elements. The response of the known orbital elements is modeled by General Mission Analysis Tools. In the GMAT, Gravity model JGM-2 and Atmosphere Model MSISE90 at low Earth orbiting satellite has been introduced. The atmospheric model considered the valid assumption used for standard/ accurate results.



Figure (4.37) The orbital variations of Pratham satellite Epoch at 08 Jul 2018

## 04:53:04.000 UTC Gregorian EarthMU2000Eq, $\boldsymbol{\Omega}=\mathbf{0} .0010346 \mathrm{rad} / \mathrm{s}$ in LEO (a)

 Semi major axis (b) Argument of perigee (c) Eccentricity (d) Orbital Inclination(e) Right Ascension of Ascending Node (f) True Anomaly

The simulation error tolerance is considered as $\mathrm{e}^{-8}$ and step size is 30 Graphics simulations. In Figure (4.37) shows the Pratham satellite design of Cowells simulations and GMAT simulations magnitudes of actual orbital trajectory (orange color) and magnitudes of predicted trajectory (blue color) in the orbit.

Table (4.7) (a) Magnitudes of errors of Keplerian elements of Pratham satellite with actual and predicted trajectory in the orbit

| Simulation <br> Periods <br> (Days) | Semi-Major Axis (Km) |  | Errors deviations (\%) | Argument of Perigee (Deg) |  | Errors deviations (\%) | Eccentricity |  | Errors deviations (\%) |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | Magnitudes of actual data (GMAT) Simulations | Magnitudes of Predicted data (Cowel's) Simulations | 0.003 | Magnitudes of actual data (GMAT) Simulations | Magnitudes of Predicted data (Cowel's) Simulations | 0.92 | Magnitudes of actual data (GMAT) Simulations | Magnitudes of Predicted data (Cowel's) Simulations | 1.92 |
| 0.005 | 7057 | 7057.2 |  | 54 | 54 |  | 0.00347 | 0.00337 |  |
| 0.0075 | 7055.1 | 7055 |  | 52 | 51 |  | 0.00329 | 0.00322 |  |
| 0.01 | 7053.8 | 7053.4 |  | 49 | 49 |  | 0.00325 | 0.00319 |  |
| 0.0125 | 7051.6 | 7051 |  | 46 | 45 |  | 0.00305 | 0.003 |  |
| 0.015 | 7049.5 | 7049.2 |  | 44 | 44 |  | 0.00299 | 0.00297 |  |
| 0.0175 | 7047.7 | 7047.7 |  | 42 | 42 |  | 0.0029 | 0.0028 |  |
| 0.02 | 7046 | 7045.6 |  | 40 | 39 |  | 0.00276 | 0.00257 |  |

The Pratham satellite modeled by GMAT initial step size 0.001 maximum step size 2700 for 50 attempts considered the coordinate system EarthMJ2000Eq. The simulation period of the Pratham satellite is taking into 0.020 times (days). In table (4.7) shown the simulation of ODE obtained with minimum error deviations.

Table (4.7) (b) Magnitudes of errors of Keplerian elements of Pratham satellite with actual and predicted trajectory in the orbit (Orbital Inclination, RAAN, True Anomaly)

| Orbital inclination |  | Errors deviations (\%) | RAAN |  | Errors deviations (\%) | True Anomaly |  | Errors deviations (\%) |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Magnitudes of actual data (GMAT) Simulation | Magnitudes of Predicted data Cowel's Simulation | 0.0001 | Magnitudes of actual data <br> (GMAT) <br> Simulation | $\begin{array}{\|l\|} \hline \text { Magnitudes } \\ \text { of } \\ \text { Predicted } \\ \text { data } \\ \text { Cowel's } \\ \text { Simulation } \\ \hline \end{array}$ | 0.0006 | Magnitudes of actual data (GMAT) <br> Simulation | $\begin{array}{\|l\|} \hline \text { Magnitudes } \\ \text { of } \\ \text { Predicted } \\ \text { data } \\ \text { Cowel's } \\ \text { Simulation } \\ \hline \end{array}$ | 0.85 |
| 98.1058 | 98.1058 |  | 354.5666 | 354.55 |  | 250 | 250 |  |
| 98.1069 | 98.107 |  | 354.572 | 354.56 |  | 220 | 225 |  |
| 98.108 | 98.108 |  | 354.576 | 354.57 |  | 190 | 190 |  |
| 98.1094 | 98.10892 |  | 354.581 | 354.58 |  | 160 | 160 |  |
| 98.1104 | 98.1102 |  | 354.584 | 354.59 |  | 140 | 138 |  |
| 98.1117 | 98.1112 |  | 354.586 | 354.592 |  | 120 | 120 |  |
| 98.1123 | 98.1125 |  | 354.588 | 354.595 |  | 98 | 85 |  |

In table 4.7 (b) shows the Pratham satellite's minimum error deviation of Keplerian elements with reference low earth's orbit at UTC 08 Jul 2018 04:53:04.000 Gregorian. The magnitudes of actual and predicted trajectories of GMAT simulation and Cowell's simulation are close to the nearest integer. The Cowell's predicts the actual path from perturbed path of satellite in the orbit. This data is very much useful to estimate the errors in the attitudes and to design the proper attitude control techniques discussed in chapter 6.

### 4.6.4. Perturbation analysis of SRM Satellite

The Keplerian simulations considered for SRM satellite is epoch 08 Jul 2018 04:53:04.000 UTC Gregorian EarthMJ2000Eq, the gravitational parameters consider for simulation is $3.98 * 10^{14}\left(\mathrm{~km}^{3} / \mathrm{Sec}^{2}\right)$. The perturbation forces like aerodynamic drag, solar pressure is considered [136].

The following data is used for simulations.
Table (4.8) SRM Satellite design parameters [136]

| Epoch Format | UTC Gregorian |
| :--- | :--- |
| Epoch | 08 Jul 2018 10:32:46.000 |
| Coordinate System | EarthMU2000Eq |
| State Type | Keplerian |

The simulation minimum step size 0.001 and maximum step size is 2700 up to 50 attempts. The Keplerian elements are referred from NORAD 2-line elements.


Figure (4.38) LEO analysis of SRM Satellite epoch at 08 Jul 2018 10:32:46.000

## UTC Gregorian EarthMU2000Eq GMAT with Keplerian Elements

In table (4.38) shows the SRM satellite design parameters at Universal Time Coordinates as Epoch at 08 Jul 2018 10:32:46.000 UTC Gregorian with Keplerian coordinates EarthMJ2000Eq In Figure (4.38) shows the SRM satellite GMAT simulation considered the UTC 08 Jul 2018 10:32:46.000.

Table (4.9) SRM satellite NORAD data [137]

| Keplerian Elements | NORAD Data |
| :---: | :---: |
| NORAD ID | 37841 |
| Int'l Code | $2011-058 \mathrm{D}$ |
| Perigee | 855.7 km |
| Apogee | 873.1 km |
| Inclination | $20.0^{\circ}$ |
| Period | 102.1 minutes |
| Semi major axis | 7235 km |

In table (4.9) shows the SRM satellite NORAD data at 08 Jul 2018 10:32:46.000

UTC Gregorian EarthMU2000Eq. The atmospheric model has been introduced MSISE90 for GMAT simulations. The simulation minimum step size 0.001 and maximum step size is 2700 up to 50 attempts.

| Integrator |  |  |
| :---: | :---: | :---: |
| Type | RungeKutta89 |  |
| Initial Step Size | 60 | sec |
| Accuracy | 9.999999999999999e-012 |  |
| Min Step Size | 0.001 | sec |
| Max Step Size | 2700 | sec |
| Max Step Attempts | 50 |  |
|  | $\checkmark$ Stop If Accuracy Is Violated |  |



Figure (4.39) Selection of Propagator of SRM satellite Primary bodies are EarthSatellite given Gravity model JGM-2 and Atmosphere Model MSISE90
In the Figure (4.39) the propagator considered the Runge-Kutta ODE Integrator with initial step size is 60 seconds, primary body an earth's surface. The Gravity models JGM-2 for degree and order is 4 .


Figure (4.40) Ground Track Plot at Epoch - 08 Jul 2018 10:32:46.000 (UTC)

The path travelled by satellite can be visualized by the help of GMAT (General Mission Analysis Tool) software in which the Cartesian state vectors, i.e., position and velocity or Keplerian elements are used as the input at a point of time, known as Epoch and the path of the satellite over the Earth's surface is the output shown in the
simulations. The ground track plots. In Figure 4.41 shows the orbit view of the satellite corresponding to the epoch time 08 Jul 2018 04:53:04.000 with an atmospheric drag model in GMAT.


Figure (4.41) Pratham orbit view at EarthMJ2000Eq UTC
The Figure 4.41 shows the orbit view at Epoch with corresponding to the final values of position and velocity. The atmospheric Drag model developed in GMAT/ MATLAB by ordinary differential equation (ODE) with the help of Runge - Kutta methods.

The SRM satellite Keplerian elements considered for simulation is Mass, 10 kilograms ( 22 lb ), Orbital angular velocity, $\Omega=0.0010239$ ( $\mathrm{rad} / \mathrm{sec}$ ), 867 km altitude, Epoch 08 Jul 2018 10:32:46.000 UTC Gregorian EarthMU2000Eq. The gravitational parameters consider for simulation is $3.98 * 10^{14}\left(\mathrm{~km}^{3} / \mathrm{Sec}^{2}\right)$ [136]. The SRM Satellite perturbation forces of aerodynamic drag, solar pressure is considered. The magnitude of actual path and predicted path explained in Cowell's/GMAT simulations. The validation of Cowell's simulation results compared with known Keplerian elements. The response of the known orbital elements is modeled by General Mission Analysis Tools. In GMAT has been introduced the Gravity model EGM-96 and Atmosphere Model MSISE90 at Low Earth Orbit satellite [137]. The atmospheric model considered the valid assumption used for standard/ accurate results.



Figure (4.42) The orbital variations of SRM satellite Epoch at 08 Jul 2018

## 10:32:46.000 UTC Gregorian EarthMU2000Eq, $\boldsymbol{\Omega}=\mathbf{0} .0010239 \mathrm{rad} / \mathrm{s}$ in LEO (a)

## True Anomaly (b) Argument of perigee (c) Orbital Inclination (d) Eccentricity

(e) Right Ascension of Ascending Node (f) Semi major axis

The errors tolerance is considered as $\mathrm{e}^{-8}$ and step size is 30 Graphics simulations. The result shown in figure (4.40) RED color indicates the actual orbital trajectory and BLUE color indicates the Cowell's perturbation trajectory in the orbit. The SRM satellite is modeled by GMAT initial step size 0.001 maximum step size 2700 for 50 attempts considered coordinate system EarthMU2000Eq. The simulation period of the SRM satellite taking into 0.020 times (days). The simulation of ODE obtained with minimum error percentage shown in table (4.4)

Table (4.10) (a) SRM satellite magnitudes of errors with actual and predicted trajectory in the orbit (True Anomaly, Argument of perigee, Orbital Inclination)

| Orbital Elements | True Anomaly |  | Errors <br> deviations (\%) | Argument | of Perigee | Errors deviations (\%) | Orbital in | inclination | Errors deviations (\%) |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | Magnitudes of actual data (GMAT) Simulation | Magnitudes of <br> Predicted data <br> Cowel's <br> Simulation | 0.42 | Magnitudes of actual data (GMAT) Simulation | $\begin{aligned} & \text { Magnitudes } \\ & \text { of } \\ & \text { Predicted } \\ & \text { data } \\ & \text { Cowel's } \\ & \text { Simulation } \end{aligned}$ | 2.43 | $\begin{gathered} \text { Magnitudes } \\ \text { of actual } \\ \text { data } \\ \text { (GMAT) } \\ \text { Simulation } \end{gathered}$ | $\begin{aligned} & \text { Magnitudes } \\ & \text { of } \\ & \text { Predicted } \\ & \text { data } \\ & \text { Cowel's } \\ & \text { Simulation } \end{aligned}$ | 0.0001 |
| 0.005 | 250 | 250 |  | 90 | 95 |  | 19.974 | 19.9725 |  |
| 0.0075 | 210 | 221 |  | 96 | 102 |  | 19.9746 | 19.9749 |  |
| 0.01 | 162 | 168 |  | 103 | 110 |  | 19.975 | 19.9759 |  |
| 0.0125 | 132 | 125 |  | 113 | 120 |  | 19.9761 | 19.976 |  |
| 0.015 | 98 | 95 |  | 122 | 126 |  | 19.9768 | 19.9771 |  |
| 0.0175 | 68 | 68 |  | 136 | 130 |  | 19.9771 | 19.9775 |  |
| 0.02 | 25 | 22 |  | 142 | 139 |  | 19.9781 | 19.978 |  |

Table (4.10) (b) SRM satellite magnitudes of errors with actual and predicted trajectory in the orbit (Eccentricity, RAAN, Semi-major axis)

| Eccentricity |  | Errors <br> deviations (\%) | RA | AN | Errors <br> deviations (\%) | Semi-M | ajor Axis | Errors deviations (\%) |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Magnitudes of actual data (GMAT) Simulation | Magnitudes <br> of <br> Predicted <br> data <br> Cowel's <br> Simulation | 1.9 | Magnitudes of actual data (GMAT) Simulation | Magnitudes of <br> Predicted data <br> Cowel's Simulation | 0.03 | Magnitud es of actual data (GMAT) Simulation | Magnitudes of <br> Predicted data <br> Cowel's <br> Simulation | 0.0004 |
| 0.00129 | 0.00131 |  | 43.12 | 43.16 |  | 7235.14 | 7235.2 |  |
| 0.00134 | 0.00139 |  | 43.11 | 43.1 |  | 7235.2 | 7235.25 |  |
| 0.00141 | 0.00143 |  | 43.08 | 43.05 |  | 7235.24 | 7235.29 |  |
| 0.00146 | 0.0015 |  | 43.06 | 43.02 |  | 7235.35 | 7235.37 |  |
| 0.00152 | 0.00158 |  | 43.03 | 43 |  | 7235.39 | 7235.39 |  |
| 0.0016 | 0.00162 |  | 43 | 42.97 |  | 7235.42 | 7235.46 |  |
| 0.00168 | 0.00167 |  | 42.96 | 42.95 |  | 7235.5 | 7235.52 |  |

In the table 4.10 (a) and (b) shows the SRM satellite actual and predicted trajectory minimum error deviation of Keplerian elements with reference low earth's orbit at UTC 08 Jul 2018 10:32:46.000 Gregorian. The orbital path of GMAT/ Cowell's simulation is close to the nearest integer. The Cowell's algorithm predicts the accurate trajectory the satellite from disturbed orbit. This data is very much useful to estimate the errors in the attitudes and to design the proper attitude control techniques discussed in chapter 6. Attitude control design

|  | SRM satillite |  |  |  |  |  | Pratham satelite |  |  |  |  |  | Internationa Space Station (ISS) |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Period <br> (Days) | Semi- <br> Major <br> Axis <br> (Km) | Orbital <br> Inclindion <br> (deg) | Argument <br> of <br> Pergeee <br> (deg) | True <br> Anomady <br> (deg) | $\binom{\mathrm{RAAN}}{(\mathrm{deg})}$ | Ecc | Semi- <br> Major <br> Axis <br> (Km) | Orbital <br> Inclination <br> (deg) | Argument <br> of <br> Perigee <br> (deg) | True <br> Anomady <br> (deg) | $\begin{aligned} & \text { RAAN } \\ & (\mathrm{deg}) \end{aligned}$ | Ecc | Semi- <br> Major <br> Axis <br> (Km) | Orbital <br> Inclination <br> (deg) | Argment <br> of <br> Pergeee <br> (deg) | True <br> Anomady <br> (deg) | RAAN <br> (deg) | Ecc |
| 0.005 | 7235.1 | 19.974 | 90 | 250 | 43.12 | 0.0012 | 7057 | 98.1058 | 54 | 250 | 354.567 | 0.00347 | 6786.2 | 51.875 | 62 | 110 | 36 | 0.00039 |
| 0.0075 | 7235.2 | 19.9746 | 96 | 210 | 43.1 | 0.0012 | 7055.1 | 98.1069 | 52 | 220 | 354.572 | 0.00329 | 6785.9 | 51.872 | 8 | 155 | 35 | 0.00039 |
| 0.01 | 7235.2 | 19.975 | 103 | 162 | 43.05 | 0.00141 | 7053.8 | 98.108 | 49 | 190 | 354.576 | 0.00325 | 6,784 | 51.867 | 101 | 22 | 34 | 0.00039 |
| 0.012 | 7235.4 | 19.9761 | 113 | 132 | 43.02 | 0.00146 | 7051.6 | 98.1094 | 46 | 160 | 354.581 | 0.00305 | 6780 | 51.85 | 120 | 280 | 33 | 0.00027 |
| 0.015 | 7235.4 | 19.976 | 12 | 98 | 43 | 0.00152 | 7049.5 | 98.1104 | 44 | 140 | 354.584 | 0.0029 | 6777.2 | 51.845 | 120 | 306 | 32.8 | 0.002 |



|  | SRM satillite |  |  |  |  |  | Pratham satellite |  |  |  |  |  | Interational Space Station (ISS) |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $\begin{aligned} & \text { Period } \\ & \text { (Days) } \end{aligned}$ | Semi- <br> Major <br> Axis <br> (Km) | $\begin{gathered} \text { Oftital } \\ \text { Inclimation } \\ \left(\begin{array}{c} \text { degg } \end{array}\right. \\ \hline \end{gathered}$ | Argment <br> of <br> Perigee <br> (deg) | Tne <br> Anomady <br> (deg) | $\begin{aligned} & \text { RAAN } \\ & (\operatorname{deg}) \end{aligned}$ | Ecc | Semi- <br> Major <br> Axis <br> (Km) | Orbita <br> Inclination <br> (deg) | Argment <br> of <br> Perigee <br> (deg) | True <br> Anomady <br> (deg) | $\left.\left\lvert\, \begin{array}{c} \text { RAAN } \\ (\mathrm{deg} \end{array}\right.\right)$ | Ecc | Semi- <br> Major <br> Axis <br> (Km) | Orbital <br> Inclination <br> (deg) | Argment <br> of <br> Pergree <br> (deg) | True <br> Anomady <br> (deg) | RAAN <br> (deg) | Ecc |
| 0.005 | 7235.2 | 19.975 | 95 | 250 | 43.16 | 0.00131 | 7057.2 | 98.1058 | 54 | 250 | 354.55 | 0.00337 | 6786 | 51.875 | 55 | 96 | 36 | 0.00036 |
| 0.0075 | 7235.3 | 19.974 | 102 | 221 | 43.1 | 0.00139 | 7055 | 98.107 | 51 | 225 | 354.56 | 0.0032 | 6785.2 | 51.866 | 80 | 150 | 35 | 0.00037 |
| 0.01 | 7235.3 | 19.975 | 110 | 168 | 43.05 | 0.0014 | 7053.4 | 98.108 | 49 | 190 | 354.57 | 0.00319 | 6782.8 | 51.864 | 100 | 22 | 34 | 0.00037 |
| 0.0125 | 7235.4 | 19.976 | 120 | 125 | 43.02 | 0.0015 | 7051 | 98.10892 | 45 | 160 | 354.58 | 0.003 | 6779.9 | 51.855 | 110 | 280 | 33 | 0.00027 |
| 0.015 | 7235.4 | 19.9771 | 126 | 95 | 43 | 0.00158 | 7049.2 | 98.1102 | 44 | 138 | 354.59 | 0.0029 | 677 | 51.845 | 120 | 304 | 32 | 0.002 |



The SRM satellite, Pratham (IITB) Satellite, International Space station (ISS) variations in Keplerian elements using GMAT Simulation shown in table (4.6). The table clearly indicates the changes in ISS position and velocity due to the perturbation forces action on it. This result shows the change in altitude, earth radius with the altitude of the satellite changes with atmospheric perturbation. ISS perigee position is at 403 km and apogee position is 406 km and inclination of 51.36 deg . The period of one revolution of this satellite is about 92.5 minutes and it completes about 15.54 orbits per day an orbital velocity of 7.67 km . Since, the altitude of ISS satellite is about 350 to 420 km from the Earth's surface. Therefore, the Semi-Major axis will be from satellite altitude and earth center distance. Epoch Format: UTC Gregorian, Epoch: 03 Apr 2018 13:19:31.516, the time period consider for the simulation is 0.005 to 0.015 (Days), it is clearly understood the semi-major axis vary from 6786.2 Km to 6777.2 Km . The altitudes nearly decrease to 9 km . Also, discuss the other Keplerian elements variation in shown in the table (4.6). The Keplerian simulations results obtained for Pratham (IIT Bombay) satellite at Epoch 08 Jul 2018 04:53:04.000UTC Gregorian EarthMJ2000Eq (Refers Appendix C). The Keplerian simulations considered for Pratham satellite (IIT Bombay) Mass, 10 kg , Size $26 \times 26 \times 26 \mathrm{~cm}$, Orbital angular velocity, $\Omega=0.0010346,817 \mathrm{~km}$ altitude. The Gravity models, JGM-2 and atmospheric model, MSISE90 has been introduced for degree and order is 4. The SRM satellite considers the Universal Time Coordinates as Epoch at 08 Jul 2018 10:32:46.000 UTC Gregorian. The Runge-Kutta propagator considered ODE Integrator with initial step size is 60 seconds, primary body an earth's surface. In table (4.7) shows the variation in the Keplerian elements Cowell's Simulation of SRM satellite, Pratham (IITB) Satellite, International Space station (ISS). The table (4.6) \& (4.7) shows the comparative analysis of Keplerian elements and error deviation in GMAT/Cowell's simulation. The Cowell's simulation used to validate the GMAT simulation measures the minimum error deviations. This is used to generate the controlled input to actuator re-orient satellite in to the desired attitude with atmospheric conditions. Also, it is helping to design the suitable controller of satellite.

## CHAPTER 5

## ATTITUDE ERROR ESTIMATION ANALYSIS USING KALMAN FILTER

The linear estimates of measurement and prediction of the state (or) information from the attitude sensor using the Kalman Filter (KF) are widely used. The KF is more efficient and accurate method used to predict the performance in the linear system [138]. The Non-linear estimate of measurements and predictions of the state (or) information's from the attitude sensor using Extended Kalman Filter (EKF). The Extended KF is more efficient and accurate method used to predict the performance in non-linear system. The unscented transformation (UT) occurs in the state equations [139]. The nonlinear system estimated by first and second order ordinary differential equations (ODE) in UKF, EKF and UKF are unlike for derivatives or Jacobians in the estimate the performance equation and state equation. It requires the propagating state covariance matrix, filter circulates state, and standard deviations are circulated. This indicates the random changes in the state or uncertainty of the system. In Figure (5.1) illustrate the estimated attitude with angular rates of satellite using Kalman filter [140].


Figure (5.1) Attitude Error corrections with Kalman Filter [140]

### 5.1. Algorithms of error estimation using Kalman Filter

The state vectors in the system/plant dynamics predict and estimate of errors using Kalman filter. It is based upon the past estimates in process and present measurement with the disturbance signal from attitude sensors [141]. The mean square error is minimized from the filter.

For estimating the state vectors Kalman filter requires two steps
> Prediction
> Update

Estimating the Process: State is controlled by stochastic equation (linear form)
The process (or) plant represents the input of the system and various state space parameters along with process noise $\left(\mathrm{w}_{\mathrm{k}}\right)$ and measurement equation represents the output of the system along with noises in the sensor measurements $\left(\mathrm{v}_{\mathrm{k}}\right)$. The Process (5.1) and measurements (5.2) equation as given below [142], [143]

Process: $x_{k+1}=\mathrm{A}_{\mathrm{k}} \mathrm{x}_{\mathrm{k}}+\mathrm{Bu}_{\mathrm{k}}+$ Process Noise $\left(\mathrm{w}_{\mathrm{k}}\right)$

Measurement: $\mathrm{Z}_{\mathrm{k}}=\mathrm{H}_{\mathrm{k}} \mathrm{x}_{\mathrm{k}}+$ Measurements Noise ( $\mathrm{v}_{\mathrm{k}}$ )

The process noise and measurement noise both or not related to each other's. The $\mathrm{w}_{\mathrm{k}}$ and $\mathrm{v}_{\mathrm{k}}$ considered at standard probability distribution.
" $\mathrm{A}(\mathrm{nxn})$ " Matrix representing the state at time step k
" B " Matrix representing the requires the input signal to x (state)
" H " Measurement matrix representing to the state $\mathrm{Z}_{\mathrm{k}}$
$\overline{x_{\text {prıorl }}}=$ Priori state estimate k of the process
$\overline{x_{\text {posteriorl }}}=$ Posteriori state estimate k of the measurement
The probability methods predict the future estimate with the help of previous state information and current measurement from the satellite attitude sensor [144]

$$
\begin{equation*}
\overline{x_{\text {posterıorl }}}=\overline{x_{\text {proorl }}}+K\left(Z_{k}-H_{k} \overline{x_{\text {proorl }}}\right) \tag{5.3}
\end{equation*}
$$

" K " is the Kalman Gain,
" $Z_{k}$ " is the Actual Measurements
" $H_{k} \overline{x_{\text {priorl }}}$ " is the Predicted Measurements

The Gain Matrix K is ( n Xm ) is reducing the posteriori error covariance. The system for estimating the process of Kalman filter requires the feedback signal, output feedback into the input of the model. In the model estimating the measurements of feedback signals includes the disturbances (or) errors in the signal [145].

### 5.1.1. Numerical modeling of Kalman filter

$>$ Time update
> Measurement update
Time update methods mainly focus to estimate error covariance and present state information to find previous estimates from next time step [146].

Measurements update method focus to the measurement from the previous signal from feedback to find the best future estimate. Time update methods used to predict response of the system, it is called as "Predictor". Measurement update methods used to correct the response of the system; it is called as "corrector"

## Mathematical calculation using predictor-Corrector method

Time updates find the current estimate from the system [147]. The measurement updates find the future estimate (Shown in Figure 5.2) from the system by actual measurement at that times.


Figure (5.2) Predict/Estimate the Errors in the system [149]

The proposed estimation algorithm considered the state of system and error covariance estimate the step k to step $\mathrm{k}+1$. The measurements update equation to find the Kalman gain, update the measurement with $Z_{k}$ and update the error covariance of the state [148]. The process $\mathbf{P}_{\mathbf{k}}$ and measurement $\mathbf{R}_{\mathbf{k}}$ is and covariance matrix. The measurements updating the state with the help of priori and posteriori state and the Kalman gain to find the actual measurement, $\mathrm{Z}_{\mathrm{k}}$. The process of time-measurement (See Figure-5.3) update steps is repeated with a previous estimate to predict the new estimate [149], [150]. This is called as "recursive nature" of the Kalman filter.


## Figure (5.3) Implementation of Kalman filter algorithm [149]

After the covariance matrix implementations of Kalman filter (KF) used to calculate the past estimates before starting operation of the time - measurement updates. The Attitude sensor data gives the measurement information used to determine variance error. The process covariance is less accurate or less deterministic; it indicates the uncertainty of the process model. For tuning the $P_{k}$ and $R_{k}$ are more important to get the accurate measurements in the satellite/plant till the Kalman gain $\mathrm{K}_{\mathrm{k}}$ stabilize the value [150]. The process covariance $\mathrm{Q}_{\mathrm{k}}$ changes dynamically because of the noisy measurements. So, this value should be adjusted to the differential dynamic equation. The magnitude of process covariant changes the dynamics of the system. The KF predicts the position and angular rates of satellite from attitude sensors (INS/GPS \& IMU) [151], [152].

The magnetometer used to estimate the attitude rates (yaw angle, pitch angle, and roll angle) for the vehicle/satellite. The variation in earth magnetic flux produces the noisy measurements in the plant or model [153]. The state future estimates (5.4) and actual measurements (5.5) as given below

$$
\begin{gather*}
\overrightarrow{x_{k+1}}=\phi_{k} \overrightarrow{x_{k}}+\underline{\Delta}_{k} \overrightarrow{u_{k}}+\overrightarrow{w_{k}}  \tag{5.4}\\
\overrightarrow{Z_{k}}=H \overrightarrow{x_{k}}+\overrightarrow{v_{k}} \tag{5.5}
\end{gather*}
$$

$\mathrm{w}_{\mathrm{k} \text { : Process or plant covariance white noise, } \mathrm{Q}, ~(1)}$
$\mathrm{v}_{\mathrm{k}}$ : Sensor noise covariance, R

The attitude sensor noise (process/measurement) $\left(R_{k} / Q_{k}\right)$ Signals (5.6) \& (5.7) covariance matrix defined by, [153]

$$
\begin{align*}
& R_{k}=\mathrm{E}\left[\begin{array}{ll}
\overrightarrow{v_{k}} & \overrightarrow{v_{k}^{T}}
\end{array}\right]  \tag{5.6}\\
& Q_{k}=\mathrm{E}\left[\begin{array}{ll}
\overrightarrow{w_{k}} & \overrightarrow{w_{k}^{T}}
\end{array}\right] \tag{5.7}
\end{align*}
$$

### 5.1.2. Motion equation of satellite

The satellite accurately measures the attitude rates where it crosses to the line of nodes in the orbit. The motion equations describe the satellite movements of a particular position in the orbits.

The State variable is [41]

$$
\vec{x}=\left[\begin{array}{llllll}
\phi & \dot{\phi} & \theta & \dot{\theta} & \psi & \dot{\psi}
\end{array}\right]^{T}
$$

The satellite dynamic state space equation (SSE)

$$
\frac{\overrightarrow{d x}}{d t}=\mathrm{A} \vec{x}+\mathrm{B} \vec{u}
$$

The acceleration of satellite body (5.8), (5.9) and (5.10) given as
$\ddot{\phi}=\left[\frac{-4 \Omega^{2}\left(I_{y}-I_{z}\right) \phi+\Omega h_{y} \phi-\Omega\left(-I_{x}+I_{y}-I_{z}\right) \dot{\psi}+h_{y} \dot{\psi}-h_{z} \dot{\theta}+I_{x} \dot{\Omega} \psi}{I_{x}}\right]-\left[\frac{\dot{h}_{x}}{I_{x}}\right]+\left[\frac{T_{s p x}+\Omega h_{z}}{I_{x}}\right]$
$\ddot{\theta}=\left[\frac{-3 \Omega^{2}\left(I_{x}-I_{z}\right) \theta-h_{x} \dot{\psi}-\Omega h_{z} \psi-\Omega h_{x} \phi+h_{z} \dot{\phi}}{I_{y}}\right]-\left[\frac{\dot{h}_{y}}{I_{y}}\right]+\left[\frac{T_{s p x}+I_{y} \dot{\Omega}}{I_{y}}\right]$
$\ddot{\psi}=\left[\frac{-\Omega^{2}\left(-I_{x}+I_{y}\right) \psi+\Omega h_{y} \psi-\Omega\left(I_{x}-I_{y}+I_{z}\right) \dot{\phi}+h_{x} \dot{\theta}-h_{y} \dot{\phi}-I_{z} \dot{\Omega} \phi}{I_{z}}\right]-\left[\frac{\dot{h}_{z}}{I_{z}}\right]+\left[\frac{T_{s p z}-\Omega h_{x}}{I_{z}}\right]$
A is process model [41]
Plant matrix, A

$$
\begin{aligned}
& =\left[\begin{array}{lll}
01000
\end{array} 100 \frac{-4 \Omega^{2}\left(I_{y}-I_{z}\right)+\Omega h_{y}}{I_{x}} 00-\frac{h_{z}}{I_{x}} \dot{\Omega} \frac{-\Omega\left(-I_{x}+I_{y}-I_{z}\right)+h_{y}}{I_{x}} ; 000100 ;\right. \\
& -\Omega \frac{h_{y}}{I_{y}}-\frac{h_{z}}{I_{y}}-\frac{3 \Omega^{2}\left(I_{x}-I_{z}\right)}{I_{y}} 0-\Omega \frac{h_{z}}{I_{y}}-\frac{h_{x}}{I_{y}} ; 000001 ;-\dot{\Omega} \frac{-\Omega\left(I_{x}-I_{y}+I_{z}\right)-h_{y}}{I_{z}} 0 \\
& \left.\frac{h_{x}}{I_{z}} \frac{-\Omega^{2}\left(-I_{x}+I_{y}\right)+\Omega h_{y}}{I_{z}} 0\right]
\end{aligned}
$$

Control matrix, $\mathrm{B}=[000 ; 100 ; 000 ; 010 ; 000 ; 001]$
Control to the Actuator of satellite, $\mathrm{u}=-F x+u_{d}$
" F " is the controller Gain
" $u_{d}$ " is the sum of perturbations forces

$$
\begin{equation*}
\frac{\overrightarrow{d x}}{d t}=(\mathrm{A}-B F) x_{o}+\mathrm{B} \overrightarrow{u_{d}} \tag{5.12}
\end{equation*}
$$

This is a satellite equation of motion (5.12) of body. For implementing the Kalman filter, it requires data from on-board attitude sensors (INS/GPS \& IMU) to predict the next state. Refer Annexure B. The satellite attitude control system needs the current attitude data. Using the error estimation KF algorithms used to predicts the future (or) next state [154].

Parameter estimators \& classifications of weights: [155], [156]

Weight for state and estimate parameter

$$
\text { Weight }=W_{0}^{m}=\frac{\text { state }}{\text { measuement }+ \text { state }}
$$

Prepare the filter factor with estimated significance of constraints

$$
\begin{gathered}
\text { State value }\left(\mathrm{t}_{\mathrm{o}}\right)=\text { Expected Value }\left\{\mathrm{x}_{0}\right\} \\
\mathrm{x}_{0}=\text { Initial state parameter }
\end{gathered}
$$

Covariance for the parameters

$$
\mathrm{px}_{0}=\text { Expected Value }\left\{\left(\operatorname{state}\left(\mathrm{t}_{0}\right)-\mathrm{x}_{\mathrm{o}}\right)\left(\text { state }\left(\mathrm{t}_{0}\right)-\mathrm{x}_{\mathrm{o}}\right)^{\mathrm{T}}\right\}
$$

The standard deviations are calculated. The initial conditions simulation considered $\sqrt{\text { covariance matrix }}$ for the estimator of parameter. Equation for the state:

$$
\mathrm{X}_{\text {initial }}=\text { function of } \mathrm{xi}, \mathrm{u}, \mathrm{w}, \text { and } \mathrm{t}
$$

Covariance form for the state:

$$
=\sum_{t=0}^{2 L} W_{i}^{C}(x i-x)(x i-x) T
$$

Measurements (Expected): Output $=\mathrm{h}$ (state)
Mean measurement: $\quad=\sum_{t=0}^{2 L} W_{i}^{M} y i$

The covariance's is [156]

$$
\begin{aligned}
& \mathrm{P}_{\mathrm{yy}}=\mathrm{y}=\sum_{t=0}^{2 L} W_{i}^{C}(y i-y)(y i-y) T \\
& \mathrm{P}_{\mathrm{xy}}=\mathrm{y}=\sum_{t=0}^{2 L} W_{i}^{C}(x i-x)(y i-y) T
\end{aligned}
$$

The Kalman gain $\left(K_{\text {Gain }}\right)=P_{x y} p_{y y}{ }^{-1}$

### 5.1.3. Methodology

In the flow chart shown in Figure (5.4) describes the state flow steps for estimating the errors in attitude sensors (INS/GPS and IMU) using Kalman algorithm. The Kalman algorithm used to calculates the state variables $\left[\begin{array}{llllll}\boldsymbol{\phi} & \dot{\boldsymbol{\phi}} & \boldsymbol{\theta} & \dot{\boldsymbol{\theta}} & \boldsymbol{\psi} & \dot{\boldsymbol{\psi}}\end{array}\right]$. Considered the six-state variables for simulations


Figure (5.4) Flow chart for attitudes error estimation using Kalman Filter

The state update is $\mathrm{X}=$ state + Kalman Gain (K) (Actual measurement $(\mathrm{y})-\dot{y}$ )
Covariance updates $\left(\mathrm{P}_{\mathrm{Cu}}\right):[155] \quad \mathrm{P}_{\mathrm{CU}}=\mathrm{P}-\mathrm{K}_{\text {Gain }} \mathrm{P}_{\mathrm{yy}} \mathrm{K}^{\mathrm{T}}$
' $y$ ' is the actual measurement matches the time for updated state.
The update classification starts by including constraint design random sequence of covariance or uncertainty (Quncer) $\mathrm{P}_{\mathrm{CU}}=\mathrm{P}+\mathrm{Quncer}$

The parameter update: $w=w+K(y-\dot{y})$

Covariance updates $\left(\mathrm{P}_{\mathrm{CU}}\right): \mathrm{P}_{\mathrm{CU}}=\mathrm{P}-\mathrm{K}_{\text {Gain }} \mathrm{P}_{\mathrm{yy}} \mathrm{KT}$
' $y$ ' is the actual measurement matches the time for updated state.

Hence, nonlinear state equations are $\dot{r}=\mathrm{v}$

$$
\dot{v}+\frac{\text { Gravitational Parameter } * r}{\left(r^{T} r\right) 3 / 2}+\sum_{K} a k=\frac{f}{m} \quad \text { [156] }
$$

The satellite mass is ' $m$ ' position vector is ' $r$ ' solar pressure force on the satellite is ' $f$ '. In Figure (5.5) Shows the attitude error estimation using KF.

$$
u_{d} \text { (Disturbance (3X1) Matrix) }
$$



Controller Gain (3X6) Matrix
Figure (5.5) Block diagram of Kalman filter state estimation

Where, A = Process or Plant (6X6) Matrix
$B=$ Control (6X3) Matrix
$\mathrm{H}=$ Measurement (3X6) Matrix
$x=$ State (6X1) Matrix, $=\left[\begin{array}{llllll}\phi & \dot{\phi} & \theta & \dot{\theta} & \psi & \dot{\psi}\end{array}\right]^{T}$

F= PD Controller (3X6) Matrix [136]

State space equation [157]
State, $\dot{x}=A x+B u$

Output, $y=H x$

Control Input, $u=-F x+u_{d}$


Figure (5.6) Roll attitude dynamics of NPSAT- 1 to consider the principle moment of inertia ( $I_{x}=24.67, I_{y}=22.63, I_{z}=11$ ) kg-m ${ }^{2}$ the actual attitude data
(0-5000 Seconds) measured from the on-board attitude sensor IMU and Magnetometer with reference to low earth orbit trajectory

The Kalman filter algorithms used to estimate the errors in the real time raw data (Roll angle, Pitch angle, Yaw angle) measured from on-board attitude sensor. The KF simulation considered the step time $0-5000$ second attitude data from reference low earth orbiting satellite in the entire orbit [158]. Refers Appendix-D, for
implementing the Kalman filter algorithm, it is necessary to consider the known constants. In the simulation considered total disturbance torque $\left(\mathrm{T}_{\mathrm{d}}\right)$ is $1.04 \times 10^{-4}$ N.m. The disturbance torques due to atmospherics parameters is aerodynamics perturbation, solar radiation pressure [32]. The Kalman filter simulation employed in one entire orbit period, the time taken to complete one cycle. The satellite it's starting from one node with respect to the equator and complete with same node [24]. The one orbit simulation is very useful to estimate the changes in orbit from normal orbit to perturbed/disturbed orbit. In figure (5.6) shows the NPSAT-1 actual roll data.


Figure (5.7) Roll attitude dynamics of NPSAT-1 to consider the principle moment of inertia ( $I_{x}=24.67, I_{y}=22.63, I_{z}=11$ ) kg-m ${ }^{2}$ the Predicted/Estimated attitude data (0-5000 Seconds) using a Kalman filter algorithm

The initial simulation parameters considered is angular momentum of momentum wheel, $\mathrm{h}=10$ (Nms.), angular velocity, $\omega=0.0011068(\mathrm{rad} / \mathrm{s})$, Disturbance torque $1.04 \times 10^{-4}(\mathrm{Nm})$, step time duration, $\mathrm{dt}=0.1$ (seconds) [158]. The attitude data from on-board sensors in the IMU and Magnetometer size is referred in the code is mag = magneto', $[\mathrm{m} \mathrm{n}]=\operatorname{size}(\mathrm{mag})$; ' m ' is No of raw in the data; ' n ' is No of column in data attached in the report. The initial state matrix $\boldsymbol{X}_{\boldsymbol{0}}$ is 6 X 1 , represents
six states (roll angles, roll rates, pitch angles, pitch rates, yaw angles, yaw rates). For considered the suitable assumptions for estimate the accurate measurements with less error in the system. In figure (5.7) shows the NPSAT-1 estimated data with K.

State vector: $\vec{x}=\left[\begin{array}{llllll}\phi & \dot{\phi} & \theta & \dot{\theta} & \psi & \dot{\psi}\end{array}\right]^{T}$
To initialize the Kalman filter, for Initial error covariance matrix, 6X6 all the diagonal elements considered as $10^{-6}$, process noise covariance matrix, Q is $\mathrm{e}^{-6}$, measurement noise covariance matrix, R is $\mathrm{e}^{-2}$. [157]

Table (5.1) Estimation of NPSAT-1 satellite attitude errors between actual and estimated attitudes, principle MOI $\left(\mathrm{I}_{\mathrm{x}}=24.67, \mathrm{I}_{\mathrm{y}}=22.63, \mathrm{I}_{\mathrm{z}}=11\right) \mathbf{k g - m}{ }^{2}$

| Satellite <br> Attitudes <br> (deg) | Roll <br> angle <br> (deg) | Roll_est <br> (deg) |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $(\mathrm{Sec})$ |  |  | c

In table (5.1) shows the variations in attitudes with reference of low earth orbit satellite data collected from the initial position to next 10 Seconds. The actual attitude information is measured by IMU and Magnetometers. It clearly understood the roll angle slightly oscillates from 52.25 degrees to 53.46 degrees as shown in the figure (5.5). The pitch angles nearly constant ( 0 to 5000 Sec ) in the entire orbit [157]. The yaw angles oscillate from 185.21 degrees to 185.94 degrees. The NPSAT-1Kalman algorithm estimates the attitude error deviations in the orbit.


Figure (5.8) Pitch/Yaw attitude dynamics of NPSAT-1 to consider the principle moment of inertia ( $I_{x}=24.67, I_{y}=22.63, I_{z}=11$ ) kg-m ${ }^{2}$ the (a), (c) is Actual attitudes (pitch/yaw) data (b), (d) is Predicted/Estimated attitude data (0-5000

## Seconds) using a Kalman filter algorithms

In the Figure (5.8) shows the numerical simulation of NPSAT-1 pitch/yaw attitude dynamics compare with actual attitudes and predicted/estimated attitudes. The measurement noise covariance matrix $\mathrm{R}, \mathrm{e}^{-2}$ used to minimize the errors in the system [158]. The Kalman simulation generates the error covariance matrix as shown in below. The Kalman filter simulation used to estimates the State vectors

Table (5.2) NPSAT-1 State vectors measurements from IMU and Magnetometers

| Time in <br> Sec | $\boldsymbol{\phi} \boldsymbol{d e g}$ | $\dot{\boldsymbol{\phi} \text { deg }}$ | $\boldsymbol{\theta} \boldsymbol{\operatorname { d e g }}$ | $\dot{\boldsymbol{\theta} \boldsymbol{d e g}}$ | $\boldsymbol{\psi} \boldsymbol{d e g}$ | $\dot{\boldsymbol{\psi} \boldsymbol{d e g}}$ |
| :---: | ---: | ---: | ---: | :---: | :---: | :---: |
| 1 | 0.010576 | 0.007941 | 0.00034 | 0.00051 | 0.03723 | -0.0057 |
| 2 | 0.013368 | 0.013435 | 0.00746 | -0.0111 | 0.0605 | -0.0042 |
| 3 | 0.021196 | 0.026652 | -0.0072 | -0.0224 | 0.07906 | -0.0265 |
| 4 | 0.031255 | 0.024125 | -0.0154 | -0.0198 | 0.06837 | -0.0428 |
| 5 | 0.028917 | 0.021914 | -0.0132 | -0.0175 | 0.06663 | -0.0365 |
| 6 | 0.026345 | 0.02574 | -0.0106 | -0.0217 | 0.07773 | -0.0322 |
| 7 | 0.028964 | 0.03255 | -0.0137 | -0.0287 | 0.09133 | -0.0385 |
| 8 | 0.036382 | 0.034036 | -0.0214 | -0.0299 | 0.08867 | -0.0529 |
| 9 | 0.0368 | 0.02928 | -0.0216 | -0.0249 | 0.07863 | -0.0521 |
| 10 | 0.032376 | 0.029406 | -0.017 | -0.0252 | 0.08289 | -0.0431 |

Error covariance matrix: (Calculated from Kalman Simulations)

| $3.35 \mathrm{E}-06$ | $1.86 \mathrm{E}-07$ | $-1.21 \mathrm{E}-06$ | $-6.78 \mathrm{E}-07$ | $-8.23 \mathrm{E}-07$ | $-3.24 \mathrm{E}-06$ |
| ---: | :---: | ---: | ---: | ---: | ---: |
| $1.86 \mathrm{E}-07$ | $2.36 \mathrm{E}-06$ | $-7.34 \mathrm{E}-07$ | $-1.21 \mathrm{E}-06$ | $2.56 \mathrm{E}-06$ | $-8.23 \mathrm{E}-07$ |
| $-1.21 \mathrm{E}-06$ | $-7.34 \mathrm{E}-07$ | $3.51 \mathrm{E}-06$ | $1.94 \mathrm{E}-07$ | $7.77 \mathrm{E}-07$ | $3.39 \mathrm{E}-06$ |
| $-6.78 \mathrm{E}-07$ | $-1.21 \mathrm{E}-06$ | $1.94 \mathrm{E}-07$ | $2.51 \mathrm{E}-06$ | $-2.75 \mathrm{E}-06$ | $7.77 \mathrm{E}-07$ |
| $-8.23 \mathrm{E}-07$ | $2.56 \mathrm{E}-06$ | $7.77 \mathrm{E}-07$ | $-2.75 \mathrm{E}-06$ | $8.04 \mathrm{E}-06$ | $1.49 \mathrm{E}-06$ |
| $-3.24 \mathrm{E}-06$ | $-8.23 \mathrm{E}-07$ | $3.39 \mathrm{E}-06$ | $7.77 \mathrm{E}-07$ | $1.49 \mathrm{E}-06$ | $7.04 \mathrm{E}-06$ |
| isturbance matrix): |  |  |  |  |  |

0.000452857721929469
0.0110725956694653
$-0.000996727272727273$
The results of NPSAT-1 Kalman numerical simulation produce error covariance matrix and disturbances matrix considered the step time $0-5000$ second attitude data from reference low earth orbiting satellite in the entire orbit [159]. Refers AppendixD, For Implementing the Kalman filter algorithm, it is necessary to consider the known constants.



Figure (5.9) Estimates the attitude errors from on-board sensor INS/GPS (a)
Reference latitude of INS and GPS (b) Reference longitude of INS and GPS (c) KF estimated the latitude errors (d) KF estimated the longitude errors (e), (f), (g)

KF estimated the errors in INS /GPS in Navigation frame respectively $V_{n}, V_{e}, V_{d}$
(h) Integrated INS/GPS in reference longitude (i) Integrated INS/GPS in reference latitude

Table (5.2) clearly indicates the oscillation in roll angles and yaw angles in the orbit. The pitch angle is constant in the entire orbit. The state vectors $\overrightarrow{\boldsymbol{x}}=$ $\left[\begin{array}{llllll}\boldsymbol{\phi} & \dot{\boldsymbol{\phi}} & \boldsymbol{\theta} & \dot{\boldsymbol{\theta}} & \boldsymbol{\psi} & \dot{\boldsymbol{\psi}}\end{array}\right]^{\boldsymbol{T}}$ considered the moment of inertia $\mathrm{I}_{\mathrm{x}}=24.67, \mathrm{I}_{\mathrm{y}}=22.63, \mathrm{I}_{\mathrm{z}}=$

11 [158]. The actual attitude information is measured by IMU and Magnetometers [140]. In Figure (5.9) shows the NPSAT-1 estimates the attitude errors simulation results of integration of INS/ GPS in the satellite model [160], [161].


Figure (5.10) Kalman filter estimated the attitude actual errors from INS and Kalman filter errors (a) Attitude yaw errors (b) Attitude pitch errors (c) Attitude roll errors (d) Integrated INS/GPS with reference North Velocity

In Figure（5．10）shows the Kalman filter simulation of INS／GPS attitude data attitudes （Yaw，Pitch，Roll）data in $\mathrm{x}, \mathrm{y}, \mathrm{z}$ direction in navigation frames measured by an attitude sensor in reference with low earth orbit trajectory．The simulation time consider in the model is $0-100$ Seconds［162］．

|  |  | 20 |  | $\begin{aligned} & \text { 苞 } \\ & \text { in } \end{aligned}$ | $\begin{aligned} & \text { 喜 } \\ & \mathbf{i} \end{aligned}$ | 登 | Sio | 龠 | soog io | \％ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
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|  |  |  | 尔 | 尔 | 年 |  | 年 | 夺 | 桨 | ＇ |
|  |  |  | 0 | O | O్రిరి | $\begin{array}{\|l\|l\|} \hline 0 \\ \hline 0.0 \\ \hline \end{array}$ | 苟 | Bo | $\begin{array}{l\|l} 0.00 \\ \hline \end{array}$ | 答 |
|  | E | O | \％ | 蒿 | 蒿 | 蒿 | 픙 | $\begin{aligned} & \text { By } \\ & \substack{\text { B } \\ \hline} \end{aligned}$ | － | － |
|  | 2 | O | O్ర్ర心 | Br | 蒿 | Bo | Sis | B | 㾑 | － |
|  | 5 |  | 䓂 | 欨 | 蒿 | Sis | $\begin{aligned} & \hline 8.8 \\ & \hline \mathbf{O} \\ & \hline \end{aligned}$ |  | 䓂 | 蓇 |
|  |  | :8ib | 登 | Sis | $\begin{array}{\|l\|} \hline \text { 营 } \\ 0 \end{array}$ |  | Sidy |  | $\begin{array}{\|l} \text { in } \\ \stackrel{0}{\circ} \\ 0.0 \\ \hline \end{array}$ | \％ |
|  |  |  | 응 | Cosi | 临 | $\begin{aligned} & \stackrel{\rightharpoonup}{3} \\ & 0 \\ & \hline \end{aligned}$ | $\stackrel{\rightharpoonup}{3}$ | 蒿 | B | 会 |
|  |  | ¢ |  | \％ | 答 | $\underset{0}{3}$ | Siog | \％ | 产 | 웅 |



The reference longitude data of GPS are $45^{\circ}$ and a reference longitude of INS is measured from $45.45^{0}$ in the actual orbit．

The simulation considered the step time $0-100$ second attitude data from (GPS/INS) reference at low earth orbiting satellite in the entire orbit. Refers Appendix-D, For Implementing the Kalman filter algorithm, it is necessary to consider the known constants [163], [164]. The measurement noise covariance matrix $R$, $e^{-2}$ used to minimize the errors in the system. The minimum covariance values desired the less error in the measurements due to the perturbations [165]. The table (5.3) indicates the minimum errors in GPS, and INS with implementation of Kalman algorithms [141], [160]. The attitude data of GPS/INS considered the LEO trajectory. This result mainly predicts the future estimate used to control the attitude orientation/control of the satellite.

## CHAPTER 6

## SATELLITE ATTITUDE CONTROLLER

### 6.1 Overview

The Attitude Control System (ACS) is important to maintain the satellite into prescribed/determined orbit from perturbed/disturbed orbit. The attitude (Roll angle, Pitch angle, and Yaw angle) of satellite changes due to the orbital perturbation. It is mandatory to control \& correct the attitudes of the satellite into the actual orbit [166], [167]. This chapter presents the design of the attitude controller for NANO satellites as given below

- NPSAT-1 Satellite,
- SRM Satellite
- Pratham (IIT Bombay) Satellite

The controller is also used to reduce the oscillation due to the perturbation forces where affects the attitude of the satellite. Also, it is decreasing the errors in the satellite system dynamics [168]. The effects of satellite dynamics without controller and with controllers are compared.


Figure (6.1) Satellite Attitude Control System [169]
The design of Proportional - Derivative (PD) controller has been introduced for various transient responses of the NANO satellite attitude corrections. In Figure (6.1) shows the satellite attitude control system, the actual attitudes measure by Gyros which can be compare with reference attitudes. The comparator produces the error
signals E(s) fed to the controllers. The PD controllers introduced in satellite control system. The steady-state is settled exactly at 1 (zero steady-state error). Hence there is no need for integral control. The PD controller generates the Armature voltage ( $\mathbf{V}_{\mathbf{a}}$ ) supply the DC motor to control armature current ( $\mathbf{I}_{\mathbf{a}}$ ) in the input circuit. The DC motor act as an actuator, DC motor output is Torque (T). The armature coil consists of shaft used to derive the load. The Attitudes determination and control system (ADCS) the satellite inertia considered as load [168]. In the input circuit, due to variation in the armature voltage, the output circuit satellite inertia (Load) induces the angular velocity ( $\boldsymbol{\omega}$ ). This angular velocity generates because of the mutual inductance in the secondary circuit. In secondary coil induces the back-e. m. f (b) due to the changes in the current in the primary coil. The torque is inversely proportional to armature current shown in equation (6.1)

$$
\begin{equation*}
\mathbf{T}=\mathbf{K} \mathbf{I}_{\mathbf{a}} \tag{6.1}
\end{equation*}
$$

$\mathrm{K}=$ Torque constant;
$\mathbf{T} \propto \mathbf{I}_{\mathrm{a}}$


Figure (6.2) Armature controller DC motor [168]
In figure (6.2) shows the armature-controlled DC motor, $\mathbf{R}_{\mathbf{a}}$ is armature resistance, $\mathbf{L}_{\mathbf{a}}$ is armature inductance, $\mathbf{V}_{\mathbf{b}}$ is back e. m. f voltage, $\mathbf{J}$ is the moment of inertia, $\mathbf{b}$ is the back e. m. f constant, If is field current [168]. The magnitude of back e. m. f voltage is opposite to the input voltage.

$$
\begin{equation*}
\mathbf{V}_{\mathbf{b}}=\mathrm{k} \omega \tag{6.2}
\end{equation*}
$$

Hence, the equation (6.2) back e. m. f voltage is inversely proportional to the angular velocity derive the satellite inertia. The back e. m. f is expressed as $\mathrm{V} /(\mathrm{Radian} / \mathrm{Sec})$. In the proposed design considered back e. m. f. constant is 0.85 . In ACS, $\mathrm{V}_{\mathrm{b}}=0.85 \mathrm{~V}$ generate the angular velocity $\omega$ (1 Radian/Seconds). The one Ampere current (I) produce the torque $0.85(\mathrm{Nm})$. Since, the torque is proportional to the armature current. This is highly used to control the attitudes of Nano satellite [169]

The transfer function of armature-controlled DC motor is given in equation (6.3)

$$
\begin{equation*}
\frac{T(s)}{V_{a}(s)}=\frac{K}{L_{a} s+R_{a}} \tag{6.3}
\end{equation*}
$$

Table (6.1) ACS DC Motor design Parameters [168]

| Torque constant, K | 0.85 |
| :---: | :---: |
| Armature Inductance, $L_{a}$ | 0.003 H |
| Armature Resistance, $R_{a}$ | $1 \mathrm{Ohms}(\Omega)$ |
| Back e. m. f. constant, b | 0.85 |

In table (6.1) shows the ACS design parameters, the sensitive of Gyro is $(\mathrm{V} /$ (deg/Sec). For, measure the actual attitude from satellite the Gyro gain chosen as one. (Gyro Gain =1). The second order charactertics equations consider for satellite attitude dynamics. The various transient response of the NANO satellite time domain signal reached the maximum value Rise time ( $\mathrm{t}_{\mathrm{r}}$ ), the time at maximum amplitude occurs peak time $\left(\mathrm{t}_{\mathrm{p}}\right)$, the ratio of oscillation of critical damping to actual damping Maximum overshoot (\%MP), at the time the signal reached the final value Settling time $\left(\mathrm{t}_{\mathrm{s}}\right)$, the value of percentage of errors in the system Steady state errors ( $\mathrm{e}_{\mathrm{ss}}$ ) [168]. This control torques used to stabilize the orientation of a satellite into actual path in orbit [169].

## Discussed Control System Design Response Specifications:

- Settling time $\leq 0.2$ seconds
- Improve the Peak time, Rise time
- Reduces the overshoot (Damping)
- Zero steady state error


### 6.2. Methodology

The proposed method used to design the Nano satellite attitude control transient response of using PD controllers.


Figure (6.3) Flow chart for satellite attitude controller design
The Proportional-Derivative (PD) controllers have been introduced in Nano satellite/plant attitude control. In flow chart shows (See Figure 6.3), when design a PD controller to improve the transient response (Rise time, Overshoot, Peak time and Settling time). Theoretically, the damped frequency of oscillation is $\omega_{d}=\omega_{n} \sqrt{1-\xi^{2}}$, $\omega_{n}$ is the natural frequency of oscillation (rad/seconds) [166]. The plant is referred as satellite dynamics. The PD controller, proportional Gain $\mathbf{K}_{\mathbf{p}}$-Errors proportional to the
system, Derivative Gain $\mathbf{K}_{\mathbf{d}}$ - Derivative of errors in the system [168]. The design of satellite control systems consists of two parts one is Attitude determination (AD) another Attitude control (AC). The satellite subsystem used to determine the attitude, predict the future attitude, and control the attitude of satellite [167]. This chapter includes the attitude control part of the NANO spacecraft/satellite attitude corrections [170]. The main part of ADCS is actuator and spacecraft dynamics, controllers and attitude sensors. The satellite attitude sensor such as Gyro's used to determine the actual attitudes. The satellite reference input signal compares with feedback signal from attitude sensor. If there is any variation in the feedback signal with respect to reference/pre-determined attitude, then initiate the actuator to generate the control torque/forces. The damping ratio desired the types of oscillation. The equations $-\xi \omega_{n} \pm j \omega_{d}$ first part indicates the real value; second part indicates the imaginary value. As per the design specification the settling time less than 0.2 seconds is calculated by $\xi \omega_{n}=\frac{4}{0.2}=\mathbf{2 0}[166]$.

$$
\tan (34.3)=\frac{\omega_{d}}{20} \Rightarrow \omega_{d}=\mathbf{1 3 . 6 4}
$$



Figure (6.4) Location of closed dominant pole in the S-Plane
To determine the imaginary part value it require to find the angles from closed dominant pole, draw the line from closed dominant pole to origin of the s-plane (or) frequency plane, $\theta=\cos ^{-1}(\xi)=\cos ^{-1}(0.826)=\mathbf{3 4 . 3}^{\circ}$ as shown in Figure 6.4 (Here, for design point of view, the overshoot percent is taken as $\mathbf{1 \%}$ (equal to a damping ratio of $\mathbf{0 . 8 2 6}$ ) [166]. As per the design specification of Nano satellites (NPSAT-1, SRM Satellite, Pratham Satellite) is settling time of 0.2 seconds.

### 6.3 Algorithms -Various steps to follows design of Attitude controller

STEP 1. Defining the desired control system specification (transient response and steady-state error).

STEP 2. Finding the dominant pole for the desired specifications.
STEP 3. Design the compensator/controller pole and/or zero.
STEP 4. Calculates the loop gain of the overall control system.
STEP 5. Verify the response through simulation.
An algorithm of Nano satellite (NPSAT-1, Pratham Satellite, and SRM Satellite) attitude control implemented using MATLAB Tools. Also, the closed loop poles help to find the gain of the system using Root Locus (RL) methods. In design simulation considered the satellite attitude control, Satellite Attitude determination (SAD), Satellite Attitude Prediction (SAP), Satellite Attitude Control (SAC). The SAD is the process of computing the orientation of satellite with pre-determined point accuracy from on board sensors [171]. The SAP is the process of estimating the future attitude of the satellite model. The SAC is the process of controlling the orientation of the satellite. In table (6.2) shows the Nano satellites design parameter considered for simulations [134], [136], and [172]

Table (6.2) Nano satellite attitude control design parameters

|  | Satellite Details |  |  |
| :---: | :---: | :---: | :---: |
| Altitude | NPSAT-1 | Pratham Satellite | SRM Satellite |
| 550 km altitude, <br> Low Earth circular <br> orbit (LEO) | 817 km altitude, <br> Polar Sun <br> Synchronous Orbit <br> (PSSO) | 867 km altitude, <br> Low Earth circular <br> orbit (LEO) |  |
| Orbital <br> angular <br> velocity, $\boldsymbol{\Omega}$ | $0.0011068 \mathrm{rad} / \mathrm{s}$ | $0.0010346 \mathrm{rad} / \mathrm{s}$ | $0.0010239 \mathrm{rad} / \mathrm{s}$ |
| $\boldsymbol{I}_{x x}$ | $24.67 \mathrm{~kg}-\mathrm{m}^{2}$ | $0.116 \mathrm{kg-m} \mathrm{~m}^{2}$ | $6.2911 \mathrm{~kg}-\mathrm{m}^{2}$ |
| $\boldsymbol{I}_{y y}$ | $22.63 \mathrm{~kg}-\mathrm{m}^{2}$ | $0.109 \mathrm{~kg}-\mathrm{m}^{2}$ | $5.9162 \mathrm{~kg}-\mathrm{m}^{2}$ |
| $\boldsymbol{I}_{z z}$ | $11 \mathrm{~kg}-\mathrm{m}^{2}$ | $0.114 \mathrm{~kg}-\mathrm{m}^{2}$ | $4.6085 \mathrm{~kg}-\mathrm{m}^{2}$ |
| Three attitude axes (Roll, Pitch, Yaw) are decoupled |  |  |  |

### 6.4. Attitude Control: NPS Aurora Satellite (NPSAT-1) [172]

The principle moment of inertia of NPSAT-1 satellite is $[24.67,22.63,11] \mathrm{kg}-\mathrm{m}^{2}$. The angular velocity of the satellite $0.0011068 \mathrm{rad} / \mathrm{s}$ at 550 km Altitude [172]. The closed loop response of roll attitude dynamics is $\frac{\phi(s)}{T_{x}(s)}$. The input of the model is torque and output is actual attitudes (Roll angle, Pitch angle, and Yaw angle) of satellite [173]. The Gyro gain is taken as one. The sensitive of Gyro is (V/(deg/Sec). For getting the actual attitudes from model, this feedback signal compares with predetermined attitudes. The errors signal is generated from comparators fed to Armature control DC motor. The DC motor act as an actuator.

### 6.4.1. NPSAT-1 Roll attitude control system

The comparison of reference roll and feedback signals measured from Rate Gyro's (RG) produces (Ref. from figure 6.1) the errors in the system. These errors are minimized from controllers and generate the control torque to the satellite dynamics. Principal moments of inertia of NPSAT-1 [172]

$$
=\left[\begin{array}{ccc}
24.67 & 0 & 0 \\
0 & 22.63 & 0 \\
0 & 0 & 11
\end{array}\right] \mathrm{kg}-\mathrm{m}^{2}
$$

As per the design specification of NPSAT- 1 is settling time is $\leq 0.2$ seconds considered the closed dominant pole is $\mathbf{- 2 0} \mathbf{\pm 1 3 . 6 4 j}$. The error signal expressed in e $(\mathrm{t})$, Satellite input/reference roll signal ( $\phi_{\text {ref }}$ ) expressed in degree. The output of the satellite system is $\phi(\mathrm{t})$.


Figure (6.5) NPSAT-1 Roll attitude control system

The step response Figure (6.5) of the roll attitude transfer function without controller given below.


Figure (6.6) NPSAT-1 Roll attitude response with step command
From Figure 6.6, it is clear the damping of the roll attitude system overshoot is $55.1 \%$. The oscillation gradually reduces with settling down at time 108 seconds (we need to reduce it to $\leq 0.2$ seconds). The system doesn't have steady state transient response errors so, not requires the Integral (I) controller [166]. The NPSAT-1 roll attitude dynamics clearly indicates, there are three poles in the dynamics one on the origin of the s-plane another is -0.0639 and -333.264 illustrate in Figure (6.7) finding the angles $\theta_{1}, \theta_{2}$ and $\theta_{3}$ from closed dominant pole to other poles and zeros [167]. There are no zeros in the pole-zero plots. This method used to design the PD compensated controller by using root locus analysis find the zero location. For, finding the PD controller zero $\left(z_{c}\right)$, The line connecting from complex pole to pole at the origin and pole at -0.0639 and -333.264 (Refers Appendix E). Let us assume complex pole considered as ' $\mathbf{A}$ ' [166]. The angles from complex pole (Equation 6.4) $\mathrm{A}=180^{\circ}-$ (summation of angle measure from complex pole to other poles) + (summation of angle measure from complex pole to other zeros)

$$
\begin{equation*}
=180-\left(\theta_{1}+\theta_{2}+\theta_{3}\right) \tag{6.4}
\end{equation*}
$$



Figure (6.7) PD compensated design (Complex pole to others pole)
The angles from closed dominant pole to others pole are
$\theta_{1}=180^{\circ}-\left[\tan ^{-1}\left[\frac{13.64}{20}\right]\right]=145.7^{\circ}$
$\theta_{2}=180^{\circ}-\left[\tan ^{-1}\left[\frac{13.64}{20-0.0639}\right]\right]=145.614^{\circ}$
$\theta_{3}=\left[\tan ^{-1}\left[\frac{13.64}{-333.264-20}\right]\right]=2.493^{\circ}$
Substitute the angles from closed dominant pole to others poles $\theta_{1}, \theta_{2}$, and $\theta_{3}$ in equation (6.4) [167]
$=180-(145.7+145.614+2.493)=-\mathbf{1 1 3 . 8 0 7}^{\circ}$
Now, find the PD compensated zero from Figure (6.8) from angle measured from complex pole A (-113.807 ${ }^{\mathbf{}}$ )


Figure (6.8) Finding the location of the attitude PD controller
Using the geometry shown in Figure 6.8, $\tan \left(180^{0}-113.807^{\circ}\right)=\frac{13.64}{20-z_{c}}$
To design the PD compensator zero $\left(z_{c}\right)$ from above geometry calculates the value 13.98. Now, the PD controller dynamics is $K(s+13.98)$. The $K$ is the loop gain [166]. Since the system have three poles, so there are three root locus, one travel from origin travel to negative real axis, another two-root locus start at $-159 \pm 108 i$ (Break in point) then, travel perpendicular to real axis. In figure (6.9) shows the NPSAT-1 Roll Root Locus response indicates gain is 3700 at overshoot ( $1 \%$ ), the damping ratio, $\boldsymbol{\xi}$ is 0.826 , undamped natural frequency, $\boldsymbol{\omega}_{\boldsymbol{n}}$ is $192(\mathrm{rad} / \mathrm{Sec})$. Here, for design point of view considered one percent overshoot ( $1 \%$ ) a damping ratio equal to $\mathbf{0 . 8 2 6}$. To Calculates the angle from given equation (6.5)
$\theta=\cos ^{-1}(\xi)=\cos ^{-1}(0.826)=\mathbf{3 4 . 3}{ }^{\circ}$
Draw the straight line at angle $34.3^{0}$ from origin left of the s-plane, when this line crossing to RL at pole $-159 \pm 108$ i, Record the loop Gain as per damping ration 0.826 .


Figure (6.9) NPSAT-1 Roll attitude Root locus response
The MATLAB Simulation for closed loop system includes the DC motor dynamics, satellite roll attitude dynamics, and PD compensator. (Refers Appendix E)


Figure (6.10) NPSAT-1 Roll Attitude Response with PD controller

In Figure 6.10 shows, as per the design specification considered system settling time is ( $\leq 0.2$ ), the system responses after implemented the PD compensated controller is
0.129 Sec , rise time is 0.0106 , overshoot is improved $11.2 \%$, and peak time is 0.025 with 1.11 peak amplitude were achieved. The NPSAT-1 SIMULINK responses of roll attitude control diagram shown in Figure 6.11. This model includes the dynamics of armature-controlled DC motor cascade with spacecraft inertia considered the back e. m. f. constant 0.85 . To introduce the PD controllers $\boldsymbol{K}(\mathbf{s}+\mathbf{1 3 . 9 8})$, K is 3700 loop gain calculating from Root locus analysis. The input step signal considered for satellite attitude control model, the variation in transient response output is measured from scope block. In figure (6.11) shows the NPSAT-1 Roll response without controller, it is noticed peak amplitude occurs at 18 seconds peak time [174], [175]. Also, the response settles down at settling time 108 seconds with zero steady state errors



The NPSAT-1 Roll control PD-compensated response shown in Figure (6.12). As per the design consideration settling time is $\leq 0.2$. The derivative block is introduced for 'D' controller [176], [177]. Thus, the PD-compensated transient response is improved, the same as recorded from MATLAB simulations.

### 6.4.2. NPSAT-1 Pitch attitude control system

The closed loop response of pitch attitude dynamics is $\frac{\theta(s)}{T_{y}(s)}$.


Figure (6.13) NPSAT-1 Pitch Attitude Control
The step response shown in figure (6.13) of pitch attitude transfer function without controller is given below. (Refers Appendix E)


Figure (6.14) NPSAT-1 Pitch attitude response with step command (without controller)
From Figure 6.14, it is clear the damping of the roll attitude system has more overshoot is $53.2 \%$. The oscillation gradually reduces with settling down at time 102 seconds (we need to reduce it to $\leq 0.2$ seconds) [176]. And, the steady-state is settled exactly at 1 (zero steady-state error). Hence there is no need for integral control. The

NPSAT-1 roll attitude dynamics clearly indicates, there are three poles in the dynamics one on the origin of the s-plane another is -0.0761 and -333.2572 shown in Figure (6.15) finding the angles $\theta_{1}, \theta_{2}$ and $\theta_{3}$ from closed dominant pole to other poles and zeros [166]. There are no zeros in the pole-zero plots. This method used to design the PD compensated controller by using root locus analysis


Figure (6.15) PD compensated design (Complex pole to others pole)

The angles from closed dominant pole to others pole are
$\theta_{1}=180^{\circ}-\left[\tan ^{-1}\left[\frac{13.64}{20}\right]\right]=145.7^{\circ}$
$\theta_{2}=180^{\circ}-\left[\tan ^{-1}\left[\frac{13.64}{20-0.0761}\right]\right]=145.6^{\circ}$
$\theta_{3}=\left[\tan ^{-1}\left[\frac{13.64}{-333.2572-20}\right]\right]=2.493^{0}$
This angle is used to find the zero location for finding the PD controller zero $\left(z_{c}\right)$, The line connecting from complex pole to others pole at the origin and pole at -0.0761 and -333.2572 . Let us assume complex pole considered as A .

The angles from complex pole A [153]
$=180^{\circ}-$ (summation of angle measure from complex pole to other poles) + (summation of angle measure from complex pole to other zeros)

$$
=180-\left(\theta_{1}+\theta_{2}+\theta_{3}\right)=180-(145.7+145.6+2.493) \approx-\mathbf{1 1 3 . 7 9 3}{ }^{\circ}
$$

Now, find the PD compensated zero from Figure (6.16) from angle measured from complex pole A (-113.793 ${ }^{0}$ ) [166]


Figure (6.16) Finding the location of the attitude PD controller (zc)
Using the geometry shown in Figure 6.16, $\tan \left(180^{\circ}-113.793^{\circ}\right)=\frac{13.64}{20-z_{c}}$
To design the PD compensator zero ( $z_{c}$, from above geometry calculates the value 13.986. Now, the PD controller dynamics is $K(\mathbf{s}+13.986)$. The $K$ is the loop gain [167]. The loop gain found from the root locus at $1 \%$ maximum overshoot is 3350 , damping ratio, $\boldsymbol{\xi}$ is 0.826 , undamped natural frequency, $\boldsymbol{\omega}_{\boldsymbol{n}}$ is $192(\mathrm{rad} / \mathrm{Sec})$ shown in figure (6.17)


Figure (6.17) NPSAT-1 Pitch Attitude Root locus response
The MATLAB Simulation for closed loop system includes the DC motor dynamics, satellite inertia, and PD compensator in feed forward loop. (Refers Appendix E)


Figure (6.18) NPSAT-1 Pitch Attitude Response with PD controller

In Figure 6.18 shows the NPSAT-1 pitch response, as per design specifications Settling time 0.129 Sec, Rise time 0.0106 Sec, Overshoot (\%Mp) $11.2 \%$, and peak time 0.025 Sec . at peak amplitude 1.11 were achieved.

### 6.4.3. NPSAT-1 Yaw attitude control system

The closed loop response of yaw attitude dynamics is $\frac{\psi(s)}{T_{Z}(s)}$.


Figure (6.19) Dynamics of Yaw attitude control system
In Figure (6.19) shows the NPSAT-1 Yaw attitude step response of the transfer function without controller given below. (Refers Appendix E)


Figure (6.20) NPSAT-1 Yaw attitude dynamics response with step command (without controller)

From Figure 6.20, it is clear the damping of the yaw attitude system has more overshoot is $40 \%$. The oscillation gradually reduces with settling down at time 49.9 seconds (we need to reduce it to $\leq 0.2$ seconds) [176]. And, the steady-state is settled exactly at 1 (zero steady-state error). Hence there is no need for integral control. As
per the design specification of NPSAT-1 is settling time is $\leq 0.2$ seconds at closed dominant pole occurs at $\mathbf{2 0} \pm \mathbf{1 3 . 6 4} \mathbf{j}$.


Figure (6.21) PD compensated design (Complex pole to others pole)
The NPSAT-1 Yaw attitude dynamics clearly indicates, there are three poles (See Figure 6.21) in the dynamics one on the origin of the s-plane another is -0.1551 and 333.1783. To finding the angles $\theta_{1}, \theta_{2}$ and $\theta_{3}$ from closed dominant pole to other poles and zeros. There are no zeros in the pole-zero plots. This method used to design the PD compensated controller by using root locus analysis [166]

The angles from closed dominant pole to others pole are
$\theta_{1}=180^{\circ}-\left[\tan ^{-1}\left[\frac{13.64}{20}\right]\right]=145.7^{\circ} \quad \theta_{2}=180^{\circ}-\left[\tan ^{-1}\left[\frac{13.64}{20-0.1551}\right]\right]=145.5^{\circ}$
$\theta_{3}=\left[\tan ^{-1}\left[\frac{13.64}{-333.1783-20}\right]\right]=2.494^{\circ}$
This angle is used to find the zero location for finding the PD controller zero $\left(z_{c}\right)$, The line connecting from complex pole to pole at the origin and pole at -0.1551 and 333.1783. Let us assume complex pole considered as $\mathbf{A}$ [167]. The angles from complex pole $\mathrm{A}=180^{\circ}$ - (summation of angle measure from complex pole to other poles) + (summation of angle measure from complex pole to other zeros) $=180-$ $\left(\theta_{1}+\theta_{2}+\theta_{3}\right)=180-(145.7+145.5+2.494) \approx \mathbf{- 1 1 3 . 6 9 4}{ }^{\circ}$. Now, find the PD compensated zero from angle measured from complex pole A (-113.694 ${ }^{\boldsymbol{}}$ )


Figure (6.22) NPSAT-1 Yaw Attitude Root locus response
The MATLAB Simulation for closed loop system includes the DC motor dynamics, satellite inertia, and PD compensator in feed forward loop. From figure (6.22) shows the NPSAT-1 yaw attitude RL Responses. The loop gain is found as 1640.


Figure (6.23) Finding the location of the attitude PD controller ( $z c$ )
Using the geometry shown in Figure 6.23, $\tan \left(180^{\circ}-113.694^{\circ}\right)=\frac{13.64}{20-z_{c}}$

To design the PD compensator zero ( $z_{c}$, from above geometry calculates the value 14.014. Now, the PD controller dynamics is $K(\mathbf{s}+\mathbf{1 4 . 0 1 4})$. The $K$ is the loop gain. The loop gain found from the root locus at $1 \%$ maximum overshoot is 1650 [166]. From figure 6.24 shows the NPSAT-1 yaw attitude MATLAB Simulation for closed loop system includes satellite yaw attitude dynamics and PD compensator.


Figure (6.24) NPSAT-1 Yaw Attitude Response with PD controller
In Figure 6.24, The NPSAT-1 yaw attitude response with PD compensator. This includes the dynamics of satellites and orbital perturbation at the low earth orbiting satellite. Thus, PD controllers has been introduced $K(s+14.014) K$ is the loop gain calculating from root locus analysis 1650 at $1 \%$ overshoot. The oscillation of satellite due to the perturbation forces measured different time domain specification and variations in maximum overshoot expressed in (\%). This indicates the damping of actual to the critical value [167]. As per the design specifications settling time is less than 0.2 Sec , after implemented the PD controller in the feed forward loop, the response's reaches settling time 0.129 Sec , Rise time 0.0106 Sec were achieved.

### 6.5 Attitude Control: SRM Satellite

The next spacecraft to be considered for attitude (pitch, yaw, and roll) control is SRM Satellite. The SRM satellite is a Nano satellite developed by students from the Sri Ramaswamy Memorial University in India [178]. These types of satellite are used for atmospheric measurement such as measures the Green House Gases. The launch was multi payload mission shared with Jugnu (IIT Kanpur satellite). The attitude control design of the Nano satellite at altitude of 867 km in low earth circular orbit. The orbital angular velocity is $0.0010239 \mathrm{rad} / \mathrm{s}$. The disturbances torque considered in the spacecraft is Magnetic Disturbance Torque ( $1.367 \times 10^{-5}$ ), Aerodynamic Torque (3.166 x $10^{-2}$ ), and Solar Radiation Torque ( $6.647 \times 10^{-6}$ ). [178]

SRMSAT Principal Moments of inertia [178]

$$
=\left[\begin{array}{ccc}
6.2911 & 0 & 0 \\
0 & 5.9162 & 0 \\
0 & 0 & 4.6085
\end{array}\right] \mathrm{kg}-\mathrm{m}^{2}
$$

### 6.5.1 SRM Satelite Roll Response:

The step response of the roll attitude transfer function without controller is shown in Figure 6.25, it is clear the SRM satellite roll transient responses, settling time 14.7 Sec, rise time 3.25 Sec, peak time 7.5 Sec, and damping of the roll attitude system overshoot is $14.3 \%$.


Figure (6.25) SRM Satellite Roll attitude dynamics response with step command

As per the design specifications (we need to reduce it to $\leq 0.2$ seconds). And, the steady-state is settled exactly at 1 (zero steady-state error). Hence, here also there is no need for integral control [166]. The dynamics of the controller mentioned in the Appendix E. The angles from closed dominant pole to others pole are $\theta_{1}=145.7^{\circ}, \theta_{2}$ $=144.962^{\circ}, \theta_{3}=2.494^{\circ}$. The angles from complex pole $\mathrm{A}=180^{\circ}-$ (summation of angle measure from complex pole to other poles) + (summation of angle measure from complex pole to other zeros $)=180-\left(\theta_{1}+\theta_{2}+\theta_{3}\right)=180-(145.7+144.9620+$ 2.494) $\approx-293.162$. In figure 6.26 shows SRM satellite RL Responses used to finding loop gain 465 at damping ration 0.826 (one percent overshoot) [179]


Figure (6.26) SRM Satellite roll attitude Root Locus Response

Using the geometry calculating the PD compensator, $\tan \left(180^{\circ}-113.162^{\circ}\right)=\frac{13.64}{20-z_{c}}$, [154] Now, find the PD compensated zero from angle measured from complex pole A $\left(-113.162^{0}\right)$. To design the PD compensator zero ( $z_{c}$, ) from above geometry calculates the value $\mathbf{- 1 4 . 1 6}$. Now, the PD controller dynamics is $\boldsymbol{K}(\mathbf{s}+\mathbf{1 4 . 1 6})$. The $K$ is the loop gain. The loop gain found from the root locus at $1 \%$ maximum overshoot is 465 .


Figure (6.27) SRM satellite Roll Attitude Response with PD controller
Thus, PD controllers has been introduced $K(s+14.16) K$ is the loop gain calculating from root locus analysis 465 at $1 \%$ overshoot. The oscillation of satellite due to the perturbation forces measured different time domain specification and variations in maximum overshoot expressed in (\%) [167]. This indicates the damping of actual to the critical value. As per the design specifications settling time is less than 0.2 Sec , after implemented the PD controller in the feed forward loop, the response's reaches settling time 0.127 Sec , Rise time 0.01067 Sec were achieved.

### 6.5.2. SRM Satellite Pitch Response:

The step response of the pitch attitude transfer function without controller is shown in Figure (6.28) it is clear the SRM satellite roll transient responses, settling time 22.2 Sec, rise time 3.79 Sec , peak time 9 Sec , and damping of the roll attitude system overshoot is $22.2 \%$ [178]. The oscillation gradually reduces with settling down at time 41.6600 seconds (we need to reduce it to $\leq 0.2$ seconds). And, the steady-state is settled exactly at 1 (zero steady-state error). Hence, here also there is
no need for integral control [166]. The dynamics of the controller mentioned in the Appendix E [167]. The angles from closed dominant pole to others pole are $\theta_{1}=$ $145.7^{\circ}, \theta_{2}=145.313^{\circ}, \theta_{3}=2.4922^{\circ}$. The angles from complex pole A $[156]=180^{\circ}-$ (summation of angle measure from complex pole to other poles) + (summation of angle measure from complex pole to other zeros $)=180-\left(\theta_{1}+\theta_{2}+\theta_{3}\right)=180-$ (145.7 $+\quad 145.313+2.4922) \approx-293.5121$


Figure (6.28) SRM Satellite Pitch response with step input


Figure (6.29) SRM satellite Pitch attitude Root locus response

Using the geometry calculating the PD compensator, $\tan \left(180^{0}-113.512^{0}\right)=\frac{13.64}{20-z_{c}}$, Now, find the PD compensated zero from angle measured from complex pole A ($\mathbf{1 1 3 . 5 1 2}{ }^{0}$ ) [156]. To design the PD compensator zero ( $z c$, ) from above geometry calculates the value $\mathbf{- 1 4 . 0 6}$. Now, the PD controller dynamics is $\boldsymbol{K}(\mathbf{s}+\mathbf{1 4 . 0 6})$. The K is the loop gain. The loop gain found from the root locus at $1 \%$ maximum overshoot is 876 .


Figure (6.30) SRM Satellite Pitch Attitude Response with PD controller
Thus, PD controllers has been introduced $\boldsymbol{K}(\mathbf{s}+\mathbf{1 4 . 1 0 6}) \mathrm{K}$ is the loop gain calculating from root locus analysis 876 (See figure 6.29) at $1 \%$ overshoot. The oscillation of satellite due to the perturbation forces measured different time domain specification and variations in maximum overshoot expressed in (\%). This indicates the damping of actual to the critical value [179], [180]. As per the design specifications settling time is less than 0.2 Sec , after implemented the PD controller in the feed forward loop. In figure (6.30) shows the response's reaches settling time 0.128 Sec , Rise time 0.0107 Sec, maximum overshoot 11.1 (\%) at 1.11peak amplitude were achieved.

### 6.5.3. SRM Satellite Yaw Response:

The step response of the yaw attitude transfer function without controller is shown in Figure (6.31) it is clear the SRM satellite roll transient responses, settling time 19.5 Sec, rise time 3.54 Sec , peak time 8 Sec , and damping of the roll attitude system overshoot is $19.5 \%$ [178]. As per the design specifications (we need to reduce it to $\leq 0.2$ seconds). And, the steady-state is settled exactly at 1 (zero steady-state error). Hence, here also there is no need for integral control.



Figure (6.32) SRM satellite Yaw attitude Root locus response


Figure (6.33) SRM Satellite Yaw Attitude Response with PD controller

Using the geometry calculating the PD compensator, $\tan \left(180^{0}-113.399^{\circ}\right)=\frac{13.64}{20-z_{c}}$, Now, find the PD compensated zero from angle measured from complex pole A ($113.399^{\circ}$ ). To design the PD compensator zero $(z c$, , from above geometry calculates the value $\mathbf{- 1 9 . 8 4}$ [158]. Thus, PD controllers has been introduced $\boldsymbol{K}(\mathbf{s}+\mathbf{1 9 . 8 4}) \mathrm{K}$ is the loop gain calculating from root locus analysis 690 (See figure 6.32) at $1 \%$ overshoot [166]. The oscillation of satellite due to the perturbation forces measured different time domain specification and variations in maximum overshoot expressed in (\%). This indicates the damping of actual to the critical value. As per the design specifications settling time is less than 0.2 Sec , after implemented the PD controller in the feed forward loop. In figure (6.33) shows the response's reaches settling time 0.107 Sec , Rise time 0.0101 Sec , maximum overshoot 14.7 (\%) at 1.15peak amplitude were achieved.

### 6.6. Attitude Control: Pratham Satellite

The specifications of Pratham Satellite are 817 km altitude, Polar Sun Synchronous Orbit (PSSO), Mass, 10 kg , Size $26 \times 26 \times 26 \mathrm{~cm}$, Orbital angular velocity $=0.0010346$ rad/s [134]
Principal moments of inertia [134] of PRATHAM

$$
=\left[\begin{array}{ccc}
0.116 & 0 & 0 \\
0 & 0.109 & 0 \\
0 & 0 & 0.114
\end{array}\right] \mathrm{kg}-\mathrm{m}^{2}
$$

### 6.6.1. Pratham Satellite Roll Response

The step response of the roll attitude transfer function without controller is shown in Figure (6.34) the settling time 7.72 seconds, rise time 4.3 Seconds. As per the design specifications, we need to reduce the settling time to $\leq 0.2$ seconds. The dynamics of the controller mentioned in the Appendix E. The angles formed between the dominant pole and all other poles can be obtained $\theta_{1}=145.7^{\circ}, \theta_{2}=109.595^{\circ}, \theta_{3}=$ 2.5451. Now, the angle contribution required for the PD controller zero $\left(z_{c}\right)$ to make the root locus to pass through the desired dominant pole can be obtained as [166]

Angle contribution $=180-\left(\theta_{1}+\theta_{2}+\theta_{3}\right)$

$$
=180-(145.7+109.595+2.5451) \approx-257.846
$$



Figure (6.34) Pratham Satellite Roll attitude response with step command


Figure (6.35) Pratham Satellite Roll attitude root locus response
Using the geometry calculating the PD compensator, $\tan \left(180^{\circ}-77.846^{\circ}\right)=$ $\frac{13.64}{20-z_{c}}$, from which the location of the PD compensator zero, $z_{c}$, is found to be 17.0625 [166], [167]. Now, the loop gain $K$ for the PD-compensated system is 16.1 from root locus analysis at $1 \%$ maximum overshoot.


Figure (6.36) Pratham Satellite Roll Attitude Response with PD controller
Thus, PD controllers has been introduced $K(s+17.0625) \mathrm{K}$ is the loop gain calculating from root locus analysis 16.1 (See figure 6.35) at $1 \%$ overshoot. The oscillation of satellite due to the perturbation forces measured different time domain specification and variations in maximum overshoot expressed in (\%). This indicates the damping of actual to the critical value [181]. As per the design specifications settling time is less than 0.2 Sec , after implemented the PD controller in the feed forward loop. In figure (6.36) shows the response's reaches settling time 0.0323 Sec , Rise time 0.0126 Sec , maximum overshoot 2.3 (\%) at 1.02 peak amplitude were achieved.

### 6.6.2. Pratham Satellite Pitch response

The step response of the pitch attitude transfer function without controller is shown in Figure (6.37). The settling time is 7.73 seconds and rise time 4.31 seconds
(we need to reduce it to $\leq 0.2$ seconds). The dynamics of the controller mentioned in the Appendix E. The angles formed between the dominant pole and all other poles can be obtained $\theta_{1}=145.7^{\circ}, \theta_{2}=105.845^{\circ}, \theta_{2}=2.547^{\circ}$. Now, the angle contribution required for the PD controller zero $\left(z_{c}\right)$ in order to make the root locus to pass through the desired dominant pole can be obtained as [166], [167]. Angle contribution $=180-$ $\left(\theta_{1}+\theta_{2}+\theta_{3}\right)=180-(145.7+105.845+2.547) \approx-254.098$


Figure (6.37) Pratham Satellite Pitch attitude response with step command (without controller)
Using the geometry calculating the PD compensator, $\tan \left(180^{0}-74.098^{\circ}\right)=\frac{13.64}{20-z_{c}}$, from which the location of the PD compensator zero, $z c$, is found to be $\mathbf{- 1 6 . 1 1 4}$. Now, the loop gain $K$ for the PD-compensated system is 15 from root locus analysis at $1 \%$ maximum overshoot [166].

Pratham(IITB)Satellite Pitch RL Response


Figure (6.38) Pratham Satellite Pitch attitude Root locus response


Figure (6.39) Pratham Satellite Pitch attitude response with PD controller

Thus, PD controllers has been introduced $\boldsymbol{K}(\mathbf{s}+\mathbf{1 6 . 1 1 4}) \mathrm{K}$ is the loop gain calculating from root locus analysis 15 (See figure 6.38) at $1 \%$ overshoot. The oscillation of satellite due to the perturbation forces measured different time domain specification and variations in maximum overshoot expressed in (\%). This indicates the damping of actual to the critical value [181]. As per the design specifications settling time is less than 0.2 Sec , after implemented the PD controller in the feed forward loop. In figure (6.39) shows the response's reaches settling time 0.0201 Sec , Rise time 0.013 Sec , maximum overshoot 0.977 (\%) at 1.01 peak amplitude were achieved.

### 6.6.3. Pratham Satellite Yaw response:

The step response of the yaw attitude transfer function without controller is shown in Figure (6.40). The settling time is 7.72 seconds and rise time 4.3 seconds. As per the design specification, we need to reduce rise time to $\leq 0.2$ seconds) [134]. The dynamics of the controller mentioned in the Appendix E. The angles formed between the dominant pole and all other poles can be obtained $\theta_{1}=145.7^{\circ}, \theta_{2}=108.551^{\circ}$, $\theta_{3}=2.5474^{0}$. Now, the angle contribution required for the PD controller zero $\left(z_{c}\right)$ to make the root locus to pass through the desired dominant pole can be obtained as angle contribution $=180-\left(\theta_{1}+\theta_{2}+\theta_{3}\right)=180-(145.7+105.551+2.5474) \approx-$ $256.8052^{0}$


Figure (6.40) Pratham Satellite Yaw attitude response with step command


Figure (6.41) Pratham satellite Yaw attitude Root locus response


Figure (6.42) Pratham Satellite Yaw Attitude Response with PD controller
Using the geometry calculating the PD compensator, $\tan \left(180^{0}-76.805^{\circ}\right)=\frac{13.64}{20-z_{c}}$, from which the location of the PD compensator zero, $z_{c}$, is found to be $\mathbf{- 1 6 . 8 0 1}$. Now, the loop gain $K$ for the PD-compensated system is 15.8 (shown in Fig 6.41) calculates from root locus analysis at $1 \%$ maximum overshoot [166], [167]. Thus, PD controllers has been introduced $K(s+16.801) K$ is the loop gain calculating from root locus analysis 16.1 (See figure 6.41) at $1 \%$ overshoot. The oscillation of satellite due to the perturbation forces measured different time domain specification and variations in
maximum overshoot expressed in (\%). This indicates the damping of actual to the critical value. As per the design specifications settling time is less than 0.2 Sec , after implemented the PD controller in the feed forward loop [181]. The output response in figure (6.42) shows the response's reaches settling time 0.0193 Sec, Rise time 0.0127 Sec were achieved. In table 6.3 shows the Nano satellite attitudes (Roll, Pitch, Yaw) transient response of control system for NPSAT-1, SRM satellite, Pratham satellite (IITB) without controller and with PD compensated controller were compared.
Table 6.3: Nano Satellites Attitude Responses (a) NPSAT-1 (b) SRM satellite (c) Pratham satellite

| Specifications | Roll Attitude Dynamics |  | Pitch Attitude Dynamics |  | Yaw Attitude Dynamics |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | Without <br> controller | With <br> controller | Without <br> controller | With <br> controller | Without <br> controller | With <br> controller |
| Rise time (Sec) | 6.5 | 0.0106 | 6.2 | 0.0106 | 4.7 | 0.0106 |
| Overshoot (\%) | 55.1 | 11.2 | 53.2 | 11.2 | 40 | 11.2 |
| Settling time (Sec) | 108 | 0.129 | 102 | 0.129 | 49.9 | 0.129 |
| Peak time (Sec) | 18 | 0.025 | 8 | 0.025 | 12 | 0.025 |

(a)

| Specifications | Roll Attitude Dynamics |  | Pitch Attitude Dynamics |  | Yaw Attitude Dynamics |  |
| :--- | :--- | :--- | :--- | :--- | :--- | :--- |
|  | Without <br> controller | With <br> controller | Without <br> controller | With <br> controller | Without <br> controller | With <br> controller |
| Rise time (Sec) | 3.25 | 0.0107 | 3.79 | 0.0107 | 3.54 | 0.010 |
| Overshoot (\%) | 14.3 | 11 | 27.1 | 11.1 | 22 | 14.7 |
| Settling time <br> (Sec) | 14.7 | 0.127 | 22.2 | 0.128 | 19.5 | 0.107 |
| Peak time (Sec) | 7 |  | 9 | 0.03 | 8 | 0.025 |

(b)

| Specifications | Roll Attitude Dynamics |  | Pitch Attitude Dynamics |  | Yaw Attitude Dynamics |  |
| :--- | :--- | :--- | :--- | :--- | :--- | :--- |
|  | Without <br> controller | With <br> controller | Without <br> controller | With <br> controller | Without <br> controller | With <br> controller |
| Rise time (Sec) | 4.3 | 0.0126 | 4.31 | 0.013 | 4.3 | 0.0127 |
| Settling time <br> $(S e c)$ | 7.72 | 0.0323 | 7.73 | 0.0201 | 7.71 | 0.0193 |

(c)



For, consider the NPSAT - 1 design parameters is $\left[I_{x x}=24.67 I_{y y}=22.63 I_{z z}=11\right] \mathrm{kg}$ $\mathrm{m}^{2}, \Omega=0.0010239 \mathrm{rad} / \mathrm{s}$ (at 550 km Altitude, ) Low Earth circular orbit (LEO), SRM satellite design parameters $\left[I_{x x}=6.2911 I_{y y}=5.9162 I_{z z}=4.6085\right] \mathrm{kg}-\mathrm{m}^{2}, \Omega=$ $0.0010239 \mathrm{rad} / \mathrm{s}$ (at 867 km Altitude), Pratham satellite design parameters [ $I_{x x}=I_{x x}=$ $0.116 I_{y y}=0.109 I_{z z}=0.114 \mathrm{~kg}-\mathrm{m}^{2}, \Omega=0.0010346 \mathrm{rad} / \mathrm{s}(817 \mathrm{~km})$. After compared responses of all satellite with controller gain values varying from 3700 K for NPSAT1 LEO satellite and lower value K for SRM satellite is 876 , Pratham satellite is 15 . Figures (6.43), (6.44), and (6.45) shows the outputs transient response of Nano satellites (NPSAT-1, SRM Satellite, and Pratham Satellite) analysis of without controllers and with PD compensated controller were compared. It is noticed that attitude responses of Nano satellite without controllers has large overshoot and settling time (NPSAT -1 is $\% \mathrm{Mp}=55.1 \%, \&$ ts $=108$ Seconds, SRM satellite is $\% \mathrm{Mp}$ $=14.3 \%$ \& ts $=14.7$ Seconds, Pratham satellite is $\mathrm{tr}=4.3$ Seconds \& ts $=7.72$ Seconds). The PD controllers have been introduced in the forward loop of satellite dynamically. The spacecraft control system simulation of the PD-compensated system satisfies the design requirement. It is used to increase the transient response (Overshoot and Settling time) of the system. For, settling time is (desired value $\leq 0.2$ seconds) wear achieved. The output responses of Nano satellite after implemented PD controllers of NPSAT-1 is $\% \mathrm{Mp}=11.2 \%, \& \mathrm{ts}=0.129$ Seconds, SRM satellite is $\% \mathrm{Mp}=11 \% \& \operatorname{ts}=0.127$ Seconds, Pratham satellite is $\mathrm{tr}=0.0126$ Seconds $\& \mathrm{ts}=$ 0.0323 Seconds. The Nano satellites comparative analysis with PD compensated system is achieved as per the design specification settling time $\leq 0.2$ seconds, All the Nano satellites output attitude transient response meet as per design requirements.

## CHAPTER 7

## CONCLUSION \& RECOMMENDATION OF FUTURE WORK

In present work, the Nano satellite perturbation forces and attitude control were modeled. To study variation in attitude of satellite due to orbital perturbation in low earth orbit, Aerodynamic drag, the Gravitational attraction of the Earth, and solar pressure were considered. The perturbation analysis of Low earth orbits Nano Satellite included International Space Station, Pratham (IIT Bombay) Satellite, SRM Satellite in this report. The perturbation simulation is implemented using MATLAB Tools, Python, and GMAT. The Cowell's perturbation simulation Keplerian results are validated with General mission analysis tool. The design parameters of Nano satellites Moment of inertia, NORAD two-line element, Geometry parameters are considered. It is a very accurate tool comparing with Euler angle method. Singularity problems are omitted in quaternion attitude estimation techniques.

The armature control DC motor is introduced in the Nano satellite attitude control system. The GEO magnetic fields modeling is considered in the International Geomagnetic Reference Field (IGRF-12). The attitude control for satellite using DC motor is performed. In obtained results the perturbing accelerations changes the satellite orbit in LEO. At LEO, International Space Station were investigated by the second order mathematical ODE equations by Runge - Kutta using Cowell's Method. To obtain the output of the numerical ODE integration, the initial position vectors and velocity vectors of the Nano satellite must be identified. The Kalman filter algorithm for determining the Nano-Satellite NPSAT-1 attitude (Roll, Pitch, Yaw) errors from on-board attitude sensor, like INS/GPS and Magnetometers are derived. The Kalman error estimation algorithm was developed using MATLAB package.

The detumble of spacecraft controlled by control torque generated from the actuator to plant dynamics. The attitude dynamics of pitch control, roll control, and yaw control of SRM Satellite, Pratham Satellite, and NPSAT-1 are considered. The
oscillatory response is obtained from attitude dynamics of the Nano satellites. The design of low-cost filter with suitable feedback is incorporated for all the satellite attitudes in this work.

The comparative study illustrates the improved Nano satellites transient responses with PD compensated controller using Armature controlled DC motor. The simulation of attitude control is implemented in MATLAB/SIMULINK environment. This thesis concludes the design of low-cost attitude estimation using INS/GPS and Magnetometers with low volume, less weight, less power, more accurate at lower altitudes.

This work mainly considered two body problems like Earth- Satellite at low earth orbit satellite, the future work extends to discuss the various three body problems like SUN perturbation, MOON perturbation at High earth orbit (HEO) such as geosynchronous orbit. The GEO orbit attitude estimation mostly depends upon the star sensor, sun sensor, earth sensor and horizon sensor. The actuator dynamics of HEO mostly rely on the Control momentum Gyroscope (CMG) and Flywheels or Momentum wheels. At HEO attitude information from on-board sensor secular variations are very high. The perturbation in the orbital element is greater than the orbital period. So, the attitude estimation uses Monto- Carlo Simulations instead of Kalman filter.

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## MODEL CODES

## COWELL'S PERTURBATION'S SIMULATION CODES

rho $=\operatorname{input}\left(\right.$ ( Density rho $\left[\mathrm{kg} / \mathrm{m}^{\wedge} 3\right]=$ ');
$\mathrm{Cd}=\operatorname{input}\left({ }^{( }\right.$Coefficient of Drag Cd = ');
A = input(' Area of the satellite A [m^2] = ');
$\mathrm{u}=\operatorname{input(}($ Velocity of the satellite $\mathrm{u}[\mathrm{m} / \mathrm{s}]=$ ');
$\mathrm{v}=\operatorname{input(}($ Velocity of the satellite $\mathrm{v}[\mathrm{m} / \mathrm{s}]=$ ');
$\mathrm{w}=\operatorname{input(}('$ Velocity of the satellite $\mathrm{w}[\mathrm{m} / \mathrm{s}]=$ ');
E _ad $=0.5 *$ rho $* \mathrm{Cd} * \mathrm{~A} * \mathrm{u}^{\wedge} 2$;
F_ad $=0.5 *$ rho $* \mathrm{Cd}^{*} \mathrm{~A} *{ }^{*}{ }^{\wedge} 2$;
$\mathrm{G} \_\mathrm{ad}=0.5 *$ rho $* \mathrm{Cd} * \mathrm{~A} * \mathrm{w}^{\wedge} 2$;
$\mathbf{M}=\operatorname{input}\left({ }^{\prime}\right.$ Mass of the satellite $\mathrm{M}[\mathrm{kg}]=$ ');
p_ad = E_ad / M;
q_ad = F_ad / M;
r_ad = G_ad / M;
$\mathrm{mu}=3.986 * 10^{\wedge} 14 ; \quad \% \mathrm{~m}^{\wedge} 3 / \mathrm{s}^{\wedge} 2$
$\mathrm{x}=\operatorname{input(}($ Distance between the satellite and the Earth $\mathrm{x}[\mathrm{m}]=\mathrm{I})$;
$y=\operatorname{input}($ ( Distance between the satellite and the Earth y $[\mathrm{m}]=$ ');
$\mathrm{z}=\operatorname{input}($ ' Distance between the satellite and the Earth $\mathrm{z}[\mathrm{m}]=$ ');
\% step size
$\mathrm{h}=7200$;
$\mathrm{t}=0: \mathrm{h}: 86400 ; \%$ in seconds
$\mathrm{A}=\mathrm{zeros}(1,10)$;
$\mathrm{B}=\operatorname{zeros}(1,10)$;
\% initial condition
$\mathrm{x}(1)=\mathrm{x}$;
$u(1)=u$;
$y(1)=y$;
$\mathrm{v}(1)=\mathrm{v}$;
$z(1)=z ;$
$\mathrm{w}(1)=\mathrm{w}$;
\% Functions
$\mathrm{Fx}=@(\mathrm{x}, \mathrm{t}) \mathrm{u}$;
$\mathrm{Fu}=@(\mathrm{u}, \mathrm{t}) \mathrm{p} \_\mathrm{ad}-\left((\mathrm{mu}) /\left(\mathrm{x}^{\wedge} 2\right)\right)$;
Fy = @ (y,t) v;
$\mathrm{Fv}=@(\mathrm{v}, \mathrm{t}) \mathrm{q} \_$ad $-\left((\mathrm{mu}) /\left(\mathrm{y}^{\wedge} 2\right)\right)$;
$\mathrm{Fz}=@(\mathrm{z}, \mathrm{t}) \mathrm{w}$;
$\mathrm{Fw}=@(\mathrm{w}, \mathrm{t}) \mathrm{r}_{-} \mathrm{ad}-\left((\mathrm{mu}) /\left(\mathrm{z}^{\wedge} 2\right)\right) ;$

```
for i=1:(length(t)) % calculation loop
    k_1 = Fx(u(i));
    1_1 = Fu(x(i));
    k_2 = Fx(u(i)+0.5*k_1);
    1_2 = Fu(x(i)+0.5*h);
    k_3 = Fx(u(i)+0.5*k_2);
    1_3 = Fu(x(i)+0.5*h);
    k_4 = Fx(u(i)+k_3);
    1_4 = Fu(x(i)+h);
    x(i+1) = x(i) + (1/6)*(k_1+2*k_2+2*k_3+k_4);
    u(i+1) = u(i) + (1/6)*(1_1+2*l_2+2*l_3+l_4);
    m_1 = Fy(v(i));
    n_1 = Fv(y(i));
    m_2 = Fy(v(i)+0.5*m_1);
    n_2 = Fv(y(i)+0.5*h);
    m_3 = Fy(v(i)+0.5*m_2);
    n_3 = Fv(y(i)+0.5*h);
    m_4 = Fy(v(i)+m_3);
    n_4 = Fv(y(i)+h);
    y(i+1)=y(i)+(1/6)*(m_1+2*m_2+2*m_3+m_4);
    v(i+1) = v(i) + (1/6)*(n_1+2*n_2+2*n_3+n_4);
    r_1 = Fz(w(i));
    s_1 = Fw(z(i));
    r_2 = Fz(w(i)+0.5*r_1);
    s_2 = Fw(z(i)+0.5*h);
    r_3 = Fz(w(i)+0.5*r_2);
    s_3 = Fw(z(i)+0.5*h);
    r_4 = Fz(w(i)+r_3);
    s_4 = Fw(z(i)+h);
    z(i+1) = z(i) + (1/6)*(r_1+2*r_2+2*r_3+r_4);
    w(i+1) = w(i) + (1/6)*(s_1+2*s_2+2*s_3+s_4);
    R=[x(i+1) y(i+1) z(i+1)]
    V = [u(i+1)v(i+1) w(i+1)]
    A(:,i) = sqrt(sum(R.^2))
    B(:,i) = sqrt(sum(V.^2))
end
A = A/1000;
B = B/1000;
```

A(1,:);
B(1,:);

```
figure(1)
plot(t,B, 'Linewidth', 1.5, 'color', 'blue')
xlabel('Time(sec)')
ylabel('Velocity(km/s)')
legend('Velocity vs Time')
figure(2)
plot(t,A, 'Linewidth', 1.5, 'color', 'red')
xlabel('Time(sec)')
ylabel('Position(km)')
legend('Position vs Time')
```


## KEPLERIAN CODES FOE INTERNATIONAL SPACE STATION (ISS) (ACTUAL/PREDICTED ORBITS)

```
mport matplotlib.pyplot as plt
x = [0,0.01,0.02,0.03,0.04]
y_blue =[51.875,51.866,51.866,51.855,51.845]
y_orange = [51.875,51.872,51.867,51.855,51.845]
plt.xlabel('Time (Days) ')
plt.ylabel(''Orbital Inclination(Deg)")
plt.plot(x,y_blue,label=''Predicted",marker='o')
plt.plot(x,y_orange,label=''Actual',marker='o')
b1=sum(y_orange)/5
b=sum(y_blue)/5
Pdiff = (b1-b)/b*100
print(b,b1,Pdiff)
plt.ylim(bottom=51.825,top=51.88)
plt.xlim(left=0,right=0.045)
plt.title("Orbital Inclination vs Time")
plt.text(0.0325,51.8270,"Percentage Difference=\n 0.002 %",fontsize=8)
plt.grid(color='black',linewidth=0.25,linestyle='--')
plt.legend()
plt.show()
```

import matplotlib.pyplot as plt
$\mathrm{x}=[0.025,0.03,0.035,0.04,0.045,0.05,0.055,0.06]$
y_blue $=[55,80,100,110,120,121,129,145]$
y_orange $=[62,83,101,120,120,124,130,143]$
plt.xlabel('Time (Days) ')
plt.ylabel("Argument Of Perigee")

```
plt.plot(x,y_blue,label=''Predicted',marker='o')
plt.plot(x,y_orange,label='Actual',marker='o')
b1=sum(y_orange)/8
b=sum(y_blue)/8
Pdiff = (b1-b)/b*100
print(b,b1,Pdiff)
plt.ylim(bottom=0,top=180)
plt.xlim(left=0,right=0.07)
plt.title("Argument of Perigeee vs Time")
plt.text(0.05,10,''Percentage Difference=\n 2.67 %',fontsize=8)
plt.grid(color='black',linewidth=0.25,linestyle='--'')
plt.legend()
plt.show()
import matplotlib.pyplot as plt
x = [0,0.01,0.02,0.03,0.04]
y_blue =[110,150,222,280,306]
y_orange = [96,155,222,280,304]
plt.xlabel('Time (Days) ')
plt.ylabel(''True Anomaly(Deg)")
plt.plot(x,y_blue,label=''Predicted'',marker='o')
plt.plot(x,y_orange,label='Actual'',marker='o')
b1=sum(y_orange)/5
b=sum(y_blue)/5
Pdiff = (b1-b)/b*100
print(b,b1,Pdiff)
plt.ylim(bottom=0,top=350)
plt.xlim(left=0,right=0.045)
plt.title('True Anomaly vs Time")
plt.text(0.0325,20,''Percentage Difference=\n 1.02 %',fontsize=8)
plt.grid(color='black',linewidth=0.25,linestyle='--')
plt.legend()
plt.show()
import matplotlib.pyplot as plt
x = [0,0.01,0.02,0.03,0.04]
y_blue =[6786,6785.2,6782.8,6780.3,6778.1]
y_orange = [6786.2,6785.9,6784,6780,6777.2]
plt.xlabel('Time (Days) ')
plt.ylabel(''Semi Major Axis(Km)")
plt.plot(x,y_blue,label=''Predicted',marker='o')
plt.plot(x,y_orange,label=''Actual'',marker='0')
b1=sum(y_orange)/5
b=sum(y_blue)/5
```

```
Pdiff = (b1-b)/b*100
print(b,b1,Pdiff)
plt.ylim(bottom=6776,top=6788)
plt.xlim(left=0,right=0.045)
plt.title(''Semi Major Axis vs Time')
plt.text(0.0325,6785,'Percentage Difference=\n 0.002 %'",fontsize=8)
plt.grid(color='black',linewidth=0.25,linestyle='--')
plt.legend()
plt.show()
import matplotlib.pyplot as plt
x}=[0,0.1,0.2,0.3,0.4,0.5,0.6,0.7,0.8,0.9,1
y_blue =[36,35,34,33,32.8,32.5,32.1,31,30,30,29.8]
y_orange = [36,35,34,33,32.8,32.5,32.5,32,31,30.5,30]
plt.xlabel('Time (Days) ')
plt.ylabel(''RAAN (Deg)')
plt.plot(x,y_blue,label=''Predicted",marker='o')
plt.plot(x,y_orange,label='Actual',marker='o')
b1=sum(y_orange)/11
b=sum(y_blue)/11
Pdiff = (b1-b)/b*100
print(b,b1,Pdiff)
plt.ylim(bottom=0,top=40)
plt.xlim(left=0)
plt.title('RAAN vs Time")
plt.text(0.76,2,"Percentage Difference=\n 0.87 %",fontsize=8)
plt.grid(color='black',linewidth=0.25,linestyle='--')
plt.legend()
plt.show()
import matplotlib.pyplot as plt
x = [0,0.01,0.02,0.03,0.04]
y_blue =[0.00039,0.00039,0.00039,0.000265,0.0002]
y_orange = [0.00036,0.00037,0.00038,0.00026,0.0002]
plt.xlabel('Time (Days) ')
plt.ylabel(''Eccentricity")
plt.plot(x,y_blue,label="'Predicted',marker='o')
plt.plot(x,y_orange,label=''Actual',marker='o')
b1=sum(y_orange)/5
b=sum(y_blue)/5
Pdiff = (b1-b)/b*100
print(b,b1,Pdiff)
plt.ylim(bottom=0,top=0.00045)
plt.xlim(left=0,right=0.05)
```

plt.title('Eccentricity vs Time")
plt.text( $0.0365,0.00035$,'"Percentage Difference=$=\mathbf{n} \quad 3.97 \%$ ",fontsize=8)
plt.grid(color='black',linewidth=0.25,linestyle='--')
plt.legend()
plt.show()

## KALMAN FILTER CODE FOR ATTITUDE ERRORS ESTIMATION

clc
clear all
close all
load magneto.txt;
$\mathrm{Ix}=24.67 ; \quad \% \mathrm{~kg}-\mathrm{m}^{\wedge} 2$
$\mathrm{Iy}=22.63 ; \quad \% \mathrm{~kg}-\mathrm{m}^{\wedge} 2$
$\mathrm{Iz}=11 ; \quad \% \mathrm{~kg}-\mathrm{m}^{\wedge} 2$
$\operatorname{mag}=$ magneto';
[ m n ] = size (mag);
$\mathrm{h}=10$; \% angular momentum of momentum wheel (Nms)
$\mathrm{w}=0.0011068 ; ~ \% ~ a n g u l a r ~ v e l o c i t y ~(r a d / s)$
$\mathrm{Td}=1.04 * 10^{\wedge}(-4) ; \quad \%$ Disturbance torque (Nm)
$\mathrm{dt}=0.1 ; \quad$ \% time duration (s)
$\mathrm{w} \_$dot $=\mathrm{w} / \mathrm{dt} ; \quad \%$ rate of change of angular velocity $(\mathrm{rad} / \mathrm{s}$ per second $)$
\% controller gain and matrix
pos_x $=2 ;$
vel_x $=2 ;$
pos_y $=2 ;$
vel_y $=2 ;$
$\operatorname{pos} z=2 ;$
vel_z $=2 ;$
$\mathrm{F}=[$ pos_x/Ix vel_x/Ix 0000 ;
00 pos_y/Iy vel_y/Iy 00 ;
0000 pos_x/Ix vel_x/Ix];

## \% Initial state matrix

$\mathrm{x} 0=[0 ; 0 ; 0 ; 0 ; 0 ; 0]$;
\% Initial error covariance matrix
$\mathrm{P} 0=\left[10^{\wedge}(-6) 00000 ; 0\right.$ 10^(-6) 0000 ; 00 10^(-6) 000 ; 000 10^(-6) 00 ; 0
000 10^(-6) 0; 00000 10^(-6)];

```
% process noise covariance matrix
Q = [1e-6 0000 0; 0 1e-6 000 0; 00 1e-6 000; 000 1e-6 0 0; 0000 1e-6
0; 00000 1e-6];
% measurement noise covariance matrix
R = [1e-2 0 0; 0 1e-2 0;0 0 1e-2];
% Disturbance torque input matrix
ud = [(Td + (w*h))/Ix ; (Td + (Iy*w_dot))/Iy ; (Td - (w*h))/Iz];
% B Matrix
B = [0 0 0;1 0 0;0 0 0;0 1 0;0 0 0;0 0 1];
% H Matrix
H=[10000 0; 001000;0 0 0 0 1 0];
% A matrix
A = zeros(6,6);
A(1,1) = 0;
A(1,2) = 1;
A(1,3) = 0;
A(1,4) = 0;
A(1,5) = 0;
A(1,6) = 0;
A(2,1) = ((-4*(w^2)*(Iy-Iz)) + (w*h)) / Ix;
A(2,2) = 0;
A(2,3) = 0;
A(2,4) = -h/Ix;
A(2,5) = w_dot;
A(2,6) = ((-w*(Iy-Ix-Iz))+h)/Ix;
A(3,1) = 0;
A(3,2) = 0;
A(3,3) = 0;
A(3,4) = 1;
A(3,5) = 0;
A(3,6) = 0;
A(4,1) = (-w*h)/Iy;
A(4,2) = -h/Iy;
A(4,3) = (-3* (w^2)*(Ix-Iz)) / Iy;
A(4,4) = 0;
A(4,5) = (-w*h)/Iy;
A(4,6) = -h/Iy;
A(5,1) = 0;
A(5,2) = 0;
```

```
A(5,3) = 0;
A(5,4) = 0;
A(5,5) = 0;
A(5,6) = 1;
A(6,1) = -w_dot;
A(6,2) = ((-w*(Ix-Iy+Iz))-h)/Iz;
A(6,3) = 0;
A(6,4) = h/Iz;
A(6,5) = ((-(w^2)*(Iy-Ix)) + (w*h)) / Iz;
A(6,6) = 0;
for i=1:n
    x_p = ((A - (B*F))*x0) + B * ud;
    P_p = (A * P0 * A') + Q;
    % Kalman gain
    Kk= P_p * H' * inv(H * P_p * H' + R);
    % Measurement
    zk = mag(:,n);
    % State update
    x_new = x_p + Kk*(zk - H * x_p);
    % state covariance update
    P_new = (eye(6) - Kk * H)* P_p;
    x0 = x_new;
    P0 = P_new;
    x_vector(:,i) = x_new;
    time(i,1) = i;
end
state = x_vector';
roll = magneto(:,1);
pitch = magneto(:,2);
yaw = magneto(:,3);
roll_error = state(:,1);
pitch_error = state(:,3);
yaw_error = state(:,5);
```

```
roll_est = roll + roll_error;
pitch_est = pitch + pitch_error;
yaw_est = yaw + yaw_error;
```

figure
plot(time,roll);
hold on
plot(time,roll_est,'r');
title('Estimated Roll angle - actual (black) vs estimated (red)')
xlabel('Time in seconds');
ylabel('Degrees');
figure
plot(time,pitch);
hold on
plot(time,pitch_est,'r');
title('Actual Pitch angle - actual (blue) vs estimated (red)')
xlabel('Time in seconds');
ylabel('Degrees');

## figure

\%plot(time,yaw);
\%hold on
plot(time,yaw_est,'r');
title('Yaw angle - actual (blue) vs estimated (red)')
xlabel('Time in seconds');
ylabel('Degrees');

## SIMULATION CODES FOR NANO SATELLITE (NPSAT-1) ATTITUDE CONTROL

```
%%% NPSAT-1 Roll attitude control
Dcmotor=tf(.85,[.003 1]);
Spacecraftwithinertia=tf(.04,[1 .04]);
integrator=tf(1, [1 0]);
a=Dcmotor*Spacecraftwithinertia;
cl=feedback([a],[0.85]);
openloop=cl*integrator;
withoutcontroller=feedback(openloop,1);
comp=tf([[1 13.98],1);
il=openloop*comp;
figure(1)
rlocus(il)
title ('NPSAT-1 RL Response')
```

```
cl2=feedback(openloop,1);
figure(2);
step(cl2)
title('NPSAT-1 Roll Response without controller')
gain=3700;
wc=gain*comp*openloop;
cl3=feedback(wc,1);
figure(3)
step(cl3)
title('NPSAT-1 Roll Response with PD-compensator')
```

\%\%\% Design NPSAT-1 Pitch Response with PD controller
Dcmotor=tf(.85, [.003 1]);
Spacecraftwithinertia=tf(.0442, [1 .0442]);
integrator=tf (1, [1 0]);
a=Dcmotor*Spacecraftwithinertia;
cl=feedback([a], [0.85]);
openloop=cl*integrator
withoutcontroller=feedback (openloop,1);
comp=tf([113.98],1);
il=openloop*comp;
figure (1)
rlocus(il)
title ('NPSAT-1 RL Response')
cl2=feedback (openloop,1);
figure (2) ;
step (cl2)
title('NPSAT-1 Pitch Response without controller')
gain=2500;
wc=gain*comp*openloop;
cl3=feedback (wc, 1) ;
figure (3)
step (cl3)
title('NPSAT-1 Pitch Response with PD-compensator')
\%\%\% NPSAT-1 yaw Attitude control system
Dcmotor=tf(.85, [.003 1]);
scSpacecraftwithinertia=tf(.09, [1.09]);
integrator=tf (1, [1 0]);
a=dc*Sc;

```
cl=feedback([a],[0.85]);
openloop=cl*int
withoutcontroller=feedback(openloop,1);
comp=tf([1 13.98],1);
il=openloop*comp;
figure(1)
rlocus(il)
title ('NPSAT-1 RL Response')
cl2=feedback(openloop,1);
figure(2);
step(cl2)
title('NPSAT-1 Yaw Response without controller')
gain=2500;
wc=gain*comp*openloop;
cl3=feedback(wc,1);
figure(3)
step(cl3)
title('NPSAT-1 Yaw Response with PD-compensator')
```


## APPENDIX-A (Satellite NORAD Data)



| High Quality Stock <br> Photos | PRATHAM |
| :---: | :---: |
| Trace A Cell Phone Location | Track PRATHAM now! 10-day_predictions |
| Watch Live Satellite TV | PRATHAM is classified as: <br> Amateur radio |
| Watch Live Streaming | Space \& Earth Science <br> CubeSats |
| Live Satellite Maps |  |
| Satellite Tracking System | NORAD ID: 41783 © Int'I Code 2016-059A 9 Perigee: 666.2 km © Apogee: 715.5 km © Inclination $98.1^{\circ}$ o |
| Live Satellite Images | Period: 98.4 minutes ${ }^{6}$ <br> Semi major axis: 7061 km - <br> RCS: Unknown $\boldsymbol{\varphi}$ |
| GPS Cell Phone Tracking | Launch date: September 26. 2016 <br> Source: India (IND) <br> Launch site: SRIHARIKOTA (SRI) |
| Real Time Satellite Tracking | Uplink (MHz): <br> Downlink (MHz): 437.455 <br> Beacon (MHz): 145.980 |
| Cell Phone Tracking Device | Mode: 1200 bps AFSK CW <br> Call sign: PRATHAM <br> Status: Inactive |

- $>$ PRATHAM is an Indian ionospheric research satellite which will be operated by the Indian Institute of Technology Bombay as part of the Student Satellite Initiative. Its primary mission is to count electrons in the Earth's ionosphere.



# SATVIEW <br> Tracking Satellites <br> 아를 


https://www.amsat.org/tle/current/nasa.all
https://www.n2yo.com/satellite/?s=41783\#results
https://www.celestrak.com/NORAD/elements/supplemental/
https://www.heavens-
above.com/satinfo.aspx?satid=41783\&lat=0\&lng=0\&loc=Unspecified\&alt=0\&tz=UCT\&cul= en
http://www.satview.org/
https://www.space-track.org/documentation\#/tle
https://amsat-uk.org/tag/srmsat/
http://stuffin.space/
Appendix: (TLE: Two Line Elements)
SRM Satellite Two Line Element Set (TLE): ©

1 37841U 11058D 18280.99540663 .00000306 00000-0 22990-4 09994
23784119.9713124 .72500011809289 .9476221 .840914 .10615917360831

Pratham Satellite Two Line Element Set (TLE):
1 41783U 16059A 18280.71788194 +.00000066 +00000-0 +22131-4 09993
241783098.1057340 .38870034923102 .0312258 .480814 .62946942108414

International Space Station Two Line Element Set (TLE):
1 25544U 98067A 08264.51782528-.00002182 00000-0 -11606-4 02927
22554451.6416247 .46270006703130 .5360325 .028815 .72125391563537

First Line:


| Field | Columns | Content | Example |
| :--- | :--- | :--- | :--- |
| 1 | $01-01$ | Line number | 1 |
| 2 | $03-07$ | Satellite number | 25544 |
| 3 | $08-08$ | Classification (U=Unclassified) | U |
| 4 | $10-11$ | International Designator (Last two digits of launch year) | 98 |
| 5 | $12-14$ | International Designator (Launch number of the year) | 067 |
| 6 | $15-17$ | International Designator (piece of the launch) | A |
| 7 | $19-20$ | Epoch Year (last two digits of year) | 08 |
| 8 | $21-32$ | Epoch (day of the year and fractional portion of the day) | 264.51782528 |
| 9 | $34-43$ | First Time Derivative of the Mean Motion divided by two [11] | -.00002182 |
| 10 | $45-52$ | Second Time Derivative of Mean Motion divided by six (decimal point assumed) | $00000-0$ |
| 11 | $54-61$ | BSTAR drag term (decimal point assumed) ${ }^{\text {[11] }}$ | $-11606-4$ |
| 12 | $63-63$ | The number 0 (originally this should have been "Ephemeris type") | 0 |
| 13 | $65-68$ | Element set number. Incremented when a new TLE is generated for this object.[11] | 292 |
| 14 | $69-69$ | Checksum (modulo 10) | 7 |

Second Line:


| Field | Columns | Content | Example |
| :--- | :--- | :--- | :--- |
| 1 | $01-01$ | Line number | 2 |
| 2 | $03-07$ | Satellite number | 25544 |
| 3 | $09-16$ | Inclination (degrees) | 51.6416 |
| 4 | $18-25$ | Right ascension of the ascending node (degrees) | 247.4627 |
| 5 | $27-33$ | Eccentricity (decimal point assumed) | 0006703 |
| 6 | $35-42$ | Argument of perigee (degrees) | 130.5360 |
| 7 | $44-51$ | Mean Anomaly (degrees) | 325.0288 |
| 8 | $53-63$ | Mean Motion (revolutions per day) | 15.72125391 |
| 9 | $64-68$ | Revolution number at epoch (revolutions) | 56353 |
| 10 | $69-69$ | Checksum (modulo 10) | 7 |

## APPENDIX -B

## ISS TRAJECTORY DATA

Lift off time (UTC) : N/A
Area (sq ft) : 21963.7592
Drag Coefficient (Cd) : 2.00
Monthly MSFC 50\% solar flux (F10.7-jansky) : 74.2
Monthly MSFC 50\% earth geomagnetic index (Kp) : 2.468
ET - UTC (sec) : 68.18
UT1 - UTC (sec) : 0.00

Maneuvers contained within the current ephemeris are as follows:


Coasting Arc \#1 (Orbit 3140)

Vector Time (GMT): 2018/122/12:00:00.000
Vector Time (MET): N/A
Weight (LBS) : 928423.9

M50 Cartesian
$\mathrm{X}=4422324.43$-meter
$\mathrm{Y}=\quad-214547.33-$ meter
$\mathrm{Z}=5128165.55 \quad \mathrm{I}=51.27620$
XDOT $=2036.764769 \quad \mathrm{Wp}=\quad 48.67787$
YDOT $=\quad 7253.759472$ meter $/ \mathrm{sec}$ RA $=245.45874 \mathrm{deg}$
ZDOT $=-1446.770461 \quad$ TA $=\quad 55.34489$
$\mathrm{MA}=55.27504$
$\mathrm{Ha}=\quad 221.586 \mathrm{n} . \mathrm{mi}$
Hp = 213.875
M50 Cartesian
$\mathrm{X}=14508938.43$
$\mathrm{Y}=$-703895.45-meter
$\mathrm{Z}=16824690.12$
$\mathrm{XDOT}=\quad 6682.299112$
YDOT $=23798.423466$ feet $/ \mathrm{sec} \quad$ YDOT $=\quad 7276.120938$ meter $/ \mathrm{sec}$
ZDOT $=-4746.622246 \quad$ ZDOT $=\quad-1437.056468$

TDR Cartesian
$\mathrm{X}=10665933.08 \quad \mathrm{X}=3250976.40$
$\mathrm{Y}=-9696226.22$ - meter $\quad \mathrm{Y}=-2955409.75$
$\mathrm{Z}=16920284.55 \quad \mathrm{Z}=5157302.73$
XDOT $=19605.714579 \quad$ XDOT $=\quad 5975.821804$
YDOT $=13322.224614-$ meter $/ \mathrm{sec} \quad$ YDOT $=\quad 4060.614062$
ZDOT $=-4704.230580 \quad$ ZDOT $=\quad-1433.849481$

The mean element set is posted at the UTC for which position is Just north of the next ascending node relative to the above Vector time

## APPENDIX -C

Spacecraft Control Toolbar


Spacecraft Control Toolbox
Add-ons


PlugInDemo at the end of the simulation


PlugInDemo on starting






## APPENDIX D

Satellite Attitude Data: https://www.raspberrypi.org/learning/
Kalman filter error covariance matrix: Process and Measurements

## Q matrix: Process covariance

| $1.00 E-06$ | $0.00 E+00$ | $0.00 E+00$ | $0.00 E+00$ | $0.00 E+00$ | $0.00 E+00$ |
| ---: | ---: | ---: | ---: | ---: | ---: |
| $0.00 E+00$ | $1.00 E-06$ | $0.00 E+00$ | $0.00 E+00$ | $0.00 E+00$ | $0.00 E+00$ |
| $0.00 E+00$ | $0.00 E+00$ | $1.00 E-06$ | $0.00 E+00$ | $0.00 E+00$ | $0.00 E+00$ |
| $0.00 E+00$ | $0.00 E+00$ | $0.00 E+00$ | $1.00 E-06$ | $0.00 E+00$ | $0.00 E+00$ |
| $0.00 E+00$ | $0.00 E+00$ | $0.00 E+00$ | $0.00 E+00$ | $1.00 E-06$ | $0.00 E+00$ |
| $0.00 E+00$ | $0.00 E+00$ | $0.00 E+00$ | $0.00 E+00$ | $0.00 E+00$ | $1.00 E-06$ |

R matrix: Sensor Noise covariance

| 0.01 | 0 | 0 |
| ---: | ---: | ---: |
| 0 | 0.01 | 0 |
| 0 | 0 | 0.01 |

F is control matrix (Controllers PD Gain)
$\begin{array}{lllllll}0.0810701256586948 & 0.0810701256586948 & 0 & 0 & 0 & 0\end{array}$
$\begin{array}{lllllll}0 & 0 & 0.0883782589482987 & 0.0883782589482987 & 0 & 0\end{array}$
$\begin{array}{llllll}0 & 0 & 0 & 0 & 0.0810701256586948 & 0.0810701256586948\end{array}$

H matrix:

| 1 | 0 | 0 | 0 | 0 | 0 |
| :--- | :--- | :--- | :--- | :--- | :--- |
| 0 | 0 | 1 | 0 | 0 | 0 |
| 0 | 0 | 0 | 0 | 1 | 0 |

## $Z_{k}$ (Measurements)

52.8900000000000
1.71000000000000
186.200000000000

## ATTITUDES DATA:

| ROW_ID | pitch | roll | yaw | mag_x | mag_y | mag_z | gyro_x | gyro_y | gyro_z |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | 1.49 | 52.25 | 185.21 | -46.4228 | -8.13291 | -12.1293 | 0.000942 | 0.000492 | -0.00075 |
| 2 | 1.03 | 53.73 | 186.72 | -48.779 | -8.30424 | -12.9431 | 0.000218 | -5.00E-06 | -0.00024 |
| 3 | 1.24 | 53.57 | 186.21 | -49.1619 | -8.47083 | -12.6428 | 0.000395 | 0.0006 | -3.00E-06 |
| 4 | 1.57 | 53.63 | 186.03 | -49.3419 | -8.45738 | -12.6155 | 0.000308 | 0.000577 | -0.0001 |
| 5 | 0.85 | 53.66 | 186.46 | -50.0567 | -8.12261 | -12.6783 | 0.000321 | 0.000691 | 0.000272 |
| 6 | 0.85 | 53.53 | 185.52 | -50.2465 | -8.34321 | -11.9381 | 0.000273 | 0.000494 | -5.90E-05 |
| 7 | 0.63 | 53.55 | 186.1 | -50.4473 | -7.93731 | -12.1886 | -0.00011 | 0.00032 | 0.000222 |
| 8 | 1.49 | 53.65 | 186.08 | -50.6682 | -7.7626 | -12.2842 | -4.40E-05 | 0.000436 | 0.000301 |
| 9 | 1.22 | 53.77 | 186.55 | -50.7615 | -7.26293 | -11.9811 | 0.000358 | 0.000651 | 0.000187 |
| 10 | 1.63 | 53.46 | 185.94 | -51.2438 | -6.87527 | -11.6725 | 0.000266 | 0.000676 | 0.000356 |
| 11 | 1.32 | 53.52 | 186.24 | -51.6165 | -6.81813 | -11.8608 | 0.000268 | 0.001194 | 0.000106 |
| 12 | 1.51 | 53.47 | 186.17 | -51.7817 | -6.74435 | -11.7484 | 0.000859 | 0.001221 | 0.000264 |
| 13 | 1.55 | 53.75 | 186.38 | -51.9927 | -6.52933 | -11.5983 | 0.000589 | 0.001151 | $2.00 \mathrm{E}-06$ |
| 14 | 1.07 | 53.63 | 186.6 | -52.4092 | -6.10015 | -11.7199 | 0.000497 | 0.00061 | -6.00E-05 |
| 15 | 0.81 | 53.4 | 186.32 | -52.6485 | -6.34696 | -11.596 | -5.30E-05 | 0.000593 | -0.00014 |
| 16 | 1.51 | 53.34 | 186.42 | -52.8507 | -6.04319 | -11.7441 | -0.00024 | 0.000495 | 0.000156 |
| 17 | 1.82 | 53.49 | 186.39 | -53.4491 | -6.09123 | -11.652 | 0.000571 | 0.00077 | 0.000331 |
| 18 | 0.46 | 53.69 | 186.72 | -53.68 | -5.80839 | -11.7036 | -0.00019 | 0.000159 | 0.000386 |
| 19 | 0.67 | 53.55 | 186.61 | -54.159 | -5.63871 | -11.424 | -0.0005 | $9.40 \mathrm{E}-05$ | $8.40 \mathrm{E}-05$ |
| 20 | 1.23 | 53.43 | 186.21 | -54.4006 | -5.29371 | -11.099 | -0.00034 | $1.30 \mathrm{E}-05$ | $4.10 \mathrm{E}-05$ |
| 21 | 1.44 | 53.58 | 186.4 | -54.6094 | -5.3559 | -11.3701 | -0.00027 | 0.000279 | -9.00E-06 |
| 22 | 1.25 | 53.34 | 186.5 | -54.7461 | -5.15465 | -11.5307 | 0.000139 | 0.000312 | $5.00 \mathrm{E}-05$ |
| 23 | 1.18 | 53.49 | 186.69 | -55.0914 | -4.90302 | -11.3636 | -0.00049 | 8.60E-05 | $6.50 \mathrm{E}-05$ |
| 24 | 1.34 | 53.32 | 186.84 | -55.5163 | -4.63189 | -11.6334 | 0.000312 | 0.000175 | 0.000308 |
| 25 | 1.36 | 53.56 | 187.02 | -55.561 | -4.55227 | -11.6051 | -0.0001 | $-2.30 \mathrm{E}-05$ | 0.000377 |
| 26 | 1.17 | 53.44 | 186.95 | -56.0164 | -4.51512 | -11.427 | 0.000147 | $5.40 \mathrm{E}-05$ | 0.000147 |
| 27 | 0.88 | 53.41 | 186.57 | -56.3937 | -4.35542 | -11.005 | -0.00013 | -0.00019 | 0.000269 |
| 28 | 0.78 | 53.84 | 186.85 | -56.5245 | -4.30156 | -11.2208 | -0.00018 | -0.00031 | 0.000361 |
| 29 | 0.88 | 53.41 | 186.62 | -56.7916 | -4.07602 | -11.064 | -0.00038 | -0.00025 | 0.000132 |
| 30 | 0.86 | 53.29 | 186.71 | -56.9155 | -3.8673 | -11.0856 | $3.10 \mathrm{E}-05$ | -0.00026 | $6.90 \mathrm{E}-05$ |
| 31 | 0.64 | 53.57 | 187.09 | -57.6625 | -3.47845 | -11.118 | 0.0002 | 0.000184 | 0.000275 |
| 32 | 1.02 | 53.41 | 186.79 | -57.996 | -3.39674 | -10.8548 | $8.20 \mathrm{E}-05$ | -0.00022 | -2.90E-05 |
| 33 | 0.87 | 53.54 | 187.12 | -58.3175 | -3.00518 | -10.9324 | 0.000123 | $2.90 \mathrm{E}-05$ | -3.90E-05 |
| 34 | 0.35 | 53.55 | 187.03 | -58.2887 | -3.03725 | -10.7718 | 0.000418 | -5.30E-05 | 0.00016 |
| 35 | 1.47 | 53.48 | 187.33 | -58.5083 | -2.51236 | -10.9999 | 0.000461 | 0.000641 | 0.000109 |
| 36 | 1.82 | 53.14 | 187.53 | -58.7559 | -2.48546 | -11.2323 | -0.00023 | -3.00E-06 | $2.50 \mathrm{E}-05$ |
| 37 | 1.37 | 53.3 | 187.49 | -58.9523 | -2.18282 | -11.1746 | 0.000383 | 0.000363 | 0.000154 |
| 38 | 1.68 | 53.26 | 187.54 | -58.8051 | -2.15404 | -11.1508 | -8.10E-05 | 8.80E-05 | -0.0003 |
| 39 | 0.83 | 53.58 | 187.79 | -59.2215 | -1.85493 | -11.0891 | 0.000238 | -0.00014 | $-7.50 \mathrm{E}-05$ |
| 40 | 1.37 | 53.34 | 187.62 | -59.5793 | -1.66924 | -10.7939 | -0.00032 | -0.00053 | -6.00E-06 |

## APPENDIX E (Satellite attitude dynamics)

Structure: Satellite response from with controller and without controller taken from MATLAB.


## SRM Satellite Pitch Dynamics:

Pitch dynamics $\frac{\theta(s)}{T_{y}(s)}=\frac{0.1690}{s^{2}+0.1690 s}$
Closed loop Response $=$

$$
\begin{aligned}
& 0.169 s+2.487 \\
& \text {---------------------- } \\
& s 2+0.338 s+2.487
\end{aligned}
$$

Continuous-time transfer function.

withoutcontroller $=$

RiseTime: 2.9526

SettlingTime: 41.6600
SettlingMin: 0.7336
SettlingMax: 1.5155
Overshoot: 51.5524
Undershoot: 0

Peak: 1.5155
PeakTime: 7.6299
withcontroller $=$

RiseTime: 0.0333
SettlingTime: 0.2111

SettlingMin: 0.9167
SettlingMax: 1.1716
Overshoot: 17.1642
Undershoot: 0
Peak: 1.1716
PeakTime: 0.0881

## SRM Satellite Roll Dynamics:

Pitch dynamics $\frac{\phi(s)}{T_{x}(s)}=\frac{0.1589}{s^{2}+0.1589 s}$
withoutcontroller $=$
RiseTime: 3.0218

SettlingTime: 49.2573
SettlingMin: 0.7219
SettlingMax: 1.5276
Overshoot: 52.7614

Undershoot: 0
Peak: 1.5276

PeakTime: 8.1148
withcontroller $=$
RiseTime: 0.0332
SettlingTime: 0.2111
SettlingMin: 0.9168
SettlingMax: 1.1717
Overshoot: 17.1712
Undershoot: 0
Peak: 1.1717
PeakTime: 0.0881

## SRM Satellite Yaw Dynamics:

Yaw dynamics $\frac{\psi(s)}{T_{Z}(s)}=\frac{0.2169}{s^{2}+0.2169 s}$
Closed loop response $=$

$$
\begin{gathered}
0.2169 s+3.195 \\
------------------1 \\
s^{2}+0.4338 s+3.195
\end{gathered}
$$

Continuous-time transfer function.
withoutcontroller $=$
RiseTime: 2.6813
SettlingTime: 35.8677
SettlingMin: 0.7783
SettlingMax: 1.4703
Overshoot: 47.0268
Undershoot: 0
Peak: 1.4703
PeakTime: 6.7942
withcontroller $=$
RiseTime: 0.0333
SettlingTime: 0.2112

SettlingMin: 0.9159
SettlingMax: 1.1711
Overshoot: 17.1113
Peak: 1.1711
PeakTime: 0.0882

## Pratham IITB Roll Dynamics:

Closed loop response $=$

$$
\begin{gathered}
8.62 s+160.9 \\
-------------------1 \\
s^{2}+17.24 s+160.9
\end{gathered}
$$

Continuous-time transfer function.
withoutcontroller $=$
RiseTime: 1.9424

SettlingTime: 3.5335
SettlingMin: 0.9008
SettlingMax: 0.9985
Overshoot: 0

Undershoot: 0
Peak: 0.9985
PeakTime: 5.7863
withcontroller $=$

RiseTime: 0.0588
SettlingTime: 0.2449
SettlingMin: 0.9099
SettlingMax: 1.0898
Overshoot: 8.9797

Undershoot: 0
Peak: 1.0898
PeakTime: 0.1354

## Pratham IITB Pitch Dynamics:

Closed loop Response=

$$
9.17 s+174.2
$$

$$
s^{2}+18.34 s+174.2
$$

Continuous-time transfer function.
withoutcontroller $=$
RiseTime: 1.9572
SettlingTime: 3.5591
SettlingMin: 0.9008
SettlingMax: 0.9998
Overshoot: 0
Undershoot: 0

Peak: 0.9998
PeakTime: 7.5608
withcontroller $=$
RiseTime: 0.0618
SettlingTime: 0.2494
SettlingMin: 0.9152
SettlingMax: 1.0841
Overshoot: 8.4097

Undershoot: 0
Peak: 1.0841
PeakTime: 0.1409

## Pratham IITB Yaw Dynamics:

Closed loop response $=$

$$
\begin{aligned}
& 8.77 \mathrm{~s}+164.5 \\
& s^{2}+17.54 s+164.5
\end{aligned}
$$

Continuous-time transfer function.
withoutcontroller $=$
RiseTime: 1.9466
SettlingTime: 3.5409
SettlingMin: 0.9010
SettlingMax: 0.9995
Overshoot: 0
Undershoot: 0
Peak: 0.9995

PeakTime: 6.7576
withcontroller $=$

RiseTime: 0.0598
SettlingTime: 0.2466

SettlingMin: 0.9060
SettlingMax: 1.0883
Overshoot: 8.8319
Undershoot: 0
Peak: 1.0883
PeakTime: 0.1365

## International Space Station (ISS) Roll Dynamics:

Closed loop response $=$

$$
\begin{gathered}
1.39 e-08 s+2.036 e-07 \\
-----------------------------2.036 e-07 \\
s^{2}+2.78 e-08 s+2.036
\end{gathered}
$$

Continuous-time transfer function.
withoutcontroller $=$

RiseTime: 1.5394e+04
SettlingTime: $5.6224 e+08$
SettlingMin: 0.0085

SettlingMax: 1.9955
Overshoot: 99.5531
Undershoot: 0
Peak: 1.9955

PeakTime: 1.3252e+05
withcontroller $=$
RiseTime: 0.0050
SettlingTime: 0.0474
SettlingMin: 0.9008
SettlingMax: 1.0303
Overshoot: 3.0310

Undershoot: 0
Peak: 1.0303

PeakTime: 0.0175

## International Space Station (ISS) Pitch Dynamics:

closedloop $=$

$$
\begin{gathered}
1.72 \mathrm{e}-08 \mathrm{~s}+2.52 \mathrm{e}-07 \\
--------------------------1 \\
\mathrm{~s}^{2}+3.44 \mathrm{e}-08 \mathrm{~s}+2.52 \mathrm{e}-07
\end{gathered}
$$

Continuous-time transfer function.
withoutcontroller $=$
RiseTime: 1.1085e+04
SettlingTime: $4.5453 \mathrm{e}+08$
SettlingMin: 0.0078
SettlingMax: 1.9961
Overshoot: 99.6091

Undershoot: 0

Peak: 1.9961
PeakTime: $4.5515 \mathrm{e}+05$
withcontroller $=$
RiseTime: 0.0197
SettlingTime: 0.1633
SettlingMin: 0.9070
SettlingMax: 1.1096
Overshoot: 10.9628
Undershoot: 0

Peak: 1.1096
PeakTime: 0.0550

## International Space Station (ISS) Yaw Dynamics:

Closed loop response $=$

$$
\begin{aligned}
& 8.04 e-09 s+1.178 e-07 \\
& \text {-------------------------------- } \\
& s^{2}+1.608 e-08 s+1.178 e-07
\end{aligned}
$$

Continuous-time transfer function.
withoutcontroller $=$
RiseTime: $6.4137 e+04$
SettlingTime: 9.7121e+08
SettlingMin: 0.0041
SettlingMax: 1.9880
Overshoot: 98.7951
Undershoot: 0
Peak: 1.9880
PeakTime: $2.9783 e+06$
withcontroller $=$
RiseTime: 0.0197
SettlingTime: 0.1631

SettlingMin: 0.9066
SettlingMax: 1.1095
Overshoot: 10.9529

Peak: 1.1095

PeakTime: 0.0549

## PUBLISHED RESEARCH PAPERS

1. Raja M, Dr Ugur Guven, Dr OM Prakash, "Design a Nano Satellite Attitude Control System Using Proportional-Derivative Controller", International Journal of Research and Analytical Reviews (IJRAR), PP 451-459, Volume 5, Issue 4 (2018).
2. Raja M, Dr Ugur Guven, Dr OM Prakash, Aman Saluja "Experimental Analysis of Low Earth Orbit Satellites due to Atmospheric Perturbations", IAETSD Journal for Advanced Research in Applied Sciences, Volume 5, Issue 4, April/2018
3. Raja M, Dr Ugur Guven, Dr OM Prakash, "Design and Analysis of Attitude Control Algorithm for Low Earth Orbiting Satellite with Magnetic Torquer Concepts using Non- Linear Unscented Kalman Filter", International Journal of Engineering and Technology (UAE), ISSN 2227-524X, Vol 7 \&. No 2.18 (2018).
4. Raja M, Dr Ugur Guven, Dr OM Prakash, "Experimental and analysis of Attitude Control System for Low Earth Orbit Satellite Attitude Errors Estimation with unscented Kalman Filter", International Journal of Emerging Technology and Advanced Engineering, PP 487-494, ISSN 2250-2459, ISO 9001:2008 Certified Journal, Volume 7, Issue 9, September 2017.
5. Raja M, Dr Ugur Guven, Dr OM Prakash, Saurabh Pandey "Design of trajectory and perturbation analysis for satellite orbital parameters", IAETSD Journal for Advanced Research in Applied Sciences, Volume 4, Issue 4, Sept /2017 ISSN (Online): 2394-8442
